



The boundary rigidity problem in the presence of a magnetic field

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Abstract

For a compact Riemannian manifold with boundary, endowed with a magnetic potential α , we consider the problem of restoring the metric g and the magnetic potential α from the values of the Mañé action potential between boundary points and the associated linearized problem. We study simple magnetic systems. In this case, knowledge of the Mañé action potential is equivalent to knowledge of the scattering relation on the boundary which maps a starting point and a direction of a magnetic geodesic into its end point and direction. This problem can only be solved up to an isometry and a gauge transformation of α .

For the linearized problem, we show injectivity, up to the natural obstruction, under explicit bounds on the curvature and on α . We also show injectivity and stability for g and α in a generic class \mathcal{G} including real analytic ones.

For the nonlinear problem, we show rigidity for real analytic simple (g, α) , rigidity for metrics in a given conformal class, and locally, near any $(g, \alpha) \in \mathcal{G}$. We also show that simple magnetic systems on two-dimensional manifolds are always rigid.

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1. Introduction

1.1. Statement of the problem

Let M be a compact manifold with boundary, endowed with a Riemannian metric g , and let $\pi : TM \rightarrow M$ denote the canonical projection, $\pi : (x, \xi) \mapsto x, x \in M, \xi \in T_x M$.

Denote by ω_0 the canonical symplectic form on TM , which is obtained by pulling back the canonical symplectic form of T^*M via the Riemannian metric. Let $H : TM \rightarrow \mathbb{R}$ be defined by

$$H(v) = \frac{1}{2}|v|_g^2, \quad v \in TM.$$

The Hamiltonian flow of H w.r.t. ω_0 gives rise to the geodesic flow of M . Let Ω be a closed 2-form on M and consider the new symplectic form ω defined as

$$\omega = \omega_0 + \pi^* \Omega.$$

The Hamiltonian flow of H w.r.t. ω gives rise to the *magnetic* (or *twisted geodesic*) flow $\psi^t : TM \rightarrow TM$. This flow models the motion of a unit charge of unit mass in a magnetic field whose Lorentz force $Y : TM \rightarrow TM$ is the bundle map uniquely determined by

$$\Omega_x(\xi, \eta) = \langle Y_x(\xi), \eta \rangle \tag{1.1}$$

for all $x \in M$ and $\xi, \eta \in T_x M$.

Magnetic flows were first considered by V. I. Arnold in [4] and by D.V. Anosov and Y.G. Sinai in [3]. As shown in [5,18,25–28], they are closely related to other problems of classical mechanics, mathematical physics, symplectic geometry, and dynamical systems.

It is not hard to show that the generator \mathbf{G}_μ of the magnetic flow is

$$\mathbf{G}_\mu(x, \xi) = \mathbf{G}(x, \xi) + Y_i^j(x)\xi^i \frac{\partial}{\partial \xi^j},$$

where \mathbf{G} is the generator of the geodesic flow, and that every trajectory of the magnetic flow is a curve of the form $t \mapsto (\gamma(t), \dot{\gamma}(t))$, where γ is a curve on M which satisfies the equation

$$\nabla_{\dot{\gamma}} \dot{\gamma} = Y(\dot{\gamma}), \tag{1.2}$$

which is nothing but Newton’s law of motion. Such a curve γ is called a *magnetic geodesic*. Note that time is not reversible on the magnetic geodesics, unless $\Omega = 0$. If $\Omega = 0$ we recover the ordinary geodesic flow and ordinary geodesics.

When Ω is exact, i.e. $\Omega = d\alpha$ for some magnetic potential α , the magnetic flow also arises as the Hamiltonian flow of $H(x, p) = \frac{1}{2}(p + \alpha)_g^2$ with respect to the standard symplectic form of T^*M . The function H is the symbol of the semiclassical magnetic Schrödinger operator.

Since the magnetic flow preserves the level sets of the Hamiltonian function H , every magnetic geodesic has constant speed.

Unlike the geodesic flow, where the flow is the same (up to time scale) on all energy levels, the magnetic flow depends essentially on the choice of the energy level. We fix the energy level

$H^{-1}(1/2)$, thus considering the magnetic flow on the unit sphere bundle SM of M , in consequence we consider only the unit speed magnetic geodesics. Note that fixing the energy level to be SM is no restriction at all, since one can obtain the behavior in any energy level by considering the flow on SM upon changing Ω by $\lambda\Omega$, where $\lambda \in \mathbb{R}$.

We define the action $\mathbb{A}(x, y)$ between boundary points as a minimizer of the appropriate action functional, see (1.6) and Appendix A. In the case $\Omega = 0$, $\mathbb{A}(x, y)$ coincides with the boundary distance function $d_g(x, y)$. In this case, we cannot recover g from d_g up to isometry, unless some additional assumptions are imposed on g , see, e.g., [9]. One such assumption is the simplicity of the metric, see, e.g., [22,34,37,38]. We consider below the analog of simplicity for magnetic systems.

Let Λ stand for the second fundamental form of ∂M and $\nu(x)$ for the inward unit normal to ∂M at x . We say that ∂M is *strictly magnetic convex* if

$$\Lambda(x, \xi) > \langle Y_x(\xi), \nu(x) \rangle \tag{1.3}$$

for all $(x, \xi) \in S(\partial M)$ (see Appendix A for an explanation). Note that if we replace ξ by $-\xi$, we can put an absolute value in the right-hand side of (1.3). In particular, magnetic convexity is stronger than Riemannian one.

For $x \in M$, we define the *magnetic exponential map* at x to be the partial map $\exp_x^\mu : T_x M \rightarrow M$ given by

$$\exp_x^\mu(t\xi) = \pi \circ \psi^t(\xi), \quad t \geq 0, \xi \in S_x M.$$

It is not hard to show that, for every $x \in M$, \exp_x^μ is a C^1 -smooth partial map on $T_x M$ which is C^∞ -smooth on $T_x M \setminus \{0\}$ (see Appendix A).

Definition 1.1. We say that M is *simple* (w.r.t. (g, Ω)) if ∂M is strictly magnetic convex and the magnetic exponential map $\exp_x^\mu : (\exp_x^\mu)^{-1}(M) \rightarrow M$ is a diffeomorphism for every $x \in M$ (cf. the definition of a simple Riemannian manifold [32]).

In this case, M is diffeomorphic to the unit ball of \mathbb{R}^n (so we can assume that M is this ball); therefore, Ω is exact, and we let α be a *magnetic potential*, i.e., α is a 1-form on M such that

$$d\alpha = \Omega. \tag{1.4}$$

Henceforth we call (g, α) a *simple magnetic system* on M . We will also say that (M, g, α) is a simple magnetic system.

Given $x, y \in M$, let

$$\mathcal{C}(x, y) = \{ \gamma : [0, T] \rightarrow M : T > 0, \gamma(0) = x, \gamma(T) = y, \gamma \text{ is absolutely continuous} \}.$$

The *time-free action* of a curve $\gamma \in \mathcal{C}(x, y)$ w.r.t. (g, α) is defined as

$$\mathbb{A}(\gamma) = \mathbb{A}_{g,\alpha}(\gamma) = \frac{1}{2} \int_0^T |\dot{\gamma}(t)|_g^2 dt + \frac{1}{2} T - \int_\gamma \alpha. \tag{1.5}$$

For a simple magnetic system, unit speed magnetic geodesics minimize the time-free action (Lemma A.5 in Appendix A). More precisely,

$$\mathbb{A}(x, y) := \inf_{\gamma \in \mathcal{C}(x, y)} \mathbb{A}(\gamma) = \mathbb{A}(\gamma_{x, y}) = T_{x, y} - \int_{\gamma_{x, y}} \alpha, \tag{1.6}$$

where $\gamma_{x, y} : [0, T_{x, y}] \rightarrow M$ is the unique unit speed magnetic geodesic from x to y .

The function $\mathbb{A}(x, y)$ is referred to as Mañé’s action potential (of energy $1/2$), and we call the restriction $\mathbb{A}|_{\partial M \times \partial M}$ the *boundary action function*.

Now, we state the *boundary rigidity problem in the presence of a magnetic field* as follows: To which extent is a magnetic system (g, α) determined by the boundary action function?

To be more precise, we say that two simple magnetic systems (g, α) and (g', α') are *gauge equivalent* if there are a diffeomorphism $f : M \rightarrow M$, which is the identity on the boundary, and a function $\varphi : M \rightarrow \mathbb{R}$, vanishing on the boundary, such that $g' = f^*g$ and $\alpha' = f^*\alpha + d\varphi$. Observe that gauge equivalent magnetic systems have the same boundary action function.

Now, we rephrase the above problem as follows: Given a simple magnetic system, is any other simple magnetic system on the same manifold gauge equivalent to the former as soon as it has the same boundary action function? If so, we call the given magnetic system *magnetic boundary rigid*.

Surely, this problem can be considered under various natural restrictions. For example, we can consider it for a fixed Riemannian metric and try to restore a magnetic potential, or, vice versa, fix a magnetic potential and try to restore a metric, or consider the problem for metrics in a fixed conformal class, etc. In particular, for the zero magnetic potentials we recover the ordinary boundary rigidity problem for Riemannian metrics (see recent surveys of the latter in [10,39]).

For simple magnetic systems, knowledge of the action $\mathbb{A}(x, y)$ between boundary points is equivalent to knowing the scattering relation, see Section 2.2. For non-simple systems, the problem of recovering (g, Ω) from the scattering relation is the natural problem to study. The scattering relation appears naturally in the study of the scattering operator in \mathbb{R}^n with g Euclidean outside a ball, and α compactly supported. Namely, for non-trapping metrics, the scattering operator associated to the semi-classical magnetic Schrödinger operator is a Fourier integral operator with canonical relation that determines the scattering relation on a large sphere, [1,2,15]. (It should be noted that for magnetic Schrödinger operators the resolvent is also a Fourier integral operator.)

1.2. Description of the results

In Section 2, we show that for simple magnetic systems, the action determines the jets of g and α in boundary normal coordinates. We define the scattering relation and show that for simple magnetic systems, it determines $\mathbb{A}(x, y)$ on the boundary, and vice versa. We also show that one can recover the volume from $\mathbb{A}(x, y)$.

In Section 3, we study the linearized problem. This reduces to the magnetic ray transform I . We show that potential pairs (see Definition 3.5) belong to the kernel of I . We say that I is s -injective, if the kernel of I consists only of potential pairs.

In Section 4, we show that the normal operator $N = I^*I$ is a pseudo-differential operator in the interior of M , elliptic on solenoidal pairs that are an orthogonal complement of the potential pairs. We construct a parametrix of N ; near the boundary, additional arguments are needed. This

parametrix recovers the solenoidal projection \mathbf{f}^* given $N\mathbf{f}$, up to a smoothing term. We show that s -injectivity implies a stability estimate, uniform near any (g, α) , in appropriate spaces, see Theorem 4.3. We consider in this section and in Appendix B real analytic (g, α) and show that then I is s -injective. A crucial element of the proof is that N is an analytic pseudo-differential operator in the interior of M . This is delicate since the magnetic exponential map is only C^1 smooth when $\Omega \neq 0$, even in the analytic case. To handle this, the analysis is done in polar coordinates. The s -injectivity for real analytic magnetic systems and the uniform stability estimate imply s -injectivity of the magnetic ray transform for generic (g, α) .

In Section 5, we show that I is s -injective for simple magnetic systems with an explicit bound on the curvature and Ω . This relies on an analog of Pestov's identity for our case, see [30,34] that goes back to [23,24], see also the references there.

In Section 6, we consider the non-linear magnetic rigidity problem. We prove rigidity in a given conformal class and rigidity within real analytic systems. We also show that if a simple Riemannian manifold (M, g) is boundary rigid (within the class of Riemannian metrics), then it is also magnetic boundary rigid. In this section we also study the local problem. We show rigidity near any (g, α) in the generic class \mathcal{G} using the analysis of the linear problem. This does not directly follow from the implicit function theorem, and the stability estimate in Theorem 4.3 plays a crucial role. There is an essential difficulty compared to the Riemannian case $\alpha = 0$ [37,38], since we cannot decouple g and α in the linearization argument. This difficulty is resolved by Lemma 6.7.

Section 7 is devoted to two-dimensional systems. Here we prove that two-dimensional simple magnetic systems are magnetic boundary rigidity. This generalizes the boundary rigidity theorem of [32]. Our proof essentially resembles that in [32], establishing a connection between the scattering relation of a magnetic system and the Dirichlet-to-Neumann map of the Laplace–Beltrami operator of the underlying Riemannian manifold.

2. Boundary determination, scattering relation, and volume

2.1. Boundary jets of the metric and magnetic potential

Here we show that up to gauge equivalence the boundary action function completely determines the Riemannian metric and magnetic potential on the boundary of the manifold under study.

Lemma 2.1. *If (g, α) and (g', α') are simple magnetic systems on M with the same boundary action function, then*

$$i^*g = i^*g', \quad i^*\alpha = i^*\alpha', \quad (2.1)$$

where $i : \partial M \rightarrow M$ is the embedding map.

Proof. Given $x \in \partial M$ and $\xi \in T_x(\partial M)$, let $\tau(s)$, $-\varepsilon < s < \varepsilon$, be a curve on ∂M with $\tau(0) = x$ and $\dot{\tau}(0) = \xi$. It is easy to see that

$$\lim_{s \rightarrow 0} \frac{\mathbb{A}(x, \tau(s))}{s} = |\xi|_g - \alpha(\xi).$$

A similar equality holds for the magnetic system (g', α') . Therefore,

$$|\xi|_g - \alpha(\xi) = |\xi|_{g'} - \alpha'(\xi).$$

Changing ξ to $-\xi$, we get

$$|\xi|_g + \alpha(\xi) = |\xi|_{g'} + \alpha'(\xi),$$

whence we infer (2.1). \square

Now, we prove that the boundary action function determines the full jets of the metric and magnetic potential on the boundary. This generalizes the corresponding results of [20,22].

Theorem 2.2. *If (g, α) and (g', α') are simple magnetic systems on M with the same boundary action function, then (g', α') is gauge equivalent to some $(\bar{g}, \bar{\alpha})$ such that in any local coordinate system we have $\partial^m g|_{\partial M} = \partial^m \bar{g}|_{\partial M}$ and $\partial^m \alpha|_{\partial M} = \partial^m \bar{\alpha}|_{\partial M}$ for every multi-index m .*

Proof. Denote by ν the inward unit normal to ∂M w.r.t. g . The “usual” boundary exponential map $\exp_{\partial M}(p, t) = \exp_p(t\nu(p))$, $p \in \partial M$, $t \geq 0$, takes a sufficiently small neighborhood of the set $\partial M \times \{0\}$ in $\partial M \times \mathbb{R}_+$ diffeomorphically onto some neighborhood of ∂M in M .

Let ν' and \exp' denote the corresponding objects for the metric g' .

The map $\exp_{\partial M} \circ (\exp'_{\partial M})^{-1}$ is well defined in some neighborhood of ∂M in M and is the identity when restricted to ∂M . We extend this map from a neighborhood of ∂M in M to a diffeomorphism $f: M \rightarrow M$ and put $\bar{g} = f^*g'$ and $\alpha'' = f^*\alpha'$. By Lemma 2.1, g and \bar{g} induce the same Riemannian metric on ∂M and, by construction, at each point of ∂M the inward unit normal w.r.t. \bar{g} coincides with the one w.r.t. g . Therefore, $\bar{g}|_{\partial M} = g|_{\partial M}$.

Next, by Lemma 2.1, α and α'' induce the same 1-form on ∂M . Applying Lemma 2.2 of [35] to the form $\omega = \alpha - \alpha''$ we find a function $\varphi \in C_0^\infty(M)$ such that $\omega - d\varphi$ induces the zero form on every sufficiently short geodesic of \bar{g} starting from ∂M in the normal direction. We now put $\bar{\alpha} = \alpha'' + d\varphi$.

Let us prove the equality of derivatives on the boundary. We will use the same argument as in [20]. We fix $x_0 \in \partial M$ and introduce boundary normal coordinates (x', x^n) w.r.t. g near x_0 . By construction, the same coordinates are boundary normal coordinates w.r.t. \bar{g} . Thus, the line elements ds^2 and $d\bar{s}^2$ of these metrics are given by

$$\begin{aligned} ds^2 &= g_{\iota\kappa} dx'_\iota dx'_\kappa + dx_n^2, \\ d\bar{s}^2 &= \bar{g}_{\iota\kappa} dx'_\iota dx'_\kappa + dx_n^2, \end{aligned}$$

where ι, κ vary from 1 to $n - 1$. Therefore, for $h = g - \bar{g}$ we have $h_{in} = 0$ for $i = 1, \dots, n$. Also, by the construction of $\bar{\alpha}$, for $\beta = \alpha - \bar{\alpha}$ we have $\beta_n = 0$.

It now suffices to prove that

$$\partial_n^m h_{\iota\kappa}|_{x=x_0} = 0, \quad \partial_n^m \beta_\iota|_{x=x_0} = 0 \quad \text{for } m = 0, 1, \dots; \iota, \kappa = 1, \dots, n - 1. \quad (2.2)$$

The case of $m = 0$ is granted. Assume that there is a least $m \geq 1$ such that (2.2) is not true. The Taylor expansion of h and β then implies that there is a unit vector $\xi_0 \in T_{x_0}(\partial M)$ such that

$$\frac{1}{2}h_{i\kappa}(x)\xi^i\xi^\kappa - \beta_i(x)\xi^i > 0 \quad (\text{or } < 0) \tag{2.3}$$

for $x^n > 0$ and x' both sufficiently small and ξ close to ξ_0 . (Here i and κ vary from 1 to $n - 1$ because $h_{1n} = h_{n1} = 0$ and $\beta_n = 0$.)

For arbitrary $x, y \in \partial M$, let $\gamma = \gamma_{x,y} : [0, T] \rightarrow M$ ($\bar{\gamma} = \bar{\gamma}_{x,y} : [0, \bar{T}] \rightarrow M$) be the unit speed magnetic geodesic of the system (g, α) (system $(\bar{g}, \bar{\alpha})$) from x to y .

On the one hand, since $\bar{\gamma}$ minimizes the time-free action w.r.t. $(\bar{g}, \bar{\alpha})$, we have

$$\mathbb{A}(x, y) \leq \frac{1}{2} \int_0^T \bar{g}_{ij}(\gamma(t)) \dot{\gamma}^i(t) \dot{\gamma}^j(t) dt + \frac{1}{2}T - \int_0^T \bar{\alpha}_i(\gamma(t)) \dot{\gamma}^i(t) dt.$$

On the other hand, since γ minimizes the time-free action w.r.t. (g, α) ,

$$\mathbb{A}(x, y) = \frac{1}{2} \int_0^T \left[g_{ij}(\gamma(t)) \dot{\gamma}^i(t) \dot{\gamma}^j(t) dt + \frac{1}{2}T - \int_0^T \alpha_i(\gamma(t)) \dot{\gamma}^i(t) \right] dt.$$

Therefore,

$$\int_0^T \left[\frac{1}{2}h_{ij}(\gamma(t)) \dot{\gamma}^i(t) \dot{\gamma}^j(t) - \beta_i(\gamma(t)) \dot{\gamma}^i(t) \right] dt \leq 0. \tag{2.4}$$

Similarly, we derive the inequality

$$\int_0^{\bar{T}} \left[\frac{1}{2}h_{ij}(\bar{\gamma}(t)) \dot{\bar{\gamma}}^i(t) \dot{\bar{\gamma}}^j(t) - \beta_i(\bar{\gamma}(t)) \dot{\bar{\gamma}}^i(t) \right] dt \geq 0. \tag{2.5}$$

Continuing the proof of the theorem, we now choose $x = x_0$ and choose $y = \delta(s)$, where $\delta : (-\varepsilon, \varepsilon) \rightarrow \partial M$ is a smooth curve with $\delta(0) = x_0$ and $\dot{\delta}(0) = \xi_0$. Then for $s > 0$ sufficiently small we see that, in view of (2.3), we cannot simultaneously have (2.4) and (2.5). This contradiction concludes the proof of the theorem. \square

2.2. Scattering relation

Now, we define a scattering relation and restate our problem in terms of the scattering relation.

For $(x, \xi) \in SM$, let $\gamma_{x,\xi} : [\ell^-(x, \xi), \ell(x, \xi)] \rightarrow M$ be a magnetic geodesic such that $\gamma_{x,\xi}(0) = x$, $\dot{\gamma}_{x,\xi}(0) = \xi$, and $\gamma_{x,\xi}(\ell^-(x, \xi)), \gamma_{x,\xi}(\ell(x, \xi)) \in \partial M$. Clearly, the functions $\ell^-(x, \xi)$ and $\ell(x, \xi)$ are continuous and, on using the implicit function theorem, they are easily seen to be smooth near a point (x, ξ) such that the magnetic geodesic $\gamma_{x,\xi}(t)$ meets ∂M transversally

at $t = \ell^-(x, \xi)$ and $t = \ell(x, \xi)$ respectively. By (1.3) and Lemma A.6 in Appendix A, the last condition holds everywhere on $SM \setminus S(\partial M)$. Thus, ℓ^- and ℓ are smooth on $SM \setminus S(\partial M)$.

Let $\partial_+ SM$ and $\partial_- SM$ denote the bundles of inward and outward unit vectors over ∂M :

$$\begin{aligned} \partial_+ SM &= \{(x, \xi) \in SM : x \in \partial M, \langle \xi, \nu(x) \rangle \geq 0\}, \\ \partial_- SM &= \{(x, \xi) \in SM : x \in \partial M, \langle \xi, \nu(x) \rangle \leq 0\}, \end{aligned}$$

where ν is the inward unit normal to ∂M . Note that $\partial(SM) = \partial_+ SM \cup \partial_- SM$ and $\partial_+ SM \cap \partial_- SM = S(\partial M)$.

Lemma 2.3. (Cf. [33, Lemmas 3.2.1, 3.2.2].) *For a simple magnetic system, the function $\mathbb{L} : \partial(SM) \rightarrow \mathbb{R}$, defined by*

$$\mathbb{L}(x, \xi) := \begin{cases} \ell(x, \xi) & \text{if } (x, \xi) \in \partial_+ SM, \\ \ell^-(x, \xi) & \text{if } (x, \xi) \in \partial_- SM, \end{cases}$$

is smooth. In particular, $\ell : \partial_+ SM \rightarrow \mathbb{R}$ is smooth. The ratio $\frac{\mathbb{L}(x, \xi)}{\langle \nu(x), \xi \rangle}$ is uniformly bounded on $\partial(SM) \setminus S(\partial M)$.

Proof. Let ρ be a smooth nonnegative function on M such that $\partial M = \rho^{-1}(0)$ and $|\text{grad } \rho| = 1$ in some neighborhood of ∂M . Put $h(x, \xi, t) = \rho(\gamma_{x, \xi}(t))$ for $(x, \xi) \in \partial(SM)$. Then

$$\begin{aligned} h(x, \xi, 0) &= 0, \\ \frac{\partial h}{\partial t}(x, \xi, 0) &= \langle \nu(x), \xi \rangle, \\ \frac{\partial^2 h}{\partial t^2}(x, \xi, 0) &= \text{Hess}_x \rho(\xi, \xi) + \langle \nu(x), Y(\xi) \rangle. \end{aligned}$$

Therefore, for some smooth function $R(x, \xi, t)$,

$$h(x, \xi, t) = \langle \nu(x), \xi \rangle t + \frac{1}{2} (\text{Hess}_x \rho(\xi, \xi) + \langle \nu(x), Y(\xi) \rangle) t^2 + R(x, \xi, t) t^3.$$

Since $h(x, \xi, \mathbb{L}(x, \xi)) = 0$, it follows that $L = \mathbb{L}(x, \xi)$ is a solution of the equation

$$F(x, \xi, L) := \langle \nu(x), \xi \rangle + \frac{1}{2} (\text{Hess}_x \rho(\xi, \xi) + \langle \nu(x), Y(\xi) \rangle) L + R(x, \xi, L) L^2 = 0. \tag{2.6}$$

By (1.3), for $(x, \xi) \in S(\partial M)$

$$\begin{aligned} \frac{\partial F}{\partial L}(x, \xi, 0) &= \frac{1}{2} (\text{Hess}_x \rho(\xi, \xi) + \langle \nu(x), Y(\xi) \rangle) \\ &= \frac{1}{2} (-\Lambda(x, \xi) + \langle \nu(x), Y(\xi) \rangle) < 0. \end{aligned}$$

Now, the implicit function theorem yields smoothness of $\mathbb{L}(x, \xi)$ in a neighborhood of $S(\partial M)$. Since \mathbb{L} is also smooth on $\partial(SM) \setminus S(\partial M)$, we conclude that \mathbb{L} is smooth on $\partial(SM)$.

Next, from (2.6) we get for $(x, \xi) \in \partial(SM) \setminus S(\partial M)$

$$\left[\frac{1}{2} (\text{Hess}_x \rho(\xi, \xi) + \langle v(x), Y(\xi) \rangle) + R(x, \xi, t) \mathbb{L}(x, \xi) \right] \frac{\mathbb{L}(x, \xi)}{\langle v(x), \xi \rangle} = -1.$$

Again by (1.3) this yields boundedness of the ratio $\frac{\mathbb{L}(x, \xi)}{\langle v(x), \xi \rangle}$ for (x, ξ) sufficiently close to $S(\partial M)$ (where \mathbb{L} is sufficiently small), and clearly implies boundedness of the ratio on the whole set $\partial(SM) \setminus S(\partial M)$. \square

Definition 2.4. The scattering relation $\mathcal{S} : \partial_+ SM \rightarrow \partial_- SM$ of a magnetic system (M, g, α) is defined as follows:

$$\mathcal{S}(x, \xi) = (\gamma_{x, \xi}(\ell(x, \xi)), \dot{\gamma}_{x, \xi}(\ell(x, \xi))).$$

The restricted scattering relation $\mathfrak{s} : \partial_+ SM \rightarrow M$ is defined to be the postcomposition of the scattering relation with the natural projection of $\partial_- SM$ to M , i.e.,

$$\mathfrak{s}(x, \xi) = \gamma_{x, \xi}(\ell(x, \xi)).$$

By the preceding lemma, \mathcal{S} and \mathfrak{s} are smooth maps.

The next lemma generalizes the well-known assertion of [22].

Lemma 2.5. Suppose that (g, α) and (g', α') are simple magnetic systems on M such that $g|_{\partial M} = g'|_{\partial M}$. If the boundary action functions $\mathbb{A}|_{\partial M \times \partial M}$ and $\mathbb{A}'|_{\partial M \times \partial M}$ of both the systems coincide, then the scattering relations \mathcal{S} and \mathcal{S}' of these systems coincide, $\mathcal{S} = \mathcal{S}'$.

In the opposite direction, we have the following:

Lemma 2.6. Suppose that (g, α) and (g', α') are simple magnetic systems on M such that $g|_{\partial M} = g'|_{\partial M}$ and $i^* \alpha = i^* \alpha'$. If the restricted scattering relations \mathfrak{s} and \mathfrak{s}' of the systems coincide, $\mathfrak{s} = \mathfrak{s}'$, then the boundary action functions of these systems coincide, $\mathbb{A}|_{\partial M \times \partial M} = \mathbb{A}'|_{\partial M \times \partial M}$.

To prove these lemmas, we need one fruitful result.

Lemma 2.7. If (M, g, α) is a simple magnetic system, then for $x, y \in \partial M$

$$\frac{\partial \mathbb{A}(x, y)}{\partial \xi} = -\langle \dot{\gamma}_{x, y}(0), \xi \rangle + \alpha(\xi) \quad \text{for } \xi \in T_x(\partial M),$$

$$\frac{\partial \mathbb{A}(x, y)}{\partial \eta} = \langle \dot{\gamma}_{x, y}(T), \eta \rangle - \alpha(\eta) \quad \text{for } \eta \in T_y(\partial M),$$

where $\gamma_{x, y} : [0, T] \rightarrow M$ is the unit speed magnetic geodesic from x to y .

Proof. Fix $x, y \in \partial M$, $x \neq y$, and $\xi \in T_x(\partial M)$. Let $\tau(s)$, $-\varepsilon < s < \varepsilon$, be a curve on ∂M with $\tau(0) = x$ and $\dot{\tau}(0) = \xi$. For every s , put $\gamma(t, s) = \gamma_{\tau(s), y}(t)$, $0 \leq t \leq T_s$, denoting $\gamma = \gamma(t, 0)$ and $T = T_0$, and consider

$$c(t, s) = \gamma\left(\frac{T_s}{T}t, s\right), \quad 0 \leq t \leq T.$$

Each curve $c(\cdot, s)$ is defined on the interval $[0, T]$ and its length is exactly T_s . We have

$$\frac{\partial \mathbb{A}(x, y)}{\partial \xi} = \frac{dT_s}{ds}(0) - \frac{d}{ds} \left\{ \int_{\gamma_s} \alpha \right\} \Big|_{s=0}.$$

Using the first variation formula for length and the fact that $c(t, 0) = \gamma(t)$, we have

$$\frac{dT_s}{ds}(0) = -\langle \dot{\gamma}(0), \xi \rangle - \int_0^T \langle \nabla_{\dot{\gamma}} \dot{\gamma}, V \rangle dt,$$

where $V(t) = \frac{\partial c}{\partial s}(t, 0)$ is the variation field of $c(t, s)$. Using the equation of magnetic geodesics (1.2) and the definition (1.1) of Y , we obtain

$$\frac{dT_s}{ds}(0) = -\langle \dot{\gamma}(0), \xi \rangle - \int_0^T \Omega(\dot{\gamma}, V) dt.$$

On the other hand, it is easy to see that

$$\frac{d}{ds} \left\{ \int_{\gamma_s} \alpha \right\} = -\alpha(\xi) + \int_0^T \Omega(V, \dot{\gamma}) dt.$$

This gives the first formula of the lemma. A similar calculation gives the second formula. \square

Proof of Lemma 2.5. By Lemma 2.1, $\alpha(v) = \alpha'(v)$ for all $v \in T(\partial M)$. Using Lemma 2.7 and the fact that both metrics are the same on the boundary, we easily conclude that the scattering relations are the same. \square

Proof of Lemma 2.6. Take $x \in \partial M$ and define $\mathfrak{s}_x : \partial_+ SM \setminus S(\partial M) \rightarrow \partial M \setminus \{x\}$ by $\mathfrak{s}_x(\xi) = \mathfrak{s}(x, \xi)$. This map is a diffeomorphism. Consider its inverse $\mathfrak{s}_x^{-1} : \partial M \setminus \{x\} \rightarrow \partial_+ SM$. By Lemma 2.7, for every $y \in \partial M$, $y \neq x$, we have for all $\xi \in T_x(\partial M)$

$$\frac{\partial \mathbb{A}(x, y)}{\partial \xi} = -\langle \mathfrak{s}_x^{-1}(y), \xi \rangle_g + \alpha(\xi). \tag{2.7}$$

The assumptions of the lemma imply that the right-hand side of (2.7) is the same for the second magnetic system; therefore, so does the left-hand side. Now, the claim of the lemma is immediate. \square

Thus, for simple magnetic systems, the boundary rigidity problem is essentially equivalent to the problem of restoring a Riemannian metric and a magnetic potential from the restricted scattering relation.

2.3. Determination of volume

Here we show that the boundary action function determines the volume of the manifold. This generalizes the well-known assertion of [22, Proposition 2.13].

Theorem 2.8. *If (g, α) and (g', α') are simple magnetic systems on M with the same boundary action function, then the volume $\text{Vol}_g M$ of M w.r.t. g equals the volume $\text{Vol}_{g'} M$ of M w.r.t. g' .*

Proof. By Santaló’s formula (see (A.4) in Appendix A) we have

$$\text{Vol}_g M = \frac{1}{w_{n-1}} \int_{\partial_+ SM} \ell(x, \xi) d\mu(x, \xi).$$

Using

$$\int_{SM} \alpha(x, \xi) d\Sigma^{2n-1}(x, \xi) = 0 \tag{2.8}$$

and Santaló’s formula again, we obtain

$$\text{Vol}_g M = \frac{1}{w_{n-1}} \int_{\partial_+ SM} \mathbb{A}(\gamma_{x,\xi}) d\mu(x, \xi). \tag{2.9}$$

In view of Theorem 2.2 and Lemma 2.5, we may assume that the right-hand side of this equality is the same for the magnetic system (g', α') . This yields the sought equality of volumes. \square

3. Magnetic ray transform

3.1. Derivatives of the action function

Let (g, α) be a simple magnetic system on M . It is easy to see that there is an $\varepsilon > 0$ so that every magnetic system $(g + h, \alpha + \beta)$, satisfying

$$\|h\|_{C^2} \leq \varepsilon, \quad \|\beta\|_{C^1} \leq \varepsilon, \tag{3.1}$$

is simple.

Given h and β satisfying (3.1), consider the 1-parameter family (g^s, α^s) with

$$g^s = g + sh, \quad \alpha^s = \alpha + s\beta, \quad s \in [0, 1].$$

Clearly, each of these systems is simple.

Lemma 3.1. For $x, y \in \partial M$,

$$\frac{d\mathbb{A}_{g^s, \alpha^s}(x, y)}{ds} = \frac{1}{2} \int_{\gamma_s} \langle h, \dot{\gamma}_s^2 \rangle - \int_{\gamma_s} \beta, \tag{3.2}$$

where γ_s is the unit speed magnetic geodesic from x to y w.r.t. (g^s, α^s) .

If

$$h|_{\partial M} = 0, \quad \beta|_{\partial M} = 0, \tag{3.3}$$

then

$$\left| \frac{d^2\mathbb{A}_{g^s, \alpha^s}(x, y)}{ds^2} \right| \leq C(\|h\|_{C^1}^2 + \|\beta\|_{C^1}^2), \tag{3.4}$$

with a constant C independent of h and β and C^2 locally uniform in (g, α) .

Proof. Define

$$\varphi(s, \tau) := \mathbb{A}_{g^\tau, \alpha^\tau}(\gamma_s) = \frac{1}{2} \int_0^{T_s} |\dot{\gamma}_s(t)|_{g^\tau}^2 dt + \frac{1}{2} T_s - \int_{\gamma_s} \alpha^\tau.$$

Then

$$\frac{d\mathbb{A}_{g^s, \alpha^s}(x, y)}{ds} = \frac{\partial \varphi}{\partial s}(s, s) + \frac{\partial \varphi}{\partial \tau}(s, s). \tag{3.5}$$

By Lemma A.5, unit speed magnetic geodesics minimize the time-free action; therefore, for a fixed τ , $\mathbb{A}_{g^\tau, \alpha^\tau}(\gamma_s)$ has a minimum at $s = \tau$, which yields

$$\frac{\partial \varphi}{\partial s}(\tau, \tau) = 0. \tag{3.6}$$

Next,

$$\frac{\partial \varphi}{\partial \tau} = \frac{1}{2} \int_0^{T_s} \left(\frac{\partial}{\partial \tau} |\dot{\gamma}_s(t)|_{g^\tau}^2 \right) dt - \int_{\gamma_s} \frac{d}{d\tau} \alpha^\tau = \frac{1}{2} \int_{\gamma_s} \langle h, \dot{\gamma}_s^2 \rangle - \int_{\gamma_s} \beta. \tag{3.7}$$

Combining (3.5)–(3.7) gives (3.2).

Differentiating (3.2) and using (3.3), we obtain

$$\frac{d^2\mathbb{A}_{g^s, \alpha^s}(x, y)}{ds^2} = \int_0^{T_s} \left\{ \frac{1}{2} (\nabla_{\gamma'_s} h_{ij}) \dot{\gamma}_s^i \dot{\gamma}_s^j + h_{ij} \dot{\gamma}_s^i (\nabla_{\gamma'_s} \dot{\gamma}_s^j) - (\nabla_{\gamma'_s} \beta_i) \dot{\gamma}_s^i - \beta_i (\nabla_{\gamma'_s} \dot{\gamma}_s^i) \right\} dt,$$

where $\gamma'_s = \partial\gamma_s/\partial s$. Whence

$$\left| \frac{d^2 \mathbb{A}_{g^s, \alpha^s}(x, y)}{ds^2} \right| \leq C(\|h\|_{C^1} + \|\beta\|_{C^1})(\|\gamma'_s\|_C + \|\nabla_{\gamma'_s} \dot{\gamma}_s\|_C). \tag{3.8}$$

By the equation of magnetic geodesics, we have

$$\overset{s}{\nabla}_{\dot{\gamma}_s} \dot{\gamma}_s = \overset{s}{Y}(\dot{\gamma}_s), \tag{3.9}$$

where $\overset{s}{\nabla}$ stands for the covariant derivative related to g^s , and $\overset{s}{Y}$ is the Lorentz force associated with g^s, α_s . The initial conditions are given by

$$\gamma_s(0) = x, \quad \dot{\gamma}_s(0) = (\text{exp}_x^\mu)^{-1}(y), \tag{3.10}$$

where exp_x^μ stands for the magnetic exponential map associated with g^s, α_s . From (3.9) and (3.10) we easily infer that

$$\|\gamma'_s\|_C + \|\nabla_{\gamma'_s} \dot{\gamma}_s\|_C \leq C(\|h\|_{C^1} + \|\beta\|_{C^1}). \tag{3.11}$$

Combining (3.8) and (3.11) leads to (3.4). \square

3.2. Magnetic ray transform

Let (M, g, α) be a simple magnetic system and $\phi : SM \rightarrow \mathbb{R}$ a smooth function on the unit sphere bundle. We define the *magnetic ray transform* of ϕ to be the following function on the space of unit speed magnetic geodesics going from a boundary point to a boundary point:

$$I\phi(\gamma) = \int_\gamma \phi := \int_0^T \phi(\gamma(t), \dot{\gamma}(t)) dt,$$

where $\gamma : [0, T] \rightarrow M$ is any unit speed magnetic geodesic such that $\gamma(0) \in \partial M$ and $\gamma(T) \in \partial M$. Assuming that the magnetic geodesics are parametrized by $\partial_+ SM$, we obtain a map $I : C^\infty(SM) \rightarrow C(\partial_+ SM)$,

$$I\phi(x, \xi) = \int_0^{\ell(x, \xi)} \phi(\psi^t(x, \xi)) dt, \quad (x, \xi) \in \partial_+ SM. \tag{3.12}$$

In the space of real-valued functions on $\partial_+ SM$ define the norm

$$\|\phi\|^2 = \int_{\partial_+ SM} \phi^2 d\mu$$

and the corresponding inner product. Here $d\mu(x, \xi) = \langle \xi, \nu(x) \rangle d\Sigma^{2n-2}$ (see A.4 in Appendix A). Denote the corresponding Hilbert space by $L^2_\mu(\partial_+ SM)$.

Lemma 3.2. *The operator I extends to a bounded operator*

$$I : L^2(SM) \rightarrow L^2_\mu(\partial_+ SM).$$

Proof. Indeed, it is easy to see that $(I\phi)^2 \leq C I\phi^2$, with some constant C independent of ϕ . Therefore,

$$\int_{\partial_+ SM} (I\phi)^2 d\mu \leq C \int_{\partial_+ SM} I\phi^2 d\mu = C \int_{SM} \phi^2 d\Sigma^{2n-1}$$

by the Santaló formula (A.4). \square

3.3. Solenoidal and potential pairs

For a decomposition of a symmetric tensor field into a potential and a solenoidal part (relevant when $\Omega = 0$), we refer to [33,34].

In view of the linearization formula (3.2), we are mainly interested in I applied to functions ϕ that are of the form

$$\phi(x, \xi) = h_{ij}(x)\xi^i\xi^j + \beta_j(x)\xi^j \tag{3.13}$$

(and, more generally, that are polynomials in ξ). Then, given a symmetric 2-tensor h and a 1-form β , we set for $(x, \xi) \in \partial_+ SM$

$$I[h, \beta](x, \xi) = \int_0^T h_{ij}(\gamma(t))\dot{\gamma}^i(t)\dot{\gamma}^j(t) dt + \int_0^T \beta_j(\gamma(t))\dot{\gamma}^j(t) dt, \tag{3.14}$$

where $\gamma = \gamma_{x,\xi}$, $T = \ell(x, \xi)$.

If F is a notation for a function space (C^k , L^p , H^k , etc.), then we will denote by $\mathbf{F}(M)$ the corresponding space of pairs $\mathbf{f} = [h, \beta]$, with h a symmetric covariant 2-tensor and β a 1-form, and denote by $\mathcal{F}(M)$ the corresponding space of pairs $\mathbf{w} = [v, \varphi]$, with v a 1-form and φ a function on M . In particular, $\mathbf{L}^2(M)$ is the space of square integrable pairs $\mathbf{f} = [h, \beta]$, and we endow this space with the norm

$$\|\mathbf{f}\|^2 = \int_M \left\{ |h|_g^2 + \frac{n-1}{2} |\beta|_g^2 \right\} d \text{Vol}, \tag{3.15}$$

with the corresponding inner product. (The choice of the factor $(n-1)/2$ will play its role in the proof of Theorem 5.4.) In the space $\mathcal{L}^2(M)$ we will consider the norm

$$\|\mathbf{w}\|^2 = \int_M (|v|_g^2 + \varphi^2) d \text{Vol}. \tag{3.16}$$

Clearly, the norm of a pair $[h, \beta]$ in $L^2(M)$ is equivalent to the norm of the corresponding polynomial (3.13) in $L^2(SM)$. Therefore, Lemma 3.2 implies that

$$I : L^2(M) \rightarrow L^2_\mu(\partial_+ SM)$$

and it is bounded.

Define $Y : T^*M \rightarrow T^*M$ by

$$Y(\eta) = -(Y_i^j \eta_j), \quad \eta = (\eta_j) \in T^*M.$$

Note that this definition agrees with the map $Y : TM \rightarrow TM$ and the isomorphism between the tangent and cotangent bundles via the Riemannian metric.

Clearly, I vanishes on functions $\phi(x, \xi) = \mathbf{G}_\mu \psi(x, \xi)$ if ψ vanishes for $x \in \partial M$. To find a class of such functions that are polynomials of ξ of degree at most 2, assume that

$$\mathbf{G}_\mu[v_i(x)\xi^i + \varphi(x)] = h_{ij}(x)\xi^i\xi^j + \beta_j(x)\xi^j.$$

Since

$$\mathbf{G}_\mu(v_i(x)\xi^i + \varphi(x)) = \xi^j[v_{i,j}\xi^i + v_i Y_j^i + \varphi_{,j}] = (d^s v)_{ij}\xi^i\xi^j + (\varphi_{,j} - Y(v)_j)\xi^j, \quad (3.17)$$

we get

$$h = d^s v, \quad \beta = d\varphi - Y(v), \quad (3.18)$$

where $d^s v$ is the symmetric differential of v . We used here the fact that the even part of (3.17) w.r.t. $\xi \in S_x M$ determines the quadratic form $(d^s v)_{ij}\xi^i\xi^j$, while the odd part determines $\beta(\xi)$. Next, knowing $(d^s v)_{ij}\xi^i\xi^j$ for $n(n+1)/2$ generic $\xi \in S_x M$ is enough to recover v ; similarly knowing $\beta(\xi)$ for n linear independent $\xi \in S_x M$ is enough to recover β . We have the following stronger statement.

Lemma 3.3. Fix $x \in M$. Given any open subset V of $S_x M$, there exist $N = n(n+1)/2 + n$ vectors $\xi_m \in V$, $m = 1, \dots, N$, such that $[h, \beta]$ is uniquely determined by the values of $\psi(x, \xi) := h_{ij}(x)\xi^i\xi^j + \beta_j(x)\xi^j$ at $\xi = \xi_m$, $m = 1, \dots, N$.

Proof. Since x is fixed, we denote $\psi(\xi) = \psi(x, \xi)$. We show first that $\psi(\xi)$, known for all $\xi \in S_x M$, determines $[h, \beta]$. Since $\psi(\xi)$ is a linear functional of $[h, \beta]$, and the latter belongs to a linear space that can be identified with \mathbb{R}^N , it is enough to show that $\psi(\xi) = 0$ implies $h = 0$, $\beta = 0$. By replacing ξ by $-\xi$, we get that $h_{ij}\xi^i\xi^j = 0$, $\beta_j\xi^j = 0$. The second relation easily implies that $\beta = 0$. The first one implies $h = 0$ easily as well (see, e.g., [33], where also a sharp estimate is established).

Next, assume that $\psi(\xi) = 0$ in V . Then $\psi(\xi) = 0$ on $S_x M$ by analytic continuation, thus $[h, \beta] = 0$. Finally, since for any fixed ξ , $\psi(\xi)$ is a linear functional belonging to $(\mathbb{R}^N)' \cong \mathbb{R}^N$, and $\{\psi(\xi); \xi \in V\}$ is a complete set, there exists a basis $\psi(\xi_k)$, $k = 1, \dots, N$, in it. \square

Remark 3.4. We can say a bit more. Since the determination of $[h, \beta]$ is done by inverting a linear transform, one can choose ξ_m continuously depending on x , if x belongs to a small enough set X ,

such that the map $\{\psi(x, \xi_m(x)), m = 1, \dots, N\} \mapsto [h, \beta]$ is invertible with a uniform bound on the inverse. Then this can be extended to compact sets X .

Definition 3.5. We call a pair $[h, \beta] \in \mathbf{L}^2(M)$ potential if Eqs. (3.18) hold with $[v, \varphi] \in \mathcal{H}_0^1(M)$.

This can be written as follows:

$$\begin{pmatrix} h \\ \beta \end{pmatrix} = \mathbf{d} \begin{pmatrix} v \\ \varphi \end{pmatrix}, \quad \mathbf{d} := \begin{pmatrix} d^s & 0 \\ -Y & d \end{pmatrix},$$

where d^s stands for the symmetric derivative acting on covector fields and d is the usual differential acting on functions (which coincides with the symmetric differential on functions).

Clearly, potential pairs satisfy

$$I(\mathbf{d}[v, \varphi]) = 0. \tag{3.19}$$

This follows from (3.17) if v, φ are smooth, and follows by continuity for general v, φ , once we establish the mapping properties of $N = I^*I$ below.

We will relate potential pairs to the non-linear problem. Let (g, α) and (g', α') be two gauge equivalent pairs, i.e., $g' = f^*g$ and $\alpha' = f^*\alpha + d\varphi$ with some diffeomorphism $f : M \rightarrow M$, fixing ∂M , and some function φ vanishing on ∂M . Linearize this near $f = \text{Id}$ and $\varphi = 0$. In other words, let f_τ be a smooth family of such diffeomorphisms with $f_0 = \text{Id}$ and let φ^τ be a smooth family of such functions with $\varphi^0 = 0$. Let $g^\tau = f_\tau^*g$, $\alpha^\tau = f_\tau^*\alpha + d\varphi^\tau$, and we will compute the derivatives at $\tau = 0$. It is well known [33, (3.1.5)] and is easy to calculate that $dg^\tau/d\tau|_{\tau=0} = 2d^s v$, where $v = df_\tau/d\tau|_{\tau=0}$. Let $d\varphi^\tau/d\tau|_{\tau=0} = \psi$. Since $\alpha_i^\tau = (\alpha_j \circ f_\tau)\partial f_\tau^j/\partial x^i + \partial\varphi^\tau/\partial x^i$, we get

$$\begin{aligned} \beta_i &:= \left. \frac{d}{d\tau} \right|_{\tau=0} \alpha_i^\tau = v^j \partial_j \alpha_i + \alpha_j \partial_i v^j + \psi_{,i} \\ &= v^j \alpha_{i,j} + \alpha_j v_{,i}^j + \psi_{,i} \\ &= v^j \alpha_{i,j} + (\alpha_j v^j)_{,i} - v^j \alpha_{j,i} + \psi_{,i} \\ &= -(d\alpha)_{ij} v^j + (\alpha_j v^j + \psi)_{,i}. \end{aligned}$$

Using (1.4), (1.1), and treating v as a 1-form (by lowering the index), we obtain

$$\beta = Y(v) + d(\langle \alpha, v \rangle + \psi).$$

This shows that

$$\left. \frac{d}{d\tau} \right|_{\tau=0} \left[\frac{1}{2} g^\tau, -\alpha^\tau \right] = \mathbf{d}[v, -\langle \alpha, v \rangle - \psi], \quad \text{where } v = \left. \frac{df_\tau}{d\tau} \right|_{\tau=0}, \quad \psi = \left. \frac{d\varphi^\tau}{d\tau} \right|_{\tau=0}. \tag{3.20}$$

Since ψ can be arbitrary (as long as it vanishes on ∂M), we see that the linearization of (g', α') near $f = \text{Id}, \varphi = 0$ is given by $\mathbf{d}[v, \psi]$ with v, ψ that can be arbitrary.

For future references, let us mention that by (3.17)

$$\mathbf{G}_\mu \langle \mathbf{w}, \xi \rangle = \langle \mathbf{d}\mathbf{w}, \xi \rangle \implies \frac{d}{dt} \langle \mathbf{w}(\gamma), \dot{\gamma} \rangle = \langle \mathbf{d}\mathbf{w}(\gamma), \dot{\gamma} \rangle \tag{3.21}$$

for any unit speed magnetic geodesic $\gamma(t)$, where $\mathbf{w} = [v, \varphi]$, and we used the notation $\langle \mathbf{f}, \xi \rangle = \langle h, \xi^2 \rangle + \langle \beta, \xi \rangle$, where $\mathbf{f} = [h, \beta]$, and similarly $\langle \mathbf{w}, \xi \rangle = \langle v, \xi \rangle + \varphi$.

Definition 3.6. We say that I is s-injective if $I[h, \beta] = 0$ with $[h, \beta] \in \mathbf{L}^2(M)$ implies that $[h, \beta]$ is potential, i.e.,

$$h = d^s v, \quad \beta = -Y(v) + d\varphi \tag{3.22}$$

with $[v, \varphi] \in \mathcal{H}_0^1(M)$.

Let $[h, \beta]$ be orthogonal to all potential pairs. Then

$$\int_M \left\{ (h, d^s v) + \frac{n-1}{2} (\beta, -Y(v) + d\varphi) \right\} d \text{Vol} = 0$$

for all v, φ vanishing on ∂M . Then

$$-\int_M \left\{ \left(\delta h - \frac{n-1}{2} Y(\beta), v \right) + \frac{n-1}{2} (\delta\beta)\varphi \right\} d \text{Vol} = 0,$$

where δ is the divergence. Therefore,

$$\delta h - \frac{n-1}{2} Y(\beta) = 0, \quad \delta\beta = 0. \tag{3.23}$$

Definition 3.7. We call a pair $[h, \beta]$ solenoidal if Eqs. (3.23) hold.

This can be written as

$$\delta \begin{pmatrix} h \\ \beta \end{pmatrix} = 0, \quad \delta := \begin{pmatrix} \delta & -\frac{n-1}{2} Y \\ 0 & \delta \end{pmatrix}.$$

Then $\mathbf{d} = -\delta^*$.

In terms of the operators δ, \mathbf{d} , the solenoidal pairs are defined as the ones in $\text{Ker } \delta$, and the orthogonal complement of $\text{Ker } \delta$ is $\text{Ran } \mathbf{d}$, consisting of potential pairs. Next, we will describe the projections to solenoidal pairs.

We will show first that the operator $-\delta\mathbf{d}$ is an elliptic second order (and clearly, a formally self-adjoint) operator, acting on pairs $\mathbf{w} = [v, \varphi]$, where v is a 1-form and φ is a function on M . First, notice that

$$(-\delta\mathbf{d}\mathbf{w}, \mathbf{w}) = \|\mathbf{d}\mathbf{w}\|^2, \quad \mathbf{w} \in C_0^\infty(M^{\text{int}}); \tag{3.24}$$

thus, $-\delta\mathbf{d}$ is a non-negative operator. Let $\sigma_p(P)$ stand for the principal symbol of P . Then (3.24) implies that for any such \mathbf{w} and a fixed x

$$(-\sigma_p(\delta\mathbf{d})\mathbf{w}, \mathbf{w}) = \|\sigma_p(\mathbf{d})\mathbf{w}\|^2, \tag{3.25}$$

where the inner product and the norm are in the finite dimensional space to which \mathbf{w} belongs, i.e., they are as in (3.16) before the integration. We will show that $-\sigma_p(\delta\mathbf{d})$ is in fact a positive matrix-valued symbol for $\xi \neq 0$. Since it is homogeneous in the dual variable ξ , it is enough to show that $-\sigma_p(\delta\mathbf{d})\mathbf{w} = 0$ implies $\mathbf{w} = 0$ for $\xi \neq 0$. In view of (3.25), if $-\sigma_p(\delta\mathbf{d})\mathbf{w} = 0$, then $\sigma_p(\mathbf{d})\mathbf{w} = 0$. Then $\sigma_p(d^s)v = 0$ (see (4.6)), $\xi\varphi = 0$, where $[v, \varphi] = \mathbf{w}$. It is well known [33,34], and easy to see directly, that d^s is elliptic, therefore $v = 0$. Next, $\xi\varphi = 0$ implies $\varphi = 0$ as well.

Note that the Dirichlet boundary conditions are coercive for $\delta\mathbf{d}$, because the latter is positive elliptic. We will show that the kernel and the cokernel of this elliptic problem are trivial. If \mathbf{w} belongs to the kernel of $\delta\mathbf{d}$, and satisfies Dirichlet boundary conditions, it has to be smooth and then (3.24) holds for it as well. Then $\mathbf{d}\mathbf{w} = 0$ and $\mathbf{w} = 0$ on ∂M . This easily implies $\mathbf{w} = 0$ in M by integrating (3.21) along magnetic geodesics connecting boundary and interior points, and using Lemma 3.3. Using standard arguments, one easily checks that the cokernel of $\delta\mathbf{d}$ equipped with Dirichlet boundary conditions, is trivial as well. Indeed, fix \mathbf{u} in the cokernel. Since \mathbf{u} is orthogonal to $\delta\mathbf{d}\mathbf{w}$ for all $\mathbf{w} \in C_0^\infty(M^{\text{int}})$, we get that $\delta\mathbf{d}\mathbf{u} = 0$. This in particular implies that \mathbf{u} has a trace on ∂M , and choosing \mathbf{w} 's with a dense set of normal derivatives on ∂M , we show that this trace vanishes. Therefore, $\mathbf{u} = 0$ by what we proved above. Denote by $(\delta\mathbf{d})_D$ the Dirichlet realization of $\delta\mathbf{d}$ on M . The arguments above show that $(\delta\mathbf{d})_D$ is an invertible self-adjoint operator.

We define the following projections

$$\mathcal{P} = \mathbf{d}(\delta\mathbf{d})_D^{-1}\delta, \quad \mathcal{S} = \text{Id} - \mathcal{P}. \tag{3.26}$$

Then $\mathbf{f}^s = \mathcal{S}\mathbf{f}$ is solenoidal, $\mathcal{P}\mathbf{f} = \mathbf{d}\mathbf{w}$ is potential, and we have the orthogonal decomposition

$$\mathbf{f} = \mathbf{f}^s + \mathbf{d}\mathbf{w}$$

into the solenoidal and potential parts.

Since I vanishes on $\mathcal{P}\mathbf{L}^2(M)$, the s -injectivity of I is then equivalent to the following: I is injective on $\mathcal{S}\mathbf{L}^2(M)$.

3.4. Adjoint of I

For a fixed simple (g, α) , consider

$$I : \mathbf{L}^2(M) \rightarrow L_\mu^2(\partial_+ SM)$$

and consider its dual

$$I^* : L_\mu^2(\partial_+ SM) \rightarrow \mathbf{L}^2(M).$$

We will now find an expression for I^* . Let $\mathbf{f} = [h, \beta]$ and $\phi(x, \xi)$ be as in (3.13). Let $\psi(x, \xi) \in C(\partial_+ SM)$. Then

$$(\mathbf{I}\mathbf{f}, \psi) = \int_{\partial_+ SM} \psi(x, \xi) d\mu(x, \xi) \int_0^{\ell(x, \xi)} [h_{ij}(\gamma_{x, \xi}(t)) \dot{\gamma}_{x, \xi}^i(t) \dot{\gamma}_{x, \xi}^j(t) + \beta_i(\gamma_{x, \xi}(t)) \dot{\gamma}_{x, \xi}^i(t)] dt.$$

By Santaló’s formula (A.4) of Appendix A, we get

$$(\mathbf{I}\mathbf{f}, \psi) = \int_{SM} [h_{ij}(x) \xi^i \xi^j + \beta_i(x) \xi^i] \psi^\sharp(x, \xi) d\Sigma^{2n-1}(x, \xi),$$

where $\psi^\sharp(x, \xi)$ is defined as the function that is constant along the orbits of the magnetic flow and that equals $\psi(x, \xi)$ on $\partial_+ SM$. Let $d\sigma_x(\xi)$ be the measure on $S_x M$. Then

$$\begin{aligned} (\mathbf{I}\mathbf{f}, \psi) &= \int_M h_{ij}(x) \int_{S_x M} \xi^i \xi^j \psi^\sharp(x, \xi) d\sigma_x(\xi) d\text{Vol}(x) \\ &\quad + \int_M \beta_i(x) \int_{S_x M} \xi^i \psi^\sharp(x, \xi) d\sigma_x(\xi) d\text{Vol}(x). \end{aligned}$$

Therefore, see also [35],

$$I^* \psi = \left[\int_{S_x M} \xi^i \xi^j \psi^\sharp(x, \xi) d\sigma_x(\xi), \frac{2}{n-1} \int_{S_x M} \xi^i \psi^\sharp(x, \xi) d\sigma_x(\xi) \right]. \tag{3.27}$$

3.5. Integral representation for the normal operator

Let $M_1 \supset M$ be another manifold with boundary (a domain in \mathbb{R}^n , actually) such that $M_1^{\text{int}} \supset M$. Extend g, α to M_1 . Then (M_1, g, α) is still simple if M_1 is close enough to M .

Choose $x_0 \in M_1 \setminus M$, and consider the map

$$(\exp_{x_0}^\mu)^{-1} : M_1 \rightarrow (\exp_{x_0}^\mu)^{-1}(M_1).$$

It is C^∞ away from $x = x_0$, therefore it is a C^∞ diffeomorphism to its image, if we restrict it to any compact submanifold $M_{1/2}$ of M_1^{int} . We can assume that $M_{1/2}$ has a smooth boundary, and that $M_{1/2}^{\text{int}} \supset M$. We denote $M_{1/2}$ by M_1 again, and the map above gives us global coordinates in a neighborhood of M_1 that can be identified with a subset of \mathbb{R}^n with a smooth boundary. So in this section, M_1 is considered as a subset of \mathbb{R}^n . If g, α are analytic, and M is analytic, then we will also assume that ∂M_1 is analytic.

Denote by I_1 the magnetic ray transform on $L^2(M_1)$,

$$I_1 : L^2(M_1) \rightarrow L^2_\mu(\partial_+ SM_1),$$

and denote by I_1^* its dual. Set

$$N = I_1^* I_1. \tag{3.28}$$

Extending all tensors as 0 on M_1 , we consider $\mathbf{L}^2(M)$ as a subspace of $\mathbf{L}^2(M_1)$. In this way we may consider N as the operator

$$N : \mathbf{L}^2(M) \rightarrow \mathbf{L}^2(M_1).$$

We then define s-injectivity of N as in Definition 3.6.

Lemma 3.8. $N : \mathbf{L}^2(M) \rightarrow \mathbf{L}^2(M_1)$ is s-injective if and only if I is s-injective.

Proof. Let N be s-injective. Assume that $I\mathbf{f} = 0$ with $\mathbf{f} \in \mathbf{L}^2(M)$. Then $I_1\mathbf{f} = 0$ on ∂_+SM_1 as well, therefore $N\mathbf{f} = 0$, which implies that \mathbf{f} is potential.

Now, assume that I is s-injective and let $N\mathbf{f} = 0$. Then

$$0 = (N\mathbf{f}, \mathbf{f})_{\mathbf{L}^2(M_1)} = \|I_1\mathbf{f}\|_{L^2_\mu(\partial_+SM_1)}^2,$$

therefore, $I\mathbf{f} = 0$ as an element of $L^2_\mu(\partial_+SM)$ as well, hence \mathbf{f} is potential. \square

We will find an integral representation of N and will show that it is a Ψ DO in a neighborhood of M , following [36–38].

Using (3.27), and replacing ξ by v , we get the following.

Proposition 3.9.

$$N\mathbf{f} = \left[N_{22}h + N_{21}\beta, \frac{2}{n-1}(N_{12}h + N_{11}\beta) \right], \tag{3.29}$$

where

$$\begin{aligned} (N_{11}\beta)^{i'}(x) &= \int_{S_xM_1} v^{i'} d\sigma_x(v) \int \beta_i(\gamma_{x,v}(t)) \dot{\gamma}_{x,v}^i(t) dt, \\ (N_{12}h)^{i'j'}(x) &= \int_{S_xM_1} v^{i'} d\sigma_x(v) \int h_{ij}(\gamma_{x,v}(t)) \dot{\gamma}_{x,v}^i(t) \dot{\gamma}_{x,v}^j(t) dt, \\ (N_{21}\beta)^{i'j'}(x) &= \int_{S_xM_1} v^{i'} v^{j'} d\sigma_x(v) \int \beta_i(\gamma_{x,v}(t)) \dot{\gamma}_{x,v}^i(t) dt, \\ (N_{22}h)^{i'j'j'}(x) &= \int_{S_xM_1} v^{i'} v^{j'} d\sigma_x(v) \int h_{ij}(\gamma_{x,v}(t)) \dot{\gamma}_{x,v}^i(t) \dot{\gamma}_{x,v}^j(t) dt. \end{aligned} \tag{3.30}$$

In each of the integrals above, split the t -integral into two parts: I_+ corresponding to $t > 0$, and I_- , corresponding to $t < 0$. In I_+ , make the change of variables $y = \gamma_{x,v}(t) = \exp_x^\mu(tv)$. Then $tv = (\exp_x^\mu)^{-1}(y)$, and t, v are C^∞ functions of x, y away from the diagonal $x = y$. We treat I_- in a similar way by making the substitution $t' = -t, v' = -v$. To this end, note that the simplicity assumption implies that the map

$$\exp_x^{\mu,-}(tv) = \pi \circ \psi^{-t}(-v), \quad t \geq 0, v \in S_xM_1, \tag{3.31}$$

is also a C^1 diffeomorphism for every $x \in M_1$.

We analyze the Schwartz kernel of N near the diagonal below, but at this point we just want to emphasize that it is smooth away from the diagonal and therefore N has the pseudolocal property.

4. Analysis of the operator N . Generic s-injectivity

In this section, we analyze N and prove s-injectivity for analytic (g, α) , and for generic ones.

Here we first assume that $[g, \alpha] \in \mathbf{C}^k(M)$. As a result, various objects related to g, α will have smoothness $l(k)$ such that $l(k) \rightarrow \infty$ as $k \rightarrow \infty$. To simplify the exposition, we will not try to estimate $l(k)$ (actually, $l(k) = k - k_0$ with k_0 depending on n only). We will say that a given function (or tensor field) f is smooth, or that $f \in C^l$, if such an $l = l(k)$ exists, and l may vary from line to line. In particular, if $k = \infty$, then $l = \infty$.

At the end of the section we analyze the case in which (g, α) is an analytic magnetic system.

4.1. More about the magnetic exponential map at the origin

In order to study the singularities of the kernel of N near $x = y$, we need more precise information about the exponential map at the origin. One of the interesting features of the magnetic problem is that the magnetic exponential map is not C^2 unless $\Omega = 0$ by Lemma A.7. On the other hand, in polar coordinates, it is a smooth map. We are therefore forced to work in polar coordinates.

Consider $y = \exp_x^\mu(tv) = \gamma_{x,v}(t)$, where $v \in S_x M_1$, and $t \geq 0$ are such that $y \in M_1$. Recall that M_1 now is a subdomain of \mathbb{R}^n . We are interested in the behavior of y for small $|t|$ near a fixed x_0 , therefore, we can assume that we work in $U = \{x; |y - x_0| \leq \varepsilon\}$ with $0 \leq \varepsilon \ll 1$ such that U is strictly convex w.r.t. the magnetic geodesics as well as w.r.t. the Euclidean metric.

By Lemma A.7, the map $tv = \xi \mapsto y$ is not C^2 in general. On the other hand, the map $(t, v) \mapsto y$ is smooth, and respectively analytic, if g is analytic, and in fact extends as a smooth/analytic map for small negative t as well, by the formula $y = \gamma_{x,v}(t)$. Similarly, the function $m(t, v; x) = (\gamma_{x,v}(t) - x)/t$ has the same smoothness, therefore

$$\gamma_{x,v}(t) - x = tm(t, v; x), \quad m(0, v; x) = v. \tag{4.1}$$

We introduce the new variables $(r, \omega) \in \mathbb{R} \times S_x U$ by

$$r = t|m(t, v; x)|_g, \quad \omega = m(t, v; x)/|m(t, v; x)|_g. \tag{4.2}$$

Then (r, ω) are polar coordinates for $y - x = r\omega$ in which we allow r to be negative. Clearly, (r, ω) are smooth/analytic at least for ε small enough, if g is smooth/analytic. Consider the Jacobian of this change of variables

$$J := \det \frac{\partial(r, \omega)}{\partial(t, v)}.$$

It is not hard to see that $J|_{t=0} = 1$, therefore the map $\mathbb{R} \times S_x U \ni (t, v) \mapsto (r, \omega) \in \mathbb{R} \times S_x U$ is a local C^l , respectively analytic, diffeomorphism from $(\text{a neighborhood of } 0) \times S_x U$ to its image. We can decrease ε if needed to ensure that it is a (global) diffeomorphism on its domain because

then it is clearly injective. We denote the inverse functions by $t = t(r, \omega)$, $v = v(r, \omega)$. Note that in the (r, ω) variables

$$t = r + O(|r|), \quad v = \omega + O(|r|), \quad \dot{\gamma}_{x,v}(t) = \omega + O(|r|). \tag{4.3}$$

Another representation of the new coordinates can be given by writing

$$\text{Exp}_x^\mu(t, v) = \gamma_{x,v}(t).$$

Then

$$r = \text{sign}(t) |\text{Exp}_x^\mu(t, v) - x|_g, \quad \omega = \text{sign}(t) \frac{\text{Exp}_x^\mu(t, v) - x}{|\text{Exp}_x^\mu(t, v) - x|_g},$$

and

$$(t, v) = (\text{Exp}_x^\mu)^{-1}(x + r\omega)$$

with the additional condition that r and t have the same sign (or are both zero).

4.2. Principal symbol of N

Since we showed that the Schwartz kernel is smooth away from the diagonal, and we want to prove eventually that N is a Ψ DO, it is enough to study the restriction of N on a small enough set U as above. We analyze N_{22} , for example. Perform the change of variables $(t, v) \mapsto (r, \omega)$ in (3.30), to get

$$\begin{aligned} (N_{22}h)^{i'j'}(x) &= \int_{S_x U} \int_{\mathbb{R}} v^{i'}(r, \omega; x) v^{j'}(r, \omega; x) h_{ij}(x + r\omega) w^i(r, \omega; x) w^j(r, \omega; x) \\ &\quad \times J^{-1}(r, \omega; x) dr d\sigma_x(\omega), \end{aligned} \tag{4.4}$$

where $w(r, \omega, x) = \dot{\gamma}_{x,v}(t)$, and $\text{supp } h \subset U$.

This type of integral operators is studied in Appendix B. Applying Lemma B.1 and the remark after it to N_{22} , see (4.4), and proceeding similarly for the other operators N_{kl} , we get

Proposition 4.1. N_{kl} , $k, l = 1, 2$, are Ψ DOs in M_1^{int} with principal symbols

$$\begin{aligned} \sigma_p(N_{11})^{i'i}(x, \xi) &= 2\pi \int_{S_x U} \omega^{i'} \omega^i \delta(\omega \cdot \xi) d\sigma_x(\omega), \\ \sigma_p(N_{12})^{i'ij}(x, \xi) &= 2\pi \int_{S_x U} \omega^{i'} \omega^i \omega^j \delta(\omega \cdot \xi) d\sigma_x(\omega) = 0, \\ \sigma_p(N_{21})^{i'j'i}(x, \xi) &= 2\pi \int_{S_x U} \omega^{i'} \omega^{j'} \omega^i \delta(\omega \cdot \xi) d\sigma_x(\omega) = 0, \\ \sigma_p(N_{22})^{i'j'ij}(x, \xi) &= 2\pi \int_{S_x U} \omega^{i'} \omega^{j'} \omega^i \omega^j \delta(\omega \cdot \xi) d\sigma_x(\omega). \end{aligned} \tag{4.5}$$

Here $\omega \cdot \xi = \omega^j \xi_j$. The reason for $\sigma_p(N_{12}) = 0, \sigma_p(N_{21}) = 0$ is that they are integrals of odd functions. We therefore get that, as a Ψ DO of order $-1, \sigma_p(N) = \text{diag}(\sigma_p(N_{22}), \frac{-2}{n-1}\sigma_p(N_{11}))$. Note that the principal symbols of N_{11}, N_{22} are the same as in the case of “ordinary” geodesics, i.e., they do not depend on Y . Explicit formulas for $\sigma_p(N_{11}), \sigma_p(N_{22})$ can be found in [36,37] and they are based on the analysis of the Euclidean case in [34]. Similarly, the principal symbols of \mathbf{d} and δ are independent of Y and are given by

$$\sigma_p(\mathbf{d})[v, \phi] = \text{diag}\left(\frac{1}{2}(\xi_i v_j + \xi_j v_i), \xi_i \phi\right), \quad \sigma_p(\delta)[h, \beta] = (\xi^j h_{ij}, \xi^j \beta_j). \tag{4.6}$$

It is well known [37,38], and follows immediately from Proposition 4.1, that N_{22} is elliptic on solenoidal tensors, i.e., $\sigma_p(N_{22})[h, \beta] = 0$ and $\sigma_p(\delta)[h, \beta] = 0$ imply $[h, \beta] = 0$. Similarly, N_{11} is elliptic on solenoidal (divergence free) 1-forms. As a result, N is elliptic on solenoidal pairs.

We want to emphasize here that pairs solenoidal in M and extended as zero to $M_1 \setminus M$ (that we always assume), may fail to be solenoidal in M_1 due to possible jumps at ∂M .

4.3. Parametrix of N

We proceed as in [37,38]. We will define the Hilbert space $\tilde{H}^2(M_1)$ as in [37,38]. Let $x = (x', x^n)$ be local coordinates in a neighborhood U of a point on ∂M such that $x^n = 0$ defines ∂M . Then we set

$$\|f\|_{\tilde{H}^1(U)}^2 = \int_U \left(\sum_{j=1}^{n-1} |\partial_{x^j} f|^2 + |x^n \partial_{x^n} f|^2 + |f|^2 \right) d \text{Vol}(x).$$

This can be extended to a small enough neighborhood V of ∂M contained in M_1 . Then we set

$$\|f\|_{\tilde{H}^2(M_1)} = \sum_{j=1}^n \|\partial_{x^j} f\|_{\tilde{H}^1(V)} + \|f\|_{\tilde{H}^1(M_1)}. \tag{4.7}$$

We also define the $\tilde{H}^2(M_1)$ space of symmetric 2-tensors and 1-forms, and also the $\tilde{\mathbf{H}}^2(M_1)$ space of pairs $\mathbf{f} = [h, \beta]$ which we denote by $\tilde{\mathbf{H}}^2(M_1)$. Clearly, the latter is a Hilbert space and $\mathbf{H}^2(M_1) \subset \tilde{\mathbf{H}}^2(M_1) \subset \mathbf{H}^1(M_1)$.

The space $\tilde{\mathbf{H}}^2(M_1)$ has the property that for each $\mathbf{f} \in \mathbf{H}^1(M)$ (extended as zero outside M), we have $N\mathbf{f} \in \tilde{\mathbf{H}}^2(M_1)$. This is not true if we replace $\tilde{\mathbf{H}}^2(M_1)$ by $\mathbf{H}^2(M_1)$.

Lemma 4.2. *For any $t = 1, 2, \dots, \infty$, there exists $k > 0$ and a bounded linear operator*

$$Q : \tilde{\mathbf{H}}^2(M_1) \rightarrow \mathcal{S}\mathbf{L}^2(M) \tag{4.8}$$

such that

$$QN\mathbf{f} = \mathbf{f}^s + K\mathbf{f} \quad \text{for all } \mathbf{f} \in \mathbf{H}^1(M), \tag{4.9}$$

where $K : \mathbf{H}^1(M) \rightarrow \mathcal{S}\mathbf{H}^{1+t}(M)$ extends to $K : \mathbf{L}^2(M) \rightarrow \mathcal{S}\mathbf{H}^1(M)$. If $t = \infty$, then $k = \infty$. Moreover, Q can be constructed so that K depends continuously on g in a small neighborhood of a fixed $g_0 \in C^k(M)$.

Proof. We will first reconstruct the solenoidal projection $\mathbf{f}_{M_1}^s$ from $N\mathbf{f}$ modulo smooth terms. Recall that we extend \mathbf{f} as zero outside M , and $\mathbf{f}_{M_1}^s$ is the solenoidal projection of the so-extended \mathbf{f} in M_1 . Next, we will reconstruct \mathbf{f}^s from $\mathbf{f}_{M_1}^s$. We follow [37,38].

As in [37,38], we will work with Ψ DOs with symbols of finite smoothness $k \gg 1$. All operations we are going to perform would require finitely many derivatives of the amplitude and finitely many seminorm estimates. In turn, this would be achieved if $g \in C^k, Y \in C^k, k \gg 1$, and the corresponding Ψ DOs will depend continuously on g, Y .

Since N is elliptic on solenoidal pairs, we get that

$$W := N + N_0\mathcal{P}_{M_1} : C^\infty(M) \rightarrow \mathbf{C}^l(M_1)$$

is an elliptic Ψ DO of order -1 in M^{int} , where N_0 is any properly supported parametrix of $(-\Delta)^{1/2}$ (having principal symbol $|\xi|_g^{-1}$). Next, $W\mathbf{f}$, restricted to a small neighborhood of ∂M_1 in M_1 , is smooth because $\text{supp } \mathbf{f} \subset M$. Therefore, there exists a left parametrix P to W such that $PW - \text{Id} : \mathbf{L}^2(M) \rightarrow \mathbf{H}^t(M_1)$ with $t \gg 1$, if $k \gg 1$. Then $\mathcal{S}_{M_1}(PW - \text{Id})\mathcal{S}_{M_1}$ has the same property, therefore $P_1 := \mathcal{S}_{M_1}P$ satisfies $P_1N = \mathcal{S}_{M_1} + K_2$, where K_2 has the smoothing properties above, therefore,

$$P_1N\mathbf{f} = \mathbf{f}_{M_1}^s + K_2\mathbf{f}. \tag{4.10}$$

The next step is to compare \mathbf{f}^s and $\mathbf{f}_{M_1}^s$. We have $\mathbf{f}^s = \mathbf{f}_{M_1}^s + \mathbf{d}\mathbf{u}$, where $\mathbf{u} = \mathbf{w}_{M_1} - \mathbf{w}$, and the latter are the potentials related to \mathbf{f} in M_1 and M , respectively. Notice that \mathbf{d} commutes with the extension as zero when applied to \mathbf{w} because the latter vanishes on ∂M . Then \mathbf{u} solves the boundary value problem

$$(\delta\mathbf{d})\mathbf{u} = 0 \quad \text{in } M, \quad \mathbf{u}|_{\partial M} = \mathbf{w}_{M_1}|_{\partial M}. \tag{4.11}$$

We need to express $\mathbf{w}_{M_1}|_{\partial M}$ in terms of $N\mathbf{f}$. Since $\mathbf{f} = 0$ outside M , relation (4.10) implies that

$$-\mathbf{d}\mathbf{w}_{M_1} = P_1N\mathbf{f} - K_2\mathbf{f} \quad \text{in } M_1 \setminus M. \tag{4.12}$$

Let $(x, \xi) \in S(M_1^{\text{int}} \setminus M^{\text{int}})$ be such that the magnetic geodesic $\gamma_{x,\xi}(t), t \geq 0$, can be extended in $M_1^{\text{int}} \setminus M^{\text{int}}$ for $t \in [0, \ell_1(x, \xi))$ and $\gamma_{x,\xi}(\ell_1(x, \xi)) \in \partial M_1$; and moreover, $\gamma_{x,\xi}$ is transversal to ∂M_1 . Such (x, ξ) clearly exist if M_1 is close enough to M and if $\gamma_{x,\xi}$ is close to the outgoing magnetic geodesic normal to ∂M . Using (3.21), integrate (4.12) along such $\gamma_{x,\xi}$ to get

$$\langle \mathbf{w}_{M_1}(x), \xi \rangle = \int_0^{\ell_1(x,\xi)} \langle (P_1N\mathbf{f} - K_2\mathbf{f})(\gamma_{x,\xi}), \dot{\gamma}_{x,\xi} \rangle dt. \tag{4.13}$$

Here we denote $\langle \mathbf{w}, \xi \rangle = v_j \xi^j + \varphi$, where $\mathbf{w} = [v(x), \varphi(x)]$ with a 1-form v and a function φ , see also (3.21). Similarly we define $\langle \mathbf{f}, \xi \rangle$. By Lemma 3.3, for a fixed x , one can reconstruct $v(x), \varphi$ from $\langle \mathbf{w}, \xi \rangle$ known for finitely many ξ 's in any neighborhood of a fixed ξ , and this is done by inverting a matrix. Moreover, one can do this near any fixed x , and the norm of the solution operator is uniformly bounded in x , see the remark after Lemma 3.3. By a compactness argument, one can construct an operator P_2 such that

$$\mathbf{w}_{M_1}(x)|_{\partial M} = P_2(P_1N - K_2)\mathbf{f}. \tag{4.14}$$

Arguing as in [37], we see that

$$P_2 P_1 : \tilde{\mathbf{H}}^2(M_1) \rightarrow \mathbf{H}^{1/2}(\partial M) \tag{4.15}$$

is bounded. Let $R : \mathbf{H}^{l-1/2}(M) \rightarrow \mathbf{H}^l(M)$ be the solution operator $\mathbf{u} = R\mathbf{h}$ of the boundary value problem $\delta\mathbf{d}\mathbf{u} = 0$ in M , $\mathbf{u} = \mathbf{h}$ on ∂M . Then (4.10), (4.11), (4.14), and (4.15) imply (see also [37])

$$\mathbf{f}^s = (\text{Id} + dRP_2)P_1 N\mathbf{f} + K\mathbf{f},$$

where K has the required smoothing properties. We apply S to the identity above and set $Q := S(\text{Id} + dRP_2)P_1$.

To prove the last statement of the lemma, we note that all Ψ DOs we work with depend continuously on g, α if $k \gg 1$. The same applies to S, R and P_2 . \square

4.4. Main results for C^k coefficients

Theorem 4.3. *Let (g, α) be a simple C^k magnetic system on M extended to a simple magnetic system on M_1 . Then, for $k \gg 1$,*

- (a) $\text{Ker } I \cap S\mathbf{L}^2(M)$ is finitely dimensional and included in $\mathbf{C}^l(M)$, where $l \rightarrow \infty$ as $k \rightarrow \infty$.
- (b) Assume that I is s -injective for the pair (g, α) . Then

$$\|\mathbf{f}^s\|_{\mathbf{L}^2(M)} \leq C \|N\mathbf{f}\|_{\tilde{\mathbf{H}}^2(M_1)} \tag{4.16}$$

with a constant $C > 0$ that can be chosen to be uniform in (g, α) under a small enough C^k perturbation.

Proof. This theorem is an analog of [37, Theorem 2] except for the uniformity statement which is similar to an analogous result in [38]. Part (a) follows directly from Lemma 4.2. Part (b), without the last statement, can be deduced from Lemma 4.2 as well, as in [37]. Finally, the proof of the statement about the uniformity of C is identical to that of [38, Theorem 2]. \square

Without the assumption that I is s -injective, one has a hypoelliptic a priori estimate that can be obtained from (4.16) by adding $C_s \|f\|_{\mathbf{H}^{-s}(M_1)}$, $\forall s > 0$, to its right-hand side, see [37, Theorem 2(a)].

Corollary 4.4. *The set of simple C^k magnetic systems (g, α) with s -injective magnetic ray transform I is open in the C^k topology, if $k \gg 1$.*

The next lemma is a linear version of Theorem 2.2, see also [38, Lemma 4].

Lemma 4.5. *Let (g, α) be a simple C^k magnetic system on M . Let $\mathbf{f} \in \mathbf{L}^2(M)$ be such that $I\mathbf{f} = 0$. Then there exists a C^l pair \mathbf{w} , with $l \rightarrow \infty$ as $k \rightarrow \infty$, vanishing on ∂M , such that for $\tilde{\mathbf{f}} := \mathbf{f} - d\mathbf{w}$ we have*

$$\partial^m \tilde{\mathbf{f}}|_{\partial M} = 0, \quad |m| \leq l, \tag{4.17}$$

and if $\tilde{\mathbf{f}} = [\tilde{h}, \tilde{\beta}]$, then in semigeodesic boundary normal coordinates,

$$\tilde{h}_{in} = \tilde{\beta}_n = 0 \quad \text{for all } i. \tag{4.18}$$

If $k = \infty$, then $l = \infty$ and, in particular, (4.17) holds for all m .

Proof. In view of Theorem 4.3(a), we may assume without loss of generality that $\mathbf{f} \in \mathbf{C}^l(M)$.

We now show how to construct $\tilde{\mathbf{f}}$ satisfying (4.18). We refer to [13,35] for similar arguments. Let (x', x^n) be semigeodesic boundary normal coordinates. Then $g_{in} = \delta_{in}$, $\Gamma_{nn}^i = \Gamma_{in}^n = 0$ for all i . Let $\mathbf{f} = [h, \beta]$, $\mathbf{w} = [v, \varphi]$. Condition (4.18) is equivalent to

$$\nabla_i v_n + \nabla_n v_i = 2h_{in}, \quad \partial_n \varphi - Y_n^l v_l = \beta_n. \tag{4.19}$$

The first system can be solved by setting first $i = n$ and solving $\partial_n v_n = h_{nn}$, $v_n = 0$ for $x^n = 0$ by integration. Then we solve the remaining system of $n - 1$ ODEs of the form

$$\partial_n v_l - 2\Gamma_{in}^k v_k = 2h_{in} - \partial_l v_n, \quad l = 1, \dots, n - 1, \tag{4.20}$$

with initial conditions $v_l = 0$ for $x^n = 0$. Finally, we solve the second equation in (4.19) with the same zero initial condition. This defines \mathbf{w} near ∂M . To define \mathbf{w} in the whole M , we replace \mathbf{w} by $\chi \mathbf{w}$, where χ is an appropriate smooth cut-off function equal to 1 near ∂M and supported in a larger neighborhood of ∂M .

To prove (4.17), we argue as in Theorem 2.2, following [20]. Assume that (4.17) is not true at some $x_0 \in \partial M$. Then by studying the Taylor expansion of $\langle \tilde{\mathbf{f}}, \xi \rangle = \tilde{h}_{\iota k} \xi^\iota \xi^k + \tilde{\beta}_l \xi^l$ near $x = x_0$, we see that there is $\xi_0 \in T(\partial M)$ such that $\langle \tilde{\mathbf{f}}, \xi \rangle$ is either strictly negative or strictly positive for ξ lying in some neighborhood of ξ_0 . This contradicts the fact that $I\tilde{\mathbf{f}} = 0$ if we integrate over magnetic geodesics originating from x_0 with directions close to ξ_0 . \square

4.5. Analytic magnetic systems

Assume that M is an analytic manifold with smooth boundary ∂M that does not need to be analytic (that can always be achieved by choosing an analytic atlas; and in case we have a simple system, we can start with a fixed global coordinate system and do only analytic changes of variables). We will show that then I is s -injective, if (g, α) are analytic. By “analytic,” we mean real analytic, and we say that f is analytic in the set X , not necessarily open, if f is analytic in a neighborhood of X . Then we write $f \in \mathcal{A}(X)$. Notice that one can construct an analytic M_1 as before.

The central result of this section is the following.

Theorem 4.6. *If (g, α) is an analytic magnetic system on M , then I is s -injective.*

We will give a proof at the end of this section.

Theorem 4.7. *Let g, α be analytic in M_1 . Then N_{kl} , $k, l = 1, 2$, are analytic Ψ DOs in M_1^{int} with principal symbols as in Proposition 4.1.*

Proof. Notice first that by (3.30) and the simplicity assumption, N_{kl} have Schwartz kernels that are analytic away from the diagonal. Therefore, it is enough to prove the theorem for $N_{1,2}$ restricted to an arbitrary small open subset of M_1 . This, however, follows from (4.4) and Lemma B.3 of Appendix B. \square

Lemma 4.8. *Let g, α be analytic in M_1 and assume that $I\mathbf{f} = 0, \mathbf{f} \in \mathbf{L}^2(M)$. Then $\mathbf{f}^s \in \mathcal{A}(M)$.*

Proof. Consider the solenoidal projection $\mathbf{f}_{M_1}^s = \mathbf{f} - \mathbf{d}\mathbf{w}_{M_1}$ of \mathbf{f} (extended as 0 outside M) on M_1 . Since $\delta\mathbf{f}_{M_1}^s = 0$ and $N\mathbf{f}_{M_1}^s = 0$ in M_1^{int} , and since δ and N together form an elliptic system of analytic Ψ DOs (we can apply an elliptic Ψ DO of order 2 to the left of N to make the new operator and δ of the same order, see also [37,38]), we get that $\mathbf{f}_{M_1}^s \in \mathcal{A}(M_1^{\text{int}})$. On the other hand, \mathbf{w}_{M_1} solves $\delta\mathbf{d}\mathbf{w}_{M_1} = 0$ in $M_1^{\text{int}} \setminus M, \mathbf{w}_{M_1} = 0$ on ∂M_1 , and by elliptic boundary regularity, see [38, Lemma 3] and references therein, we have that \mathbf{w}_{M_1} is analytic up to ∂M_1 , therefore $\mathbf{f}_{M_1}^s \in \mathcal{A}(M_1)$.

Next, we have $\mathbf{f}^s = \mathbf{f}_{M_1}^s + \mathbf{d}\mathbf{u}$, where $\mathbf{u} = \mathbf{w}_{M_1} - \mathbf{w}$, see also the proof of Lemma 4.2. Then \mathbf{u} solves the boundary value problem (4.11), and all we need is to prove that $\mathbf{w}_{M_1}|_{\partial M}$ is analytic. Note that in general, \mathbf{w}_{M_1} is not analytic across ∂M , because $\mathbf{d}\mathbf{w}_{M_1}$ may have a jump there but it belongs to H^1 because \mathbf{d} is elliptic; therefore the trace on ∂M is well defined. Since $\mathbf{d}\mathbf{w}_{M_1} = \mathbf{f}_{M_1}^s$ in $M_1^{\text{int}} \setminus M$, and it is analytic in its closure, in the notation of (4.13), we have

$$\langle \mathbf{w}_{M_1}(x), \xi \rangle = \int_0^{\ell_1(x, \xi)} \langle \mathbf{f}_{M_1}^s(\gamma_{x, \xi}), \dot{\gamma}_{x, \xi} \rangle dt$$

for $x \in \partial M$ and ξ in a small neighborhood of the unit exterior normal to ∂M . This, combined with Lemma 3.3 and the remark after it, shows that $\mathbf{w}_{M_1}|_{\partial M}$ is analytic. \square

Lemma 4.9. *Under the assumptions of Lemma 4.5, assume that g and α are analytic in a neighborhood of ∂M . Then $\tilde{\mathbf{f}} = 0$ in a neighborhood of ∂M .*

Proof. It is enough to notice that near $\partial M, \mathbf{w}$ in the proof of Lemma 4.5 is obtained by solving ODEs with analytic coefficients. Therefore, $\tilde{\mathbf{f}}$ is analytic in a neighborhood of ∂M , and it must vanish there by (4.17). \square

Proof of Theorem 4.6. We first find \mathbf{w} as in Lemmas 4.5 and 4.9 such that $\tilde{\mathbf{f}} = \mathbf{f}^s - \mathbf{d}\mathbf{w}$ vanishes near ∂M , i.e.,

$$\mathbf{f}^s = \mathbf{d}\mathbf{w} \quad \text{near } \partial M. \tag{4.21}$$

Our next goal is to show that one can extend $\mathbf{w} = [v, \varphi]$ to the whole M analytically; then (4.21) would be preserved by analytic continuation (recall that \mathbf{f}^s is analytic by Lemma 4.8), and this would imply $\mathbf{f}^s = 0$.

Let $u_{\pm}(x, \xi)$ be the solution of the kinetic equation $\mathbf{G}_{\mu}u = \langle \mathbf{f}^s, \xi \rangle$ with $u = 0$ on $\partial_{\pm}SM$. Integrate to get

$$u_{\pm}(x, \xi) = \mp \int_0^{\ell^{\pm}(x, \xi)} \langle \mathbf{f}^s(\gamma_{x, \xi}), \dot{\gamma}_{x, \xi} \rangle dt,$$

where ℓ^+ stands for ℓ . Since $I\mathbf{f} = 0$, we have $u_+ + u_- = 0$. For (x, ξ) such that x is close enough to ∂M and ξ close enough to $-e_n$ (in boundary normal coordinates), we have $u_+(x, \xi) =$

$\langle \mathbf{dw}, \xi \rangle = \langle v(x), \xi \rangle + \varphi(x)$ by (4.21) and (3.21). This also implies $-u_-(x, \xi) = \langle v(x), \xi \rangle + \varphi(x)$. If we replace ξ by $-\xi$, since $-\xi$ is close to e_n , we get by (4.21), $-u_-(x, -\xi) = -\langle v(x), \xi \rangle + \varphi(x)$; therefore, $u_+(x, -\xi) = -\langle v(x), \xi \rangle + \varphi(x)$.

In particular, for x close enough to ∂M and $\xi \in S_x M$ close enough to either e_n or $-e_n$,

$$\varphi(x) = u_{+, \text{even}}(x, \xi), \quad v_j(x) = \partial_{\eta^j} |\eta|_g u_{+, \text{odd}}(x, \eta/|\eta|_g) \Big|_{|\eta|_g=1}, \tag{4.22}$$

where $u_{+, \text{odd/even}}$ stands for the odd/even part of u_+ w.r.t. ξ . The derivative on the right-hand side of (4.22) can be written as $Pu_{+, \text{odd}}$, where P is a first order differential operator on $S_x M$ with coefficients analytically depending on x and ξ . A direct differentiation shows that $(Pu)_j = \partial_{\xi^j} u - \xi_j \xi^i \partial_{\xi^i} u + \xi_j u$, and the first order part is clearly tangent to $S_x M$.

We define $\varphi(x, \xi)$ and $v(x, \xi)$ by (4.22) on the whole SM . Since $u_+(x, \xi)$ is analytic on SM , so are $\varphi(x, \xi)$ and $v(x, \xi)$. By (4.22), in some open set they are independent of ξ , i.e., $\text{grad}_\xi \varphi(x, \xi) = 0$, $\text{grad}_\xi v(x, \xi) = 0$ there, where grad_ξ stands for the gradient on $S_x M$. Those equalities extend to the whole SM by analytic continuation; therefore, (4.22) defines ξ -independent φ and v in the whole M , and they are analytic functions of x . \square

We are ready now to state the generic s-injectivity result.

Definition 4.10. For a fixed manifold M , we define \mathcal{G}^k to be the set of simple C^k pairs (g, α) with s-injective magnetic ray transform $I = I_{g, \alpha}$.

Theorem 4.11. *There exists $k_0 > 0$, such that for $k \geq k_0$, the set \mathcal{G}^k is open and dense in the set of all simple C^k pairs (g, α) and contains all real analytic simple pairs.*

Proof. By Corollary 4.4, \mathcal{G}^k is open. By Theorem 4.6, it is dense. \square

5. Energy estimates method

It is easy to see that if (M, g, α) is a simple smooth magnetic system and the magnetic ray transform of a smooth function $\phi : SM \rightarrow \mathbb{R}$ vanishes, then ϕ is the flow derivative of a unique function u which is continuous on SM , smooth on $SM \setminus S(\partial M)$, and vanishes on $\partial(SM)$:

$$\mathbf{G}_\mu u = \phi, \quad u|_{\partial(SM)} = 0. \tag{5.1}$$

Indeed, u is defined by the formula

$$u(x, \xi) = - \int_0^{\ell(x, \xi)} \phi(\psi^t(x, \xi)) dt, \quad (x, \xi) \in SM. \tag{5.2}$$

In this section we will analyze the linear problem (5.1) for $\phi(x, \xi)$ of degree at most 1 in ξ (the linear problem for 1-tensors) and for $\phi(x, \xi)$ of degree at most 2 in ξ (the linear problem for 2-tensors).

5.1. Semibasic tensor fields

We recall the notion of semibasic tensor field and its derivatives (we prefer to adhere to the notations of [12]).

Let $\pi : TM \setminus \{0\} \rightarrow M$ be the natural projection, and let $\beta_s^r M := \pi^* \tau_s^r M$ denote the bundle of semibasic tensors of degree (r, s) , where $\tau_s^r M$ is the bundle of tensors of degree (r, s) over M . Sections of the bundles $\beta_s^r M$ are called semibasic tensor fields, and we denote the space of smooth sections by $C^\infty(\beta_s^r M)$ (in particular, $C^\infty(\beta_0^0 M) = C^\infty(TM \setminus \{0\})$). For such a field T , the coordinate representation

$$T = (T_{j_1 \dots j_s}^{i_1 \dots i_r})(x, \xi)$$

holds in the domain of a standard local coordinate system (x^i, ξ^i) on $TM \setminus \{0\}$ associated with a local coordinate system (x^i) in M . Under a change of a local coordinate system, the components of a semibasic tensor field are transformed by the same formula as those of an ordinary tensor field on M .

Every “ordinary” tensor field on M defines a semibasic tensor field by the rule $T \mapsto T \circ \pi$, so that the space of tensor fields on M can be treated as embedded in the space of semibasic tensor fields.

For a semibasic tensor field $(T_{j_1 \dots j_s}^{i_1 \dots i_r})(x, \xi)$, the horizontal derivative is defined by

$$\begin{aligned} T_{j_1 \dots j_s | k}^{i_1 \dots i_r} &= \frac{\partial}{\partial x^k} T_{j_1 \dots j_s}^{i_1 \dots i_r} - \Gamma_{kq}^p \xi^q \frac{\partial}{\partial \xi^p} T_{j_1 \dots j_s}^{i_1 \dots i_r} \\ &+ \sum_{m=1}^r \Gamma_{kp}^m T_{j_1 \dots j_s}^{i_1 \dots i_{m-1} p i_{m+1} \dots i_r} - \sum_{m=1}^r \Gamma_{km}^p T_{j_1 \dots j_{m-1} p j_{m+1} \dots j_s}^{i_1 \dots i_r}, \end{aligned}$$

the vertical derivative by

$$T_{j_1 \dots j_s \cdot k}^{i_1 \dots i_r} = \frac{\partial}{\partial \xi^k} T_{j_1 \dots j_s}^{i_1 \dots i_r},$$

and the modified horizontal derivative by

$$T_{j_1 \dots j_s \cdot k}^{i_1 \dots i_r} = T_{j_1 \dots j_s | k}^{i_1 \dots i_r} + |\xi| Y_k^j T_{j_1 \dots j_s \cdot j}^{i_1 \dots i_r}.$$

The operators

$$\nabla_{|} : C^\infty(\beta_s^r M) \rightarrow C^\infty(\beta_{s+1}^r M), \quad \nabla_{\cdot} : C^\infty(\beta_s^r M) \rightarrow C^\infty(\beta_{s+1}^r M),$$

and

$$\nabla_{\cdot} : C^\infty(\beta_s^r M) \rightarrow C^\infty(\beta_{s+1}^r M),$$

are defined as

$$\begin{aligned}
 (\nabla_{|} T)_{j_1 \dots j_s k}^{i_1 \dots i_r} &= \nabla_{|k} T_{j_1 \dots j_s}^{i_1 \dots i_r} := T_{j_1 \dots j_s |k}^{i_1 \dots i_r}, \\
 (\nabla \cdot T)_{j_1 \dots j_s k}^{i_1 \dots i_r} &= \nabla \cdot_k T_{j_1 \dots j_s}^{i_1 \dots i_r} = T_{j_1 \dots j_s \cdot k}^{i_1 \dots i_r}
 \end{aligned}$$

and

$$(\nabla : T)_{j_1 \dots j_s k}^{i_1 \dots i_r} = \nabla :_k T_{j_1 \dots j_s}^{i_1 \dots i_r} := T_{j_1 \dots j_s :k}^{i_1 \dots i_r}.$$

For convenience, we also define $\nabla^{|}$, ∇^{\cdot} , and $\nabla^{:}$ as

$$\nabla^{|i} = g^{ij} \nabla_{|j}, \quad \nabla^{\cdot i} = g^{ij} \nabla \cdot_j, \quad \nabla^{:i} = g^{ij} \nabla :_j.$$

In [30,34], the operators $\nabla_{|}$ and $\nabla \cdot$ were denoted by $\overset{h}{\nabla}$ and $\overset{v}{\nabla}$ respectively.

Given $u \in C^\infty(TM \setminus \{0\})$, we define

$$\mathbf{X}u(x, \xi) = \xi^i u_{:i} = \xi^i (u_{|i} + |\xi| Y_i^j u \cdot_j).$$

Note that \mathbf{X} restricted to SM coincides with \mathbf{G}_μ .

For $V = (V^i) \in C^\infty(\beta_0^1 M)$, we set

$$\overset{h}{\text{div}} V := V_{|i}^i, \quad \overset{m}{\text{div}} V := V_{:i}^i, \quad \overset{v}{\text{div}} V := V_{\cdot i}^i, \quad (\mathbf{X}V)^i = \xi^k V_{:k}^i.$$

Note if γ is a magnetic geodesic, then

$$(\mathbf{X}V)(\gamma(t), \dot{\gamma}(t)) = \nabla_{\dot{\gamma}} V(\gamma(t), \dot{\gamma}(t)).$$

Given a function $u : SM \rightarrow \mathbb{R}$, we will also denote by u its extension to a positively homogeneous function of degree 0 on $TM \setminus \{0\}$ (hoping that this will not yield any confusion). For a smooth ϕ the smoothness properties of u defined by (5.2) are determined by those of $\ell(x, \xi)$. As mentioned in the beginning of Section 2.2, the latter function is smooth on $SM \setminus S(\partial M)$. All points of $S(\partial M)$ are singular for ℓ as some derivatives of ℓ are unbounded in a neighborhood of such a point. Nonetheless, some derivatives are bounded. Let ρ be a smooth function on M such that $\partial M = \rho^{-1}(0)$ and $|\text{grad } \rho| = 1$ in some neighborhood of ∂M . Define

$$\nabla^{:} u(x, \xi) = \nabla \cdot u(x, \xi) - \langle \nabla \cdot u(x, \xi), \text{grad } \rho(x) \rangle \text{grad } \rho(x).$$

Note that $\nabla^{:} u(x, \xi)$ is completely determined by the restriction of u to the level set of ρ that contains (x, ξ) .

Lemma 5.1. (Cf. [33, Lemma 3.2.3].) *The semibasic vector fields $\nabla^{:} \ell$ and $\nabla \cdot \ell$ are bounded on $SM \setminus S(\partial M)$.*

Proof. Clearly, $\nabla^{:} \rho = 0$. Let $h(x, \xi, t) = \rho(\gamma_{x,\xi}(t))$. Since $h(x, \xi, 0) = \rho(x)$, it follows that $\nabla^{:} h(x, \xi, 0) = 0$ and therefore

$$\nabla^{:} h(x, \xi, t) = a(x, \xi, t)t$$

for some smooth field a . Differentiating the equality $h(x, \xi, \ell(x, \xi)) = 0$ for $(x, \xi) \in SM \setminus S(\partial M)$, we obtain

$$\begin{aligned} \nabla_{\rho}^{\cdot} h(x, \xi, \ell(x, \xi)) + \frac{\partial \rho}{\partial t}(x, \xi, \ell(x, \xi)) \nabla_{\rho}^{\cdot} \ell(x, \xi) \\ = \nabla_{\rho}^{\cdot} h(x, \xi, \ell(x, \xi)) + \langle \text{grad } \rho(y), \eta \rangle \nabla_{\rho}^{\cdot} \ell(x, \xi) = 0, \end{aligned}$$

where $(y, \eta) = \psi^{\ell(x, \xi)}(x, \xi) \in \partial_{-} SM$. Hence,

$$\nabla_{\rho}^{\cdot} \ell(x, \xi) = -\frac{\ell(x, \xi)}{\langle \text{grad } \rho(y), \eta \rangle} a(x, \xi, \ell(x, \xi)).$$

Since $\ell(x, \xi) \leq |\mathbb{L}(y, \eta)|$, boundedness of $\nabla_{\rho}^{\cdot} \ell(x, \xi)$ now follows from Lemma 2.3.

Boundedness of $\nabla^{\cdot} \ell$ is established similarly. \square

5.2. Identities

The next identities are particular cases of those in [12, Lemmas 4.6 and 4.7]. For $(x, \xi) \in SM$:

$$\begin{aligned} 2\langle \nabla^{\cdot} u, \nabla^{\cdot}(\mathbf{X}u) \rangle &= |\nabla^{\cdot} u|^2 + \mathbf{X}(\langle \nabla^{\cdot} u, \nabla^{\cdot} u \rangle) - \text{div}^m(\mathbf{X}u \nabla^{\cdot} u) \\ &\quad + \text{div}^v(\mathbf{X}u \nabla^{\cdot} u) - \langle \tilde{\mathbf{R}}_{\xi}(\nabla^{\cdot} u), \nabla^{\cdot} u \rangle + \langle Y(\nabla^{\cdot} u), \nabla^{\cdot} u \rangle, \end{aligned} \tag{5.3}$$

$$2\langle \nabla^{\cdot} u, \nabla^{\cdot}(\mathbf{X}u) \rangle = |\nabla^{\cdot} u|^2 + |\nabla^{\cdot}(\mathbf{X}u)|^2 + \langle Y(\xi), \nabla^{\cdot} u \rangle^2 - |\mathbf{X}(\nabla^{\cdot} u)|^2, \tag{5.4}$$

where $\tilde{\mathbf{R}}_{\xi}(Z) = (\tilde{R}_{jkl}^i \xi^j \xi^l Z^k)$,

$$\tilde{R}_{jkl}^i = R_{jkl}^i + \xi_j(Y_{k|l}^i - Y_{l|k}^i) + g_{js}(Y_l^s Y_k^i - Y_k^s Y_l^i),$$

and (R_{jkl}^i) is the Riemann curvature tensor.

For a unit speed magnetic geodesic γ , the operator \mathcal{C} on smooth vector fields along γ is defined by (see [12, (45)])

$$\mathcal{C}(Z) = R(\dot{\gamma}, Z)\dot{\gamma} - Y(\dot{Z}) - (\nabla_Z Y)(\dot{\gamma}),$$

where \dot{Z} stands for the covariant derivative along γ , $\dot{Z} = \nabla_{\dot{\gamma}} Z$.

The operator $\tilde{\mathcal{C}}$ on smooth semibasic vector fields on $TM \setminus \{0\}$ is defined by (see [12, Section 4.6])

$$\tilde{\mathcal{C}}(Z)(x, \xi) = \mathbf{R}_{\xi}(Z) - Y(\mathbf{X}Z) - (\nabla_{|Z} Y)(\xi),$$

where $\mathbf{R}_{\xi}(Z) = (R_{jkl}^i \xi^j \xi^l Z^k)$ is the curvature operator.

Notice that

$$\tilde{\mathcal{C}}(Z)(\gamma(t), \dot{\gamma}(t)) = \mathcal{C}(Z(\gamma(t), \dot{\gamma}(t))).$$

In view of this identity, we henceforth omit the tilde in the notation of $\tilde{\mathcal{C}}$, hoping that no confusion will arise.

Lemma 5.2. *The following hold for $(x, \xi) \in SM$:*

$$\begin{aligned} & |\mathbf{X}(\nabla \cdot u)|^2 - \langle \mathcal{C}(\nabla \cdot u), \nabla \cdot u \rangle - \langle Y(\xi), \nabla \cdot u \rangle^2 \\ &= |\nabla \cdot (\mathbf{X}u)|^2 - \operatorname{div}^v(\mathbf{X}u \nabla \cdot u) \\ &\quad - \operatorname{div}^h(\langle \nabla \cdot u, \nabla \cdot u \rangle \xi - (\mathbf{X}u) \nabla \cdot u) + \operatorname{div}^v(\langle \nabla \cdot u, \nabla \cdot u \rangle Y(\xi)), \end{aligned} \tag{5.5}$$

$$\begin{aligned} & |\mathbf{X}(\nabla \cdot u)|^2 - 2\langle \mathcal{C}(\nabla \cdot u), \nabla \cdot u \rangle - \langle Y(\xi), \nabla \cdot u \rangle^2 \\ &= |\nabla \cdot (\mathbf{X}u)|^2 - 2\operatorname{div}^v(\mathbf{X}u \nabla \cdot u) - |\nabla \cdot u|^2 + 2\langle \nabla \cdot u, \nabla \cdot (\mathbf{X}u) \rangle \\ &\quad - 2\operatorname{div}^h(\langle \nabla \cdot u, \nabla \cdot u \rangle \xi - (\mathbf{X}u) \nabla \cdot u) + \operatorname{div}^v(\langle \nabla \cdot u, \nabla \cdot u \rangle Y(\xi)). \end{aligned} \tag{5.6}$$

Proof. Note that

$$\begin{aligned} \mathbf{X}(\langle \nabla \cdot u, \nabla \cdot u \rangle)(x, \xi) &= \xi^j \langle \nabla \cdot u, \nabla \cdot u \rangle_{:j} = \langle \nabla \cdot u, \nabla \cdot u \rangle \xi^j_{:j} - \langle \nabla \cdot u, \nabla \cdot u \rangle \xi^j_{:j} \\ &= \operatorname{div}^m(\langle \nabla \cdot u, \nabla \cdot u \rangle \xi) - \langle \nabla \cdot u, \nabla \cdot u \rangle Y^j_j = \operatorname{div}^m(\langle \nabla \cdot u, \nabla \cdot u \rangle \xi), \end{aligned} \tag{5.7}$$

because $Y^j_j = 0$ by the skew-symmetry of Y .

Using (5.7), we first change (5.3) to

$$\begin{aligned} 2\langle \nabla \cdot u, \nabla \cdot (\mathbf{X}u) \rangle &= |\nabla \cdot u|^2 + \operatorname{div}^m V + \operatorname{div}^v(\mathbf{X}u \nabla \cdot u) \\ &\quad - \langle \tilde{\mathbf{R}}_\xi(\nabla \cdot u), \nabla \cdot u \rangle + \langle Y(\nabla \cdot u), \nabla \cdot u \rangle, \end{aligned} \tag{5.8}$$

with

$$V = \langle \nabla \cdot u, \nabla \cdot u \rangle \xi - (\mathbf{X}u) \nabla \cdot u.$$

Next,

$$\begin{aligned} \operatorname{div}^m V &= V^i_{:i} = V^i_i + Y^j_i V^i_{:j} = \operatorname{div}^h V + \operatorname{div}^v(Y(V)) \\ &= \operatorname{div}^h V + \operatorname{div}^v(\langle \nabla \cdot u, \nabla \cdot u \rangle Y(\xi)) - \operatorname{div}^v(\mathbf{X}u Y(\nabla \cdot u)) \end{aligned}$$

and

$$\begin{aligned} \operatorname{div}^v(\mathbf{X}u Y(\nabla \cdot u)) &= (\mathbf{X}u) Y^j_i g^{ik} u_{:k}{}_{:j} = (\mathbf{X}u)_{:j} Y^j_i g^{ik} u_{:k} + (\mathbf{X}u) g^{ki} Y^j_i u_{:k}{}_{:j} \\ &= \langle \nabla \cdot (\mathbf{X}u), Y(\nabla \cdot u) \rangle, \end{aligned}$$

because $g^{ki} Y^j_i u_{:k}{}_{:j} = 0$ by the skew-symmetry of Y and symmetry of mixed derivatives of u .

Then (5.8) takes the form

$$2\langle \nabla^i u, \nabla^i (\mathbf{X}u) \rangle = |\nabla^i u|^2 + \operatorname{div}^h V + \operatorname{div}^v (\langle \nabla^i u, \nabla^i u \rangle Y(\xi)) - \langle \nabla^i (\mathbf{X}u), Y(\nabla^i u) \rangle + \operatorname{div}^v (\mathbf{X}u \nabla^i u) - \langle \tilde{\mathbf{R}}_\xi(\nabla^i u), \nabla^i u \rangle + \langle Y(\nabla^i u), \nabla^i u \rangle. \tag{5.9}$$

We have (see [12, Section 4.6]):

$$\langle \mathcal{C}(\nabla^i u), \nabla^i u \rangle = \langle \mathbf{R}_\xi(\nabla^i u), \nabla^i u \rangle + \langle \nabla^i (\mathbf{X}u), Y(\nabla^i u) \rangle - \langle \nabla^i u, Y(\nabla^i u) \rangle + \langle Y(\xi), \nabla^i u \rangle^2 - \langle (\nabla_{|\nabla^i u})Y(\xi), \nabla^i u \rangle.$$

Also,

$$\begin{aligned} \langle \tilde{\mathbf{R}}_\xi(\nabla^i u), \nabla^i u \rangle &= (R^i_{jkl} + \xi_j(Y^i_{kl} - Y^i_{lk}) + g_{js}(Y^s_l Y^i_k - Y^s_k Y^i_l)) u^k v^l v^j u_{.i} \\ &= \langle R(\nabla^i u, \xi)\xi, \nabla^i u \rangle + \langle \nabla_\xi Y(\nabla^i u), \nabla^i u \rangle - \langle \nabla_{(\nabla^i u)} Y(\xi), \nabla^i u \rangle \\ &\quad + \langle Y(\xi), \xi \rangle \langle Y(\nabla^i u), \nabla^i u \rangle - \langle Y(\nabla^i u), \xi \rangle \langle Y(\xi), \nabla^i u \rangle \\ &= \langle \mathbf{R}_\xi(\nabla^i u), \nabla^i u \rangle - \langle (\nabla_{(\nabla^i u)} Y)(\xi), \nabla^i u \rangle + \langle Y(\xi), \nabla^i u \rangle^2. \end{aligned}$$

Therefore,

$$\langle \tilde{\mathbf{R}}_\xi(\nabla^i u), \nabla^i u \rangle = \langle \mathcal{C}(\nabla^i u), \nabla^i u \rangle - \langle \nabla^i (\mathbf{X}u), Y(\nabla^i u) \rangle + \langle \nabla^i u, Y(\nabla^i u) \rangle. \tag{5.10}$$

Using (5.10) in (5.9), we get

$$2\langle \nabla^i u, \nabla^i (\mathbf{X}u) \rangle = |\nabla^i u|^2 + \operatorname{div}^h V + \operatorname{div}^v (\langle \nabla^i u, \nabla^i u \rangle Y(\xi)) + \operatorname{div}^v (\mathbf{X}u \nabla^i u) - \langle \mathcal{C}(\nabla^i u), \nabla^i u \rangle. \tag{5.11}$$

Subtracting (5.11) from (5.4), we arrive at (5.5). Multiplying (5.11) by 2 and subtracting the result from (5.4), we arrive at (5.6). \square

5.3. Linear problem for 1-tensors

Theorem 5.3. *Let (M, g, α) be a simple magnetic system, v a square integrable 1-form, and φ a square integrable function on M . If the magnetic ray transform of the function $\phi(x, \xi) = v_i(x)\xi^i + \varphi(x)$ vanishes, then $\varphi = 0$ and $v = dh$ for some function $h \in H^1_0(M)$.*

Proof. Recall that the space $\mathcal{L}^2(M)$ consists of the pairs $[v, \varphi]$, where v is a square integrable 1-form and φ is a square integrable function on M , furnished with the norm (3.16) (see Section 3.3). In $\mathcal{L}^2(M)$, we consider the subspace $\mathcal{S}\mathcal{L}^2(M)$ of solenoidal pairs $[v, \varphi]$, defined by the condition $\delta v = 0$, and the subspace $\mathcal{P}\mathcal{L}^2(M)$ of potential pairs $[dh, 0]$, where $h \in H^1_0(M)$. Then $\mathcal{L}^2(M)$ decomposes as the orthogonal direct sum of $\mathcal{S}\mathcal{L}^2(M)$ and $\mathcal{P}\mathcal{L}^2(M)$.

Associating each pair $[v, \varphi] \in \mathcal{L}^2(M)$ with the function $\phi(x, \xi) = v_i(x)\xi^i + \varphi(x)$, define

$$\mathcal{I}[v, \varphi] = I\phi.$$

Clearly, \mathcal{I} vanishes on $\mathcal{P}\mathcal{L}^2(M)$ and, to prove the theorem, it suffices to show that $\ker \mathcal{I} \cap \mathcal{S}\mathcal{L}^2(M) = 0$.

By considerations similar to those in Section 4, one can prove, as in part (a) of Theorem 4.3, that $\text{Ker } \mathcal{I} \cap \mathcal{S}\mathcal{L}^2(M) \subset \mathcal{C}^l(M)$.

Thus, we may assume under assumptions of the theorem that $[v, \varphi] \in \mathcal{C}^l(M)$.

Define $u : SM \rightarrow \mathbb{R}$ by means of (5.2). Then u satisfies the boundary value problem (5.1). As before, we preserve the notation u for the extension of u to a positively homogeneous function of degree 0 on $TM \setminus \{0\}$. Then for $(x, \xi) \in TM \setminus \{0\}$ we have

$$\mathbf{X}u(x, \xi) = v_i(x)\xi^i + |\xi|\varphi(x).$$

Now, we wish to integrate identity (5.5) over SM . However, u has singularities on $T(\partial M)$, and we need some precautions against them. We proceed in the same way as in the proof of Theorem 3.4.3 in [33]. Let $\rho : M \rightarrow \mathbb{R}$ be a nonnegative smooth function such that $\partial M = \rho^{-1}(0)$ and $|\text{grad } \rho| = 1$ near ∂M . For $\varepsilon > 0$, let $M_\varepsilon = \{x \in M : \rho(x) \geq \varepsilon\}$. Then u is smooth on SM_ε . Integrating (5.5) over SM_ε , we get

$$\begin{aligned} & \int_{SM_\varepsilon} \{ |\mathbf{X}(\nabla \cdot u)|^2 - \langle \mathcal{C}(\nabla \cdot u), \nabla \cdot u \rangle - \langle Y(\xi), \nabla \cdot u \rangle^2 \} d\Sigma^{2n-1}(x, \xi) \\ &= \int_{SM_\varepsilon} |\nabla \cdot (\mathbf{X}u)|^2 d\Sigma^{2n-1} - \int_{SM_\varepsilon} \overset{v}{\text{div}}((\mathbf{X}u)\nabla \cdot u) d\Sigma^{2n-1} \\ & \quad - \int_{SM_\varepsilon} \overset{h}{\text{div}}(\langle \nabla \cdot u, \nabla \cdot u \rangle \xi - (\mathbf{X}u)\nabla \cdot u) d\Sigma^{2n-1} \\ & \quad + \int_{SM_\varepsilon} \overset{v}{\text{div}}(\langle \nabla \cdot u, \nabla \cdot u \rangle Y(\xi)) d\Sigma^{2n-1}(x, \xi). \end{aligned} \tag{5.12}$$

Now, we transform the last three integrals on the right-hand side by using the Gauss–Ostrogradskiĭ formulas of [33, Theorem 2.7.1]. They give:

$$\begin{aligned} \int_{SM_\varepsilon} \overset{v}{\text{div}}((\mathbf{X}u)\nabla \cdot u) d\Sigma^{2n-1} &= n \int_{SM_\varepsilon} \langle (\mathbf{X}u)\nabla \cdot u, \xi \rangle d\Sigma^{2n-1}(x, \xi) \\ &= n \int_{SM_\varepsilon} (\mathbf{X}u)^2 d\Sigma^{2n-1} \end{aligned}$$

because $\langle \nabla \cdot u, \xi \rangle = \mathbf{X}u$,

$$\begin{aligned} & \int_{SM_\varepsilon}^h \operatorname{div}(\langle \nabla \cdot u, \nabla \cdot u \rangle \xi - (\mathbf{X}u) \nabla \cdot u) d\Sigma^{2n-1} \\ &= (-1)^{n-1} \int_{\partial(SM_\varepsilon)} \langle \langle \nabla \cdot u, \nabla \cdot u \rangle \xi - (\mathbf{X}u) \nabla \cdot u, \operatorname{grad} \rho(x) \rangle d\Sigma^{2n-2}(x, \xi) \\ &= (-1)^{n-1} \int_{\partial(SM_\varepsilon)} (\langle \nabla_\rho \cdot u, \nabla \cdot u \rangle \langle \xi, \operatorname{grad} \rho \rangle - \langle \nabla_\rho \cdot u, \xi \rangle \langle \nabla \cdot u, \operatorname{grad} \rho \rangle) d\Sigma^{2n-2} \end{aligned}$$

because a straightforward calculation gives

$$\langle \langle \nabla \cdot u, \nabla \cdot u \rangle \xi - (\mathbf{X}u) \nabla \cdot u, \operatorname{grad} \rho(x) \rangle = \langle \nabla_\rho \cdot u, \nabla \cdot u \rangle \langle \xi, \operatorname{grad} \rho \rangle - \langle \nabla_\rho \cdot u, \xi \rangle \langle \nabla \cdot u, \operatorname{grad} \rho \rangle,$$

and

$$\begin{aligned} & \int_{SM_\varepsilon}^v \operatorname{div}(\langle \nabla \cdot u, \nabla \cdot u \rangle Y(\xi)) d\Sigma^{2n-1}(x, \xi) \\ &= (n-1) \int_{SM_\varepsilon} \langle \langle \nabla \cdot u, \nabla \cdot u \rangle Y(\xi), \xi \rangle d\Sigma^{2n-1}(x, \xi) = 0 \end{aligned}$$

because $\langle Y(\xi), \xi \rangle = 0$.

Hence, (5.12) takes the form

$$\begin{aligned} & \int_{SM_\varepsilon} \{ |\mathbf{X}(\nabla \cdot u)|^2 - \langle \mathcal{L}(\nabla \cdot u), \nabla \cdot u \rangle - \langle Y(\xi), \nabla \cdot u \rangle^2 \} d\Sigma^{2n-1}(x, \xi) \\ &= \int_{SM_\varepsilon} |\nabla \cdot (\mathbf{X}u)|^2 d\Sigma^{2n-1} - n \int_{SM_\varepsilon} (\mathbf{X}u)^2 d\Sigma^{2n-1} \\ &+ (-1)^n \int_{\partial(SM_\varepsilon)} (\langle \nabla_\rho \cdot u, \nabla \cdot u \rangle \langle \xi, \operatorname{grad} \rho \rangle - \langle \nabla_\rho \cdot u, \xi \rangle \langle \nabla \cdot u, \operatorname{grad} \rho \rangle) d\Sigma^{2n-2}. \quad (5.13) \end{aligned}$$

Observe that by Lemma 5.1 and (5.2), the derivatives $\nabla_\rho \cdot u$ and $\nabla \cdot u$ are bounded in $SM \setminus S(\partial M)$. This fact allows us to pass to the limit in (5.13), arriving at the identity

$$\begin{aligned} & \int_{SM} \{ |\mathbf{X}(\nabla \cdot u)|^2 - \langle \mathcal{L}(\nabla \cdot u), \nabla \cdot u \rangle - \langle Y(\xi), \nabla \cdot u \rangle^2 \} d\Sigma^{2n-1}(x, \xi) \\ &= \int_{SM} |\nabla \cdot (\mathbf{X}u)|^2 d\Sigma^{2n-1} - n \int_{SM} (\mathbf{X}u)^2 d\Sigma^{2n-1} \\ &+ (-1)^n \int_{\partial(SM)} (\langle \nabla_\rho \cdot u, \nabla \cdot u \rangle \langle \xi, \operatorname{grad} \rho \rangle - \langle \nabla_\rho \cdot u, \xi \rangle \langle \nabla \cdot u, \operatorname{grad} \rho \rangle) d\Sigma^{2n-2}. \end{aligned}$$

Note that the last integral on the right-hand side vanishes, because $u|_{\partial(SM)} = 0$ and hence $\nabla u|_{\partial(SM)} = 0$.

For $(x, \xi) \in SM$, $\mathbf{X}u(x, \xi) = v_i(x)\xi^i + \varphi(x)$. As follows from Lemma 4.5.3 of [34] (and is easy to check directly),

$$\int_{SM} |\nabla(v_i(x)\xi^i + \varphi(x))|^2 d\Sigma^{2n-1} \leq n \int_{SM} (v_i(x)\xi^i + \varphi(x))^2 d\Sigma^{2n-1}.$$

Setting $Z = \nabla u$, we thus have

$$\int_{SM} \{|\mathbf{X}Z|^2 - \langle C(Z), Z \rangle - \langle Y(\xi), Z \rangle^2\} d\Sigma^{2n-1} \leq 0.$$

By Santaló’s formula (A.4) and the Index Lemma A.10 of Appendix A, this is possible if and only if $Z = 0$. So $u(x, \xi)$ is independent of ξ , which clearly implies the sought conclusion. \square

5.4. Linear problem for 2-tensors

For a magnetic system (M, g, α) and $(x, \xi) \in SM$, put

$$k_\mu(x, \xi) = \sup_\eta \{2K(x, \sigma_{\xi, \eta}) + \langle Y(\eta), \xi \rangle^2 + (n + 3)|Y(\eta)|^2 - 2\langle (\nabla_\eta Y)(\xi), \eta \rangle\},$$

where the supremum is taken over all unit vectors $\eta \in T_x M$ orthogonal to ξ , and $K(x, \sigma_{\xi, \eta})$ is the sectional curvature of the 2-plane $\sigma_{\xi, \eta}$ spanned by ξ and η .

Define

$$k_\mu^+(x, \xi) = \max\{0, k(x, \xi)\}$$

and

$$k(M, g, \alpha) = \sup_\gamma \int_0^{T_\gamma} k_\mu^+(\gamma(t), \dot{\gamma}(t)) dt,$$

where the supremum is taken over all unit speed magnetic geodesics $\gamma : [0, T_\gamma] \rightarrow M$ running between boundary points.

Theorem 5.4. *If (M, g, α) be a simple magnetic system with $k(M, g, \alpha) \leq 4$, then I is s -injective.*

Remark 5.5. Note that the condition $k(M, g, \alpha) \leq 4$ holds if (M, g) is negatively curved and the C^1 -norm of Y is small enough. Also, it is easy to see that this conditions is valid for every sufficiently small simple piece of any magnetic system.

Proof. Assume that $[h, \beta] \in \ker I \cap \mathbf{SL}^2(M)$. Then, by Theorem 4.3(a), the pair $[h, \beta]$ is smooth. Define $u : SM \rightarrow \mathbb{R}$ by means of (5.2). Then u satisfies the boundary value problem (5.1). As

before, we preserve the notation u for the extension of u to a positively homogeneous function of degree 0 on $TM \setminus \{0\}$. Then for $(x, \xi) \in TM \setminus \{0\}$ we have

$$\mathbf{X}u = |\xi|^{-1} h_{ij}(x) \xi^i \xi^j + \beta_j(x) \xi^j. \quad (5.14)$$

Now, we wish to integrate identity (5.6) over SM . Again u has singularities on $T(\partial M)$. We overcome this obstacle in the same way as in the proof of the previous theorem. Thus, we deduce

$$\begin{aligned} & \int_{SM} \{ |\mathbf{X}(\nabla \cdot u)|^2 - 2\langle \mathcal{C}(\nabla \cdot u), \nabla \cdot u \rangle - \langle Y(\xi), \nabla \cdot u \rangle^2 \} d\Sigma^{2n-1}(x, \xi) \\ &= \int_{SM} |\nabla \cdot (\mathbf{X}u)|^2 d\Sigma^{2n-1} - 2 \int_{SM} \overset{v}{\operatorname{div}}(\langle \mathbf{X}u, \nabla \cdot u \rangle) d\Sigma^{2n-1} - \int_{SM} |\nabla \cdot u|^2 d\Sigma^{2n-1} \\ &+ 2 \int_{SM} \langle \nabla \cdot u, \nabla \cdot (\mathbf{X}u) \rangle d\Sigma^{2n-1} - 2 \int_{SM} \overset{h}{\operatorname{div}}(\langle \nabla \cdot u, \nabla \cdot u \rangle \xi - \langle \mathbf{X}u, \nabla \cdot u \rangle) d\Sigma^{2n-1} \\ &+ \int_{SM} \overset{v}{\operatorname{div}}(\langle \nabla \cdot u, \nabla \cdot u \rangle Y(\xi)) d\Sigma^{2n-1}. \end{aligned}$$

Again we use the Gauss–Ostrogradskiĭ formulas to transform integrals of divergent form. Denoting the leftmost side of the above formula by A , we obtain

$$\begin{aligned} A &= \int_{SM} |\nabla \cdot (\mathbf{X}u)|^2 d\Sigma^{2n-1} - \int_{SM} \{ 2n(\mathbf{X}u)^2 + |\nabla \cdot u|^2 \} d\Sigma^{2n-1} \\ &+ 2 \int_{SM} \langle \nabla \cdot u, \nabla \cdot (\mathbf{X}u) \rangle d\Sigma^{2n-1} \\ &\leq \int_{SM} |\nabla \cdot (\mathbf{X}u)|^2 d\Sigma^{2n-1} - (2n+1) \int_{SM} (\mathbf{X}u)^2 d\Sigma^{2n-1} \\ &+ 2 \int_{SM} \langle \nabla \cdot u, \nabla \cdot (\mathbf{X}u) \rangle d\Sigma^{2n-1}, \end{aligned} \quad (5.15)$$

where we have used the inequality $|\mathbf{X}u| = |\langle \nabla \cdot u, \xi \rangle| \leq |\nabla \cdot u|$.

From (5.14) we have for $(x, \xi) \in SM$

$$(Xu)_{,l} = 2h_{lj} \xi^j - \xi_l h_{ij} \xi^i \xi^j + \beta_l.$$

Therefore,

$$\int_{SM} |\nabla \cdot (\mathbf{X}u)|^2 d\Sigma^{2n-1} = \int_{SM} |2h(\xi) - \langle h, \xi^2 \rangle \xi + \beta|^2 d\Sigma^{2n-1}$$

$$\begin{aligned}
 &= \int_{SM} \{4|h(\xi)|^2 - 3\langle h, \xi^2 \rangle^2 + |\beta|^2 + 4\langle h(\xi), \beta \rangle - \langle h, \xi^2 \rangle \langle \beta, \xi \rangle\} d\Sigma^{2n-1} \\
 &= \int_{SM} \{4|h(\xi)|^2 - 3\langle h, \xi^2 \rangle^2 + |\beta|^2\} d\Sigma^{2n-1}, \tag{5.16}
 \end{aligned}$$

$$\int_{SM} (\mathbf{X}u)^2 d\Sigma^{2n-1} = \int_{SM} \{\langle h, \xi^2 \rangle^2 + \langle \beta, \xi \rangle^2\} d\Sigma^{2n-1}. \tag{5.17}$$

Now, we transform the last integral in (5.15). Once

$$\nabla \cdot u = \nabla_{|} u - Y(\nabla \cdot u),$$

it follows that

$$\langle \nabla \cdot u, \nabla \cdot (\mathbf{X}u) \rangle = \langle \nabla_{|} u, \nabla \cdot (\mathbf{X}u) \rangle - \langle Y(\nabla \cdot u), \nabla \cdot (\mathbf{X}u) \rangle. \tag{5.18}$$

Observe that

$$\begin{aligned}
 \langle \nabla_{|} u, \nabla \cdot (\mathbf{X}u) \rangle &= g^{kl} u_{|k} (2h_{lj} \xi^j - \xi_l h_{ij} \xi^i \xi^j + \beta_l) \\
 &= 2(u g^{kl} h_{lj} \xi^j)_{|k} - 2u g^{kl} h_{lj,k} \xi^j - (u \xi^k h_{ij} \xi^i \xi^j)_{|k} \\
 &\quad + u \xi^k (h_{ij} \xi^i \xi^j)_{|k} + (u \beta^k)_{|k} - u \beta^k_{,k} \\
 &= \overset{h}{\text{div}}(W) - 2u(\delta h)_j \xi^j + u \mathbf{G}(h_{ij} \xi^i \xi^j) - u \delta(\beta) \\
 &= \overset{h}{\text{div}}(W) - (n-1)\langle uY(\beta), \xi \rangle + u \mathbf{G}(\langle h, \xi^2 \rangle), \tag{5.19}
 \end{aligned}$$

with

$$W^k = 2u g^{kl} h_{lj} \xi^j - u \xi^k h_{ij} \xi^i \xi^j + u \beta^k.$$

Here we have used Eqs. (3.23) for solenoidal pairs.

Integrating (5.19) and transforming integrals by means of Gauss–Ostrogradskii formulas, we get

$$\begin{aligned}
 &\int_{SM} \langle \nabla_{|} u, \nabla \cdot (\mathbf{X}u) \rangle d\Sigma^{2n-1} \\
 &= -(n-1) \int_{SM} \langle uY(\beta), \xi \rangle d\Sigma^{2n-1} + \int_{SM} u \mathbf{G}(\langle h, \xi^2 \rangle) d\Sigma^{2n-1} \\
 &= - \int_{SM} \overset{v}{\text{div}}(uY(\beta)) d\Sigma^{2n-1} - \int_{SM} \mathbf{G}(u)\langle h, \xi^2 \rangle d\Sigma^{2n-1} \\
 &= - \int_{SM} \langle Y(\beta), \nabla \cdot u \rangle d\Sigma^{2n-1} - \int_{SM} \mathbf{G}_\mu(u)\langle h, \xi^2 \rangle d\Sigma^{2n-1}
 \end{aligned}$$

$$\begin{aligned}
& + \int_{SM} (\xi^i Y_i^j u_{\cdot j}) \langle h, \xi^2 \rangle d\Sigma^{2n-1} \\
= & \int_{SM} \langle Y(\nabla \cdot u), \beta \rangle d\Sigma^{2n-1} - \int_{SM} \langle h, \xi^2 \rangle^2 d\Sigma^{2n-1} \\
& + \int_{SM} \langle \nabla \cdot u, Y(\xi) \rangle \langle h, \xi^2 \rangle d\Sigma^{2n-1}. \tag{5.20}
\end{aligned}$$

Next,

$$\begin{aligned}
& \int_{SM} \langle Y(\nabla \cdot u), \nabla \cdot (\mathbf{X}u) \rangle d\Sigma^{2n-1} \\
= & \int_{SM} \langle Y(\nabla \cdot u), (2h_{ij} \xi^j - \xi_l h_{ij} \xi^i \xi^j + \beta_l) \rangle d\Sigma^{2n-1} \\
= & 2 \int_{SM} \langle Y(\nabla \cdot u), h(\xi) \rangle d\Sigma^{2n-1} + \int_{SM} \langle \nabla \cdot u, Y(\xi) \rangle \langle h, \xi^2 \rangle d\Sigma^{2n-1} \\
& + \int_{SM} \langle Y(\nabla \cdot u), \beta \rangle d\Sigma^{2n-1}, \tag{5.21}
\end{aligned}$$

where $h(\xi) = (g^{ki} h_{ij} \xi^j)$.

From (5.18), (5.20), and (5.21) we get

$$\begin{aligned}
& \int_{SM} \langle \nabla \cdot u, \nabla \cdot (\mathbf{X}u) \rangle d\Sigma^{2n-1} \\
= & - \int_{SM} \langle h, \xi^2 \rangle^2 d\Sigma^{2n-1} - 2 \int_{SM} \langle Y(\nabla \cdot u), h(\xi) \rangle d\Sigma^{2n-1} \\
\leq & - \int_{SM} \langle h, \xi^2 \rangle^2 d\Sigma^{2n-1} + \frac{2}{n+2} \int_{SM} |h(\xi)|^2 d\Sigma^{2n-1} \\
& + \frac{n+2}{2} \int_{SM} |Y(\nabla \cdot u)|^2 d\Sigma^{2n-1}. \tag{5.22}
\end{aligned}$$

Using (5.16), (5.17), and (5.22) in (5.15), we find that

$$\begin{aligned}
A \leq & \int_{SM} \{4|h(\xi)|^2 - 3\langle h, \xi^2 \rangle^2 + |\beta|^2\} d\Sigma^{2n-1} \\
& - (2n+1) \int_{SM} \{\langle h, \xi^2 \rangle^2 + \langle \beta, \xi \rangle^2\} d\Sigma^{2n-1} - 2 \int_{SM} \langle h, \xi^2 \rangle^2 d\Sigma^{2n-1}
\end{aligned}$$

$$\begin{aligned}
 & + \frac{4}{n+2} \int_{SM} |h(\xi)|^2 d\Sigma^{2n-1} + (n+2) \int_{SM} |Y(\nabla \cdot u)|^2 d\Sigma^{2n-1} \\
 = & \frac{4(n+3)}{n+2} \int_{SM} |h(\xi)|^2 d\Sigma^{2n-1} - 2(n+3) \int_{SM} \langle h, \xi^2 \rangle^2 d\Sigma^{2n-1} \\
 & + \int_{SM} |\beta|^2 d\Sigma^{2n-1} - (2n+1) \int_{SM} \langle \beta, \xi \rangle^2 d\Sigma^{2n-1} \\
 & + (n+2) \int_{SM} |Y(\nabla \cdot u)|^2 d\Sigma^{2n-1}.
 \end{aligned}$$

As follows from Lemma 4.5.3 of [34] (and is easy to check directly),

$$\begin{aligned}
 \int_{SM} |h(\xi)|^2 d\Sigma^{2n-1} & \leq \frac{n+2}{2} \int_{SM} \langle h, \xi^2 \rangle^2 d\Sigma^{2n-1}, \\
 \int_{SM} |\beta|^2 d\Sigma^{2n-1} & \leq n \int_{SM} \langle \beta, \xi \rangle^2 d\Sigma^{2n-1}.
 \end{aligned}$$

Therefore,

$$A \leq (n+2) \int_{SM} |Y(\nabla \cdot u)|^2 d\Sigma^{2n-1}.$$

Setting $Z = \nabla \cdot u$, we rewrite this as follows:

$$\int_{SM} \{ |\mathbf{X}(Z)|^2 - 2\langle \mathcal{C}(Z), Z \rangle - \langle Y(Z), \xi \rangle^2 - (n+2)|Y(Z)|^2 \} d\Sigma^{2n-1} \leq 0. \tag{5.23}$$

We need the following lemma whose proof is given below.

Lemma 5.6. *Let $\gamma : [0, T] \rightarrow M$ be a unit speed magnetic geodesic. If*

$$\int_0^T k_\mu^+(\gamma(t), \dot{\gamma}(t)) dt \leq 4/T,$$

then for every smooth vector field Z along γ vanishing at the endpoints of γ and orthogonal to $\dot{\gamma}$ we have

$$\int_0^T \{ |\dot{Z}|^2 - 2\langle \mathcal{C}(Z), Z \rangle - \langle Y(Z), \dot{\gamma} \rangle^2 - (n+2)|Y(Z)|^2 \} dt \geq 0 \tag{5.24}$$

with equality if and only if $Z = 0$.

Continuing the proof of the theorem, observe that, by Santaló’s formula (A.4) and Lemma 5.6, inequality (5.23) may hold if and only if $Z = 0$. So $\nabla \cdot u = 0$, which means that $u(x, \xi)$ is independent of ξ . Since the pair $[h, \beta]$ is solenoidal, this readily yields the sought conclusion. \square

Proof of Lemma 5.6. Denote the left-hand side of (5.24) by B . Since

$$C(Z) = \mathbf{R}_{\dot{\gamma}}(Z) - Y(\dot{Z}) - (\nabla_Z Y)(\dot{\gamma}),$$

we have

$$B = \int_0^T \{ |\dot{Z}|^2 - 2\langle \mathbf{R}_{\dot{\gamma}}(Z), Z \rangle + 2\langle Y(\dot{Z}), Z \rangle + 2\langle (\nabla_Z Y)(\dot{\gamma}), Z \rangle - \langle Y(Z), \dot{\gamma} \rangle^2 - (n + 2)|Y(Z)|^2 \} dt.$$

Let $e_i, i = 1, \dots, n$, be an orthonormal frame at $\gamma(0)$, with $e_n = \dot{\gamma}(0)$. Let $E_i(t), 0 \leq t \leq T$, be a vector field along γ satisfying the equation

$$\dot{E}_i = Y(E_i)$$

with the initial condition $E_i(0) = e_i$. Surely, $E_n = \dot{\gamma}$. Since

$$\frac{d}{dt} \langle E_i, E_j \rangle = \langle \dot{E}_i, E_j \rangle + \langle E_i, \dot{E}_j \rangle = \langle Y(E_i), E_j \rangle + \langle E_i, Y(E_j) \rangle = 0,$$

we see that $E_1(t), \dots, E_n(t)$ is an orthonormal frame for each t .

Consider the expansion of Z in this frame:

$$Z(t) = z^\kappa(t) E_\kappa(t), \quad \kappa \in \{1, \dots, n - 1\}.$$

Then

$$\begin{aligned} \dot{Z} &= \dot{z}^\kappa E_\kappa + z^\kappa \dot{E}_\kappa = \dot{z}^\kappa E_\kappa + z^\kappa Y(E_\kappa) = \dot{z}^\kappa E_\kappa + Y(Z), \\ |\dot{Z}|^2 &= \sum_\kappa \dot{z}^\kappa \dot{z}^\kappa + 2Y_{\kappa\lambda} z^\kappa \dot{z}^\lambda + |Y(Z)|^2, \\ \langle Y(\dot{Z}), Z \rangle &= \langle \dot{z}^\kappa Y(E_\kappa) + Y^2(Z), z^\lambda E_\lambda \rangle = -Y_{\kappa\lambda} z^\kappa \dot{z}^\lambda - |Y(Z)|^2 \end{aligned}$$

with

$$Y_{\kappa\lambda} = \langle Y(E_\kappa), E_\lambda \rangle.$$

Therefore,

$$\begin{aligned}
 B &= \int_0^T \left\{ \sum_{\kappa} \dot{z}^{\kappa} \dot{z}^{\kappa} - [2K(\sigma_{\dot{\gamma}, Z})|Z|^2 + \langle Y(Z), \dot{\gamma} \rangle^2 \right. \\
 &\quad \left. + (n+3)|Y(Z)|^2 - 2\langle (\nabla_Z Y)(\dot{\gamma}), Z \rangle] \right\} dt \\
 &\geq \int_0^T \left\{ \sum_{\kappa} \dot{z}^{\kappa} \dot{z}^{\kappa} - k_{\mu}|Z|^2 \right\} dt = \int_0^T \sum_{\kappa} \{ \dot{z}^{\kappa} \dot{z}^{\kappa} - k_{\mu} z^{\kappa} z^{\kappa} \} dt.
 \end{aligned}$$

Now, the claim follows from Lyapunov’s inequality [16, p. 346]. \square

6. General rigidity theorems

6.1. Rigidity in a given conformal class

We first state a rigidity theorem, fixing the conformal class of a metric. The theorem below generalizes the corresponding well-known result for the ordinary boundary rigidity problem, see [9,23,24].

Theorem 6.1. *Let (g, α) and (g', α') be simple magnetic systems on M whose boundary action functions $\mathbb{A}|_{\partial M \times \partial M}$ and $\mathbb{A}'|_{\partial M \times \partial M}$ coincide. If g' is conformal to g , i.e., $g' = \omega^2(x)g$ for a smooth positive function ω on M , then $\omega \equiv 1$ and $\alpha' = \alpha + dh$ for some smooth function h on M vanishing on ∂M , hence (g', α') is gauge equivalent to (g, α) .*

Proof. In view of Lemma 2.1, $\omega = 1$ on the boundary of M . Next, using Lemma 2.5, we see that the scattering relations \mathcal{S} and \mathcal{S}' of both magnetic systems coincide.

Let us show that $\omega = 1$ on the whole of M . Note that for $(x, \xi) \in \partial_+ SM$

$$\mathbb{A}'(\gamma'_{x,\xi}) \leq \mathbb{A}'(\gamma_{x,\xi}) = \frac{1}{2} \int_0^{\ell(x,\xi)} \omega^2(\gamma_{x,\xi}(t)) dt + \frac{1}{2} \ell(x, \xi) - \int_{\gamma_{x,\xi}} \alpha'. \tag{6.1}$$

Using (2.9) and (2.8), whence we obtain

$$\begin{aligned}
 \text{Vol}_{g'}(M) &= \frac{1}{\omega_{n-1}} \int_{\partial_+ SM} \mathbb{A}'(\gamma'_{x,\xi}) d\mu(x, \xi) \\
 &\leq \frac{1}{2\omega_{n-1}} \int_{\partial_+ SM} \left\{ \int_0^{\ell(x,\xi)} \omega^2(\gamma_{x,\xi}(t)) dt + \ell(x, \xi) \right\} d\mu(x, \xi) \\
 &= \frac{1}{2} \int_M \omega^2 d\text{Vol} + \frac{1}{2} \text{Vol}_g(M).
 \end{aligned}$$

On the one hand, by Hölder's inequality

$$\int_M \omega^2 d \text{Vol} \leq \left\{ \int_M \omega^n d \text{Vol} \right\}^{\frac{2}{n}} \left\{ \int_M d \text{Vol} \right\}^{\frac{n-2}{n}} = \text{Vol}_{g'}(M)^{\frac{2}{n}} \text{Vol}_g(M)^{\frac{n-2}{n}}, \quad (6.2)$$

with equality if and only if $\omega \equiv 1$.

It follows that

$$\text{Vol}_{g'}(M) \leq \frac{1}{2} \text{Vol}_{g'}(M)^{\frac{2}{n}} \text{Vol}_g(M)^{\frac{n-2}{n}} + \frac{1}{2} \text{Vol}_g(M). \quad (6.3)$$

However, by Theorem 2.8, $\text{Vol}_{g'}(M) = \text{Vol}_g(M)$, which implies that (6.3) holds with the equality sign. This means that (6.2) holds with the equality sign. Thus, $\omega \equiv 1$.

Now, (6.1) and the equality $\mathbb{A}'(\gamma'_{x,\xi}) = \mathbb{A}(\gamma_{x,\xi})$ yield

$$\int_{\gamma_{x,\xi}} \alpha \geq \int_{\gamma'_{x,\xi}} \alpha'.$$

In view of (A.4) and (2.8), we conclude now that

$$\int_{\gamma_{x,\xi}} (\alpha' - \alpha) = 0$$

for all $(x, \xi) \in \partial_+ SM$. By Theorem 5.3, we see that $\alpha' - \alpha = dh$ for some function h on M vanishing on ∂M . This completes the proof of the theorem. \square

6.2. Rigidity of analytic systems

Theorem 6.2. *If M is a real-analytic compact manifold with boundary, and (g, α) and (g', α') are simple real-analytic magnetic systems on M with the same boundary action function, then these systems are gauge equivalent.*

Proof. Consider the diffeomorphism $f : M \rightarrow M$ constructed in the proof of Theorem 2.2. Recall that in a neighborhood of ∂M , f coincides with the map $\exp_{\partial M} \circ (\exp'_{\partial M})^{-1}$, which is obviously analytic for analytic systems.

As shown in that proof, g and f^*g' have the same jets on ∂M . Hence, by analyticity, g and f^*g' coincide in some connected neighborhood of ∂M , and f is thus an analytic isometry of a neighborhood of ∂M of the analytic Riemannian manifold (M, g) onto a neighborhood of ∂M of the analytic Riemannian manifold (M, g') , which is moreover the identity when restricted to ∂M . Now, the same arguments as in the proof of [21, Theorem C(a)] show that f extends from a neighborhood of ∂M to an analytic isometry \tilde{f} of (M, g) onto (M, g') . Now, the sought conclusion follows from Theorem 6.1 applied to the magnetic systems (M, g, α) and $(M, \tilde{f}^*g', \tilde{f}^*\alpha')$. \square

6.3. Rigidity of reversible systems

A magnetic system (M, g, α) is said to be reversible if the flip $(x, \xi) \mapsto (x, -\xi)$ conjugates ψ^t with ψ^{-t} . It is easy to see that a system is reversible if and only if $d\alpha = 0$, i.e., if and only if the Lorentz force Y of the system vanishes. In this case magnetic geodesics are nothing but the ordinary geodesics of the Riemannian manifold (M, g) ; moreover, simplicity of a reversible magnetic system (M, g, α) is equivalent to simplicity of the Riemannian manifold (M, g) .

It is interesting to know whether reversibility of a magnetic system can be established by boundary measurements. Observe that if a system is reversible then

$$\mathcal{S}(-\mathcal{S}(x, \xi)) = (x, -\xi) \tag{6.4}$$

for all $(x, \xi) \in \partial_+ SM$. Let us call the systems satisfying (6.4) *boundary reversible*.

Theorem 6.3. *A simple magnetic system is boundary reversible if and only if it is reversible.*

Proof. Let (M, g, α) be a boundary reversible simple magnetic system. Consider the magnetic system $(g, -\alpha)$ on M . Note that if $\gamma : [0, T] \rightarrow M$ is a unit speed magnetic geodesic of the system (g, α) , then the curve $\bar{\gamma} : [0, T] \rightarrow M$, defined as $\bar{\gamma}(t) = \gamma(T - t)$, is a unit speed magnetic geodesic of the system $(g, -\alpha)$.

For $x, y \in M$, let $\gamma_{x,y} : [0, T_{x,y}] \rightarrow M$ denote the unit speed magnetic geodesic of the system (g, α) from x to y . By Lemma A.5,

$$\mathbb{A}(\gamma_{x,y}) \leq \mathbb{A}(\bar{\gamma}_{y,x}),$$

i.e.,

$$T_{x,y} - \int_{\gamma_{x,y}} \alpha \leq T_{y,x} + \int_{\gamma_{y,x}} \alpha. \tag{6.5}$$

Interchanging x and y , we receive

$$T_{y,x} - \int_{\gamma_{y,x}} \alpha \leq T_{x,y} + \int_{\gamma_{x,y}} \alpha. \tag{6.6}$$

From (6.5) and (6.6) we get

$$\int_{\gamma_{x,y}} \alpha + \int_{\gamma_{y,x}} \alpha \geq 0. \tag{6.7}$$

For $(x, \xi) \in \partial_+ SM$ and $y = \gamma_{x,\xi}(\ell(x, \xi))$, we have in view of (6.4)

$$\gamma_{x,y} = \gamma_{x,\xi}, \quad \bar{\gamma}_{y,x} = \bar{\gamma}_{x,\xi},$$

where $\tilde{\gamma}_{x,\xi} : [0, \tilde{\ell}(x, \xi)] \rightarrow M$ is the unit speed magnetic geodesic of the magnetic system $(M, g, -\alpha)$ with initial conditions $\tilde{\gamma}_{x,\xi}(0) = x$, $\dot{\tilde{\gamma}}_{x,\xi}(0) = \xi$, and with $\tilde{\gamma}_{x,\xi}(\tilde{\ell}(x, \xi)) = y$. Therefore, (6.7) yields

$$\int_{\gamma_{x,\xi}} \alpha - \int_{\tilde{\gamma}_{x,\xi}} \alpha \geq 0. \quad (6.8)$$

By Santaló's formula (A.4) and (2.8), we have

$$\int_{\partial_+ SM} \left\{ \int_{\gamma_{x,\xi}} \alpha - \int_{\tilde{\gamma}_{x,\xi}} \alpha \right\} d\mu(x, \xi) = 0. \quad (6.9)$$

Combining (6.9) and (6.8) yields

$$\int_{\gamma_{x,\xi}} \alpha = \int_{\tilde{\gamma}_{x,\xi}} \alpha \quad (6.10)$$

for all $(x, \xi) \in \partial_+ SM$. From (6.5) we then obtain

$$\ell(x, \xi) \leq \tilde{\ell}(x, \xi),$$

and from (6.6),

$$\tilde{\ell}(x, \xi) \leq \ell(x, \xi).$$

Thus,

$$\tilde{\ell}(x, \xi) = \ell(x, \xi). \quad (6.11)$$

Now, by (6.10) and (6.11)

$$\mathbb{A}(\tilde{\gamma}_{x,\xi}) = \mathbb{A}(\gamma_{x,\xi}) = \mathbb{A}(x, y).$$

This implies that $\tilde{\gamma}_{x,\xi}$ is a unit speed magnetic geodesic of the system (g, α) which joins x to y . By simplicity, we then have

$$\tilde{\gamma}_{x,\xi} = \gamma_{x,\xi}.$$

Valid for all $(x, \xi) \in \partial_+ SM$, this equality means clearly that the magnetic system (M, g, α) is reversible. \square

Theorem 6.4. *A simple reversible magnetic system (M, g, α) is magnetic boundary rigid if and only if the simple Riemannian manifold (M, g) is boundary rigid.*

Proof. Suppose that a simple reversible magnetic system (g, α) on M is magnetic boundary rigid. Let g' be a simple Riemannian metric on M whose boundary distance function equals the boundary distance function of g . We must prove that g' is isometric to g by an isometry which is the identity on the boundary.

Without loss of generality we may assume (see, e.g., [20, Theorem 2.1]) that

$$g'|_{\partial M} = g|_{\partial M}. \tag{6.12}$$

Then equality of the boundary distance functions of g and g' and (6.12) are well known to imply equality of the scattering relations of the Riemannian manifolds (M, g) and (M, g') (see, e.g., [22]).

Consider the magnetic system (g', α) on M . Since $d\alpha = 0$, this system is reversible as well. It is trivial that for a reversible system, the scattering relation of the magnetic system coincides with the scattering relation of the underlying Riemannian metric. Therefore, the magnetic systems (g, α) and (g', α) on M have the same scattering relations. In view of (6.12) and Lemma 2.6, the boundary actions functions of these systems coincide. From the magnetic boundary rigidity of (M, g, α) we infer that the metric g' is isometric to g by an isometry which is the identity on the boundary, as required.

Now, let (M, g, α) be a simple reversible magnetic system such that the simple Riemannian manifold (M, g) is boundary rigid. Let (g', α') be a simple magnetic system on M whose boundary action function $\mathbb{A}'|_{\partial M \times \partial M}$ equals the boundary action function $\mathbb{A}|_{\partial M \times \partial M}$ of the magnetic system (g, α) . By Theorem 2.2, we may assume that $g'|_{\partial M} = g|_{\partial M}$ and $\alpha'|_{\partial M} = \alpha|_{\partial M}$. Now, Lemma 2.5 implies that the scattering relations of the magnetic systems (M, g, α) and (M, g', α') are the same. Since (M, g, α) is reversible and hence boundary reversible, (M, g', α') too is boundary reversible and hence, by Theorem 6.3, it is also reversible. This implies readily that the scattering relations of the Riemannian manifolds (M, g) and (M, g') are the same. Applying Lemma 2.6 to the magnetic systems $(M, g, 0)$ and $(M, g', 0)$, we see that the boundary distance functions of the metrics g and g' are the same. Hence, by the boundary rigidity of the Riemannian manifold (M, g) , there is a diffeomorphism $f : M \rightarrow M$, $f|_{\partial M} = \text{identity}$, such that $g' = f^*g$. Then Theorem 6.1, applied to the magnetic systems (M, g, α) and $(M, f^*g', f^*\alpha')$, gives the gauge equivalence of the magnetic systems (g, α) and (g', α') , as required. \square

6.4. Generic local boundary rigidity

We will prove that near each (g_0, α_0) in the generic set \mathcal{G}^k of Definition 4.10, the action on the boundary determines (g, α) . Note that, by Theorem 5.4, this generic set \mathcal{G}^k contains the magnetic systems (g, α) with $k(M, g, \alpha) \leq 4$, in particular magnetic systems in which the underlying Riemannian metric is negatively curved and the magnetic field is sufficiently small.

Theorem 6.5. *Let k_0 be as in Theorem 4.11. There exists $k \geq k_0$ such that for every $(g_0, \alpha_0) \in \mathcal{G}^k$, there is $\varepsilon > 0$ such that for any two magnetic systems (g, α) , (g', α') with*

$$\|g - g_0\|_{C^k(M)} + \|\alpha - \alpha_0\|_{C^k(M)} \leq \varepsilon, \quad \|g' - g_0\|_{C^k(M)} + \|\alpha' - \alpha_0\|_{C^k(M)} \leq \varepsilon$$

we have the following:

$$\mathbb{A}_{g, \alpha} = \mathbb{A}_{g', \alpha'} \quad \text{on } \partial M \times \partial M \tag{6.13}$$

implies that (g, α) and (g', α') are gauge equivalent, i.e., $g' = \psi^*g, \alpha' = \psi^*\alpha + d\phi$ with some $C^{k+1}(M)$ -diffeomorphism $\psi : M \rightarrow M$, fixing the boundary, and some C^{k+1} function ϕ vanishing on ∂M .

Observe that g is solenoidal (with respect to itself), and if we replace α by its solenoidal projection $\alpha^s := \alpha - d\phi$, where $\Delta_g \phi = \delta\alpha, \phi|_{\partial M} = 0$, then $\delta\alpha^s = 0$ as well. Therefore, $[\frac{1}{2}g, -\alpha^s]$ is not necessarily a solenoidal pair in the sense of Definition 3.7, instead

$$\delta \left[\frac{1}{2}g, -\alpha^s \right] = \left[\frac{n-1}{2}Y(\alpha^s), 0 \right]. \tag{6.14}$$

We will prove an analogue of [11, Theorem 2.1].

Lemma 6.6. *Let $(g, \alpha) \in C^{k,\mu}, 0 < \mu < 1, k \geq 2$, be a simple magnetic system on M with $\delta\alpha = 0$. Then for any other (g', α') close enough to (g, α) , there exists a magnetic system $(\tilde{g}', \tilde{\alpha}')$ gauge equivalent to (g', α') and satisfying (6.14), i.e.,*

$$\delta \left[\frac{1}{2}\tilde{g}', -\tilde{\alpha}' \right] = \left[\frac{n-1}{2}Y(\alpha), 0 \right], \tag{6.15}$$

where δ and Y are related to (g, α) .

Moreover, if $\|(g', \alpha') - (g, \alpha)\|_{C^{k,\mu}} \leq \varepsilon$ with $\varepsilon \ll 1$, then $\|(\tilde{g}', \tilde{\alpha}') - (g, \alpha)\|_{C^{k,\mu}} \leq \varepsilon_1$, with $\varepsilon_1 = \varepsilon_1(\varepsilon) \rightarrow 0$ as $\varepsilon \rightarrow 0$.

Proof. Our argument is much the same as that in the proof of [11, Theorem 2.1]. So we merely sketch it.

We have to show that there exist a diffeomorphism f of M , fixing ∂M , and a function φ , vanishing on ∂M , such that $\tilde{g}' = f^*g', \tilde{\alpha}' = f^*\alpha' + d\varphi$, and (6.15) holds. If f is close enough to the identity, we can identify f with a certain vector field v as follows. If v is a vector field vanishing on ∂M and $|\nabla v|_g < 1$, the map

$$e_v(x) = \exp_x(v(x))$$

is well defined on M , with image in M again, and it is a diffeomorphism for $\|v\|_{C^{k,\mu}}$ small enough. Then $v \mapsto e_v$ has an inverse defined as follows. If f is a diffeomorphism close enough to the identity, set $v_f(x) = \exp_x^{-1}(f(x))$, i.e., $v_f(x) = \dot{\gamma}(0)$, where $\gamma : [0, 1] \rightarrow M$ is the geodesic such that $\gamma(0) = x, \gamma(1) = f(x)$. The existence of such a geodesic follows from the strict convexity of the boundary, following in turn from the simplicity assumption. Clearly, the differential of the map $v \mapsto e_v$ at $v = 0$ is the identity transformation.

Let $h = \frac{1}{2}(g' - g), \beta = -(\alpha' - \alpha)$. Condition (6.15) then takes the form

$$\delta(e_v^*(g/2 + h)) + \frac{n-1}{2}Y(e_v^*(\alpha + \beta) + d\varphi) = \frac{n-1}{2}Y(\alpha), \tag{6.16}$$

$$-\delta(e_v^*(\alpha + \beta) + d\varphi) = 0. \tag{6.17}$$

We define the map (recall that (g, α) is fixed)

$$\begin{aligned}
 &F([v, \varphi], [h, \beta]) \\
 &= \left[\delta(e_v^*(g/2 + h)) + \frac{n-1}{2} Y(e_v^*(\alpha + \beta) + d\varphi), -\delta(e_v^*(\alpha + \beta) + d\varphi) \right].
 \end{aligned}$$

Then (6.16), (6.17) can be rewritten as

$$F([v, \varphi], [h, \beta]) = \left[\frac{n-1}{2} Y(\alpha), 0 \right]. \tag{6.18}$$

We want to solve this equation for $[v, \varphi]$ if $[h, \beta]$ are small enough and g, α are fixed. If $[h, \beta] = [0, 0]$, then $[v, \varphi] = [0, 0]$ is a solution by (6.14). To show solvability for $[h, \beta]$ small enough, we apply an implicit function theorem in Banach spaces as in [11]. For this purpose, we need to compute the derivative $F'_{[v, \varphi]}([0, 0], [0, 0])$.

Set $[h, \beta] = 0$ in (6.18) and use (3.20) to deduce

$$\begin{aligned}
 &F'_{[v, \varphi]}([0, 0], [0, 0])[v, \varphi] \\
 &= \left[\delta d^s v + \frac{n-1}{2} Y(Y(v) + d\langle v, \alpha \rangle + d\varphi), -\delta(Y(v) + d\langle v, \alpha \rangle + d\varphi) \right] \\
 &= \delta \mathbf{d}[v, -\langle v, \alpha \rangle - \varphi],
 \end{aligned}$$

where v is considered as a 1-form by lowering the index. Thus, the derivative above is the superposition of the map $[v, \varphi] \mapsto [v, -\langle v, \alpha \rangle - \varphi]$, which is an isomorphism (and this map and its inverse preserve the zero boundary conditions), and $\delta \mathbf{d}$ with Dirichlet boundary conditions, which is also invertible in appropriate spaces. We refer to [11] for more technical details. \square

The next lemma states, loosely speaking, that gauge equivalent pairs differ by a potential one, modulo a quadratic term.

Lemma 6.7. *Let (g', α') and (g, α) be in C^k , $k \geq 2$, and gauge equivalent, i.e.,*

$$g' = \psi^* g, \quad \alpha' = \psi^* \alpha + d\varphi$$

for some diffeomorphism ψ fixing ∂M and some function φ be vanishing on ∂M . Set $\mathbf{f} = [\frac{1}{2}(g' - g), -(\alpha' - \alpha)]$. Then there exists \mathbf{w} , vanishing on ∂M , such that

$$\mathbf{f} = d\mathbf{w} + \mathbf{f}_2,$$

and for (g, α) belonging to any bounded set U in C^k , there exists $C(U) > 0$ such that

$$\|\mathbf{f}_2\|_{C^{k-2}} \leq C(U) \|\psi - \text{Id}\|_{C^{k-1}}^2, \quad \|\mathbf{w}\|_{C^{k-1}} \leq C(U) \|\psi - \text{Id}\|_{C^{k-1}}.$$

Proof. As in the proof above, set $v(x) = \exp_x^{-1}(\psi(x))$, which is a well defined vector field if ψ is close enough to the identity in C^2 (it is enough to prove the claim in this case only) and $v = 0$ on ∂M . Set $\psi_\tau(x) = \exp_x(\tau v(x))$, $0 \leq \tau \leq 1$. Let $g^\tau = \psi_\tau^* g$. Then the Taylor formula implies

$$g' = g + \left. \frac{d}{d\tau} \right|_{\tau=0} g^\tau + h = g + 2d^s v + h,$$

where

$$|h| \leq \frac{1}{2} \max_{\tau \in [0,1]} \left| \frac{d^2 g^\tau}{d\tau^2} \right|,$$

and $2d^s v$ is the linearization of g^τ at $\tau = 0$. To estimate h , write

$$g_{ij}^\tau = g_{kl} \circ \psi_\tau \frac{\partial \psi_\tau^k}{\partial x^i} \frac{\partial \psi_\tau^l}{\partial x^j},$$

and differentiate twice w.r.t. τ . Notice that

$$\left| \frac{\partial^2 \psi_\tau}{\partial \tau^2} \right| \leq C \|v\|_{L^\infty}^2, \quad \left| \frac{\partial^2 \nabla \psi_\tau}{\partial \tau^2} \right| \leq C \|v\|_{C^1}^2.$$

This yields the stated estimate for the first component of \mathbf{f}_2 for $k = 2$. The estimates for $k > 2$ go along similar lines by expressing the remainder h in its Lagrange form and estimating the derivatives of h .

The analysis of the second component is similar, using (3.20). In particular, we get that $\mathbf{w} = [v, -\alpha(v) - \phi]$ which corresponds well with the linearization formula (3.20). \square

Proof of Theorem 6.5. We can assume that (g', α') is replaced by its gauge equivalent pair that satisfies the assumptions of Lemma 6.6 (where δ and Y are related to (g, α)). Set $\mathbf{f} = [(g' - g)/2, -(\alpha' - \alpha)]$. Then $\delta \mathbf{f} = 0$. We assume that (g, α) belongs to a small enough neighborhood of (g_0, α_0) , so that $(g, \alpha) \in \mathcal{G}^k$, and the constant C in Theorem 4.3 is uniform in that neighborhood. By the second statement of Lemma 6.6 and the assumptions of Theorem 6.5,

$$\|\mathbf{f}\|_{C^k} \leq \varepsilon_1(\varepsilon), \tag{6.19}$$

where $\varepsilon_1 \rightarrow 0$ as $\varepsilon \rightarrow 0$, and $k \gg 1$ is fixed (this requires the original (g', α') to be in \mathcal{G}^{k+1} if we want to avoid the $C^{k,\mu}$ spaces). Moreover, $\|g\|_{C^k} + \|\alpha\|_{C^k} \leq A$, where $A > 0$ depends on (g_0, α_0) and on an upper bound ε_0 of ε . All constants C below will depend only on A and will be uniform in $\varepsilon \leq \varepsilon_0$.

Let ψ be a diffeomorphism in M that maps boundary normal coordinates w.r.t. g into boundary normal coordinates w.r.t. g' near ∂M , then extended to the whole M . In other words, if ψ_1 that maps a neighborhood of ∂M into $\partial M \times [0, \delta]$ defines the semigeodesic normal coordinates related to g and if ψ_2 is defined in the same way corresponding to g' , then $\psi = \psi_2 \circ \psi_1^{-1}$. Clearly,

$$\|\psi - \text{Id}\|_{C^{k-1}} \leq C \|g' - g\|_{C^k} \leq C' \|\mathbf{f}\|_{C^k}. \tag{6.20}$$

Set $\tilde{g}' = \psi^* g'$, $\tilde{\alpha}' = \psi^* \alpha' + d\varphi$, where φ is such that in boundary normal coordinates $\tilde{\alpha}'_n = 0$, see Lemma 4.5. Set $\tilde{\mathbf{f}} = [\tilde{g}'/2, -\tilde{\alpha}'] - [g/2, -\alpha]$. By Lemma 6.7 and (6.20),

$$\tilde{\mathbf{f}} = \mathbf{f} + \mathbf{dw} + \mathbf{h}, \quad \|\mathbf{h}\|_{C^{k-2}} \leq C \|\mathbf{f}\|_{C^k}^2, \tag{6.21}$$

with $\mathbf{w}|_{\partial M} = 0$; so, roughly speaking, $\tilde{\mathbf{f}}$ and \mathbf{f} differ, up to a quadratic term, only by a potential term. By Theorem 2.2, \mathbf{f} vanishes on ∂M together with its derivatives up to any fixed order m ,

if $k \gg 1$. We will choose m below. We are going to prove that $\mathbf{f} = 0$, and this would prove the theorem.

Since $\mathbb{A}_{g,\alpha} = \mathbb{A}_{g',\alpha'} = \mathbb{A}_{\tilde{g}',\tilde{\alpha}'}$ on $\partial M \times \partial M$, by the linearization Lemma 3.1

$$\|I\tilde{\mathbf{f}}\|_{L^\infty} \leq C\|\tilde{\mathbf{f}}\|_{C^1}^2.$$

By (3.27), $\|N\tilde{\mathbf{f}}\|_{L^\infty(M_1)} \leq C\|I\tilde{\mathbf{f}}\|_{L^\infty}$; therefore,

$$\|N\tilde{\mathbf{f}}\|_{L^\infty(M_1)} \leq C\|\tilde{\mathbf{f}}\|_{C^1}^2. \tag{6.22}$$

We apply Theorem 4.3 to get

$$\|\tilde{\mathbf{f}}^s\|_{L^2(M)} \leq C\|N\tilde{\mathbf{f}}\|_{H^2(M_1)}, \tag{6.23}$$

where we estimated the \tilde{H}^2 norm by the H^2 norm, which is finite for $\tilde{\mathbf{f}}$. Note that for the solenoidal projection $\tilde{\mathbf{f}}^s$ of $\tilde{\mathbf{f}}$ we have, by (6.21),

$$\tilde{\mathbf{f}}^s = \mathbf{f} + \mathbf{h}^s; \tag{6.24}$$

in particular,

$$\|\tilde{\mathbf{f}}^s\|_{L^2(M)} \geq \|\mathbf{f}\|_{L^2(M)} - \|\mathbf{h}^s\|_{L^2(M)} \geq \|\mathbf{f}\|_{L^2(M)} - C\|\mathbf{f}\|_{C^2}^2. \tag{6.25}$$

Using interpolation estimates in $H^s(M_1)$, see [41, Theorem 4.3.1/1], and (6.22), we get

$$\|N\tilde{\mathbf{f}}\|_{H^2(M_1)} \leq C\|N\tilde{\mathbf{f}}\|_{H^s(M_1)}^{2/s} \|N\tilde{\mathbf{f}}\|_{L^2(M_1)}^{1-2/s} \leq C'\|\tilde{\mathbf{f}}\|_{C^1}^{2-4/s}, \quad s > 2. \tag{6.26}$$

To estimate the first factor in the middle term, we used the fact that $\partial^\alpha \tilde{\mathbf{f}}|_{\partial M} = 0$ for $|\alpha| \leq m$ and therefore, if $m \geq s - 2$,

$$\|N\tilde{\mathbf{f}}\|_{H^s(M_1)} \leq C\|\tilde{\mathbf{f}}\|_{H^{s-1}(M)} \leq C'\|\mathbf{f}\|_{C^{s+1}(M)} \leq C''$$

if $k \geq s + 1$. The last inequality follows by comparing $\mathbf{f}, \tilde{\mathbf{f}}$ using (6.21) and the estimate on \mathbf{w} in Lemma 6.7, combined with (6.20). Combine (6.23), (6.25), and (6.26) to get

$$\|\mathbf{f}\|_{L^2(M)} \leq C(\|\mathbf{f}\|_{C^1}^{2-4/s} + \|\mathbf{f}\|_{C^2}^2).$$

Using again the fact that $\|\mathbf{f}\|_{C^k(M)} \leq C$ ($k \gg 1$), Sobolev embedding estimates, and interpolation estimates in $H^s(M)$, one gets

$$\|\mathbf{f}\|_{L^2(M)} \leq C\|\mathbf{f}\|_{L^2(M)}^{(2-4/s)\mu}$$

with $0 < \mu < 1$ that can be chosen as close to 1 as needed, provided that $k \gg 1$. It is enough now to choose $s > 2$ and μ so that $(2 - 4/s)\mu > 1$ to deduce that $\mathbf{f} = 0$ if $\varepsilon \ll 1$, see (6.19). \square

7. Rigidity of two-dimensional systems

The main result of this section is the following rigidity theorem for two-dimensional systems:

Theorem 7.1. *If $\dim M = 2$ and (g, α) and (g', α') are simple magnetic systems on M with the same boundary action function, then these systems are gauge equivalent.*

This theorem generalizes the boundary rigidity theorem for simple Riemannian surfaces which was established in [32]. Our proof of Theorem 7.1 mimics that of the mentioned theorem in [32].

First of all, by Theorem 2.2 and Lemma 2.5 we may assume that g' coincides with g on ∂M ,

$$g'|_{\partial M} = g|_{\partial M},$$

and that the scattering relations of the magnetic systems (M, g, α) and (M, g', α') coincide:

$$S' = S. \quad (7.1)$$

The crucial step then consists in establishing that the scattering relation of a two-dimensional magnetic system (M, g, α) determines the Dirichlet-to-Neumann (DN) map associated to the Laplace–Beltrami operator of the Riemannian manifold (M, g) .

It is proved in [19,21] for two-dimensional manifolds that the DN map determines the conformal class of the Riemannian metric up to an isometry that is the identity on the boundary.

Afterwards, the proof of Theorem 7.1 is finished by applying Theorem 6.1 which claims magnetic boundary rigidity within a given conformal class.

Derivation of the connection between the scattering relation and the DN map is based on the properties of the magnetic ray transform and the commutation formula between the magnetic flow and the fiberwise Hilbert transform.

We proceed with describing the needed properties of the magnetic ray transform.

7.1. More about the magnetic ray transform

We recall that, given a notation F for a function space (C^k , L^p , H^k , etc.), we denote by $\mathcal{F}(M)$ the corresponding space of pairs $[v, \varphi]$, with v a 1-form and φ a function on M . Also, recall that in the space $\mathcal{L}^2(M)$ we consider the norm (3.16) defined as

$$\|w\|^2 = \int_M (|v|_g^2 + \varphi^2) d \text{Vol}.$$

Associating each pair $[v, \varphi] \in \mathcal{L}^2(M)$ with the function $\phi(x, \xi) = v_i(x)\xi^i + \varphi(x)$, we may consider $\mathcal{L}^2(M)$ as a subspace of $L^2(SM)$.

Consider the restriction \mathcal{I} of the magnetic ray transform I to $\mathcal{L}^2(M)$:

$$\begin{aligned} \mathcal{I}[v, \varphi](x, \xi) &= \int_0^{\ell(x, \xi)} v_i(\gamma_{x, \xi}(t)) \dot{\gamma}_{x, \xi}^i(t) dt + \int_0^{\ell(x, \xi)} \varphi(\gamma_{x, \xi}(t)) dt \\ &= \mathcal{I}_1 v(x, \xi) + \mathcal{I}_0 \varphi(x, \xi), \quad (x, \xi) \in \partial_+ SM. \end{aligned}$$

By Lemma 3.2, we receive an operator

$$\mathcal{I} : \mathcal{L}^2(M) \rightarrow L^2_\mu(\partial_+ SM).$$

Let

$$\mathcal{I}^* : L^2_\mu(\partial_+ SM) \rightarrow \mathcal{L}^2(M)$$

be its dual. The same calculations as in Section 3.4 show that

$$\mathcal{I}^* w = [\mathcal{I}_1^* w, \mathcal{I}_0^* w]$$

with

$$\mathcal{I}_1^* w = \left(\int_{S_x M} \xi^i w^\sharp(x, \xi) d\sigma_x(\xi) \right), \quad \mathcal{I}_0^* w = \int_{S_x M} w^\sharp(x, \xi) d\sigma_x(\xi).$$

The following holds:

$$\ker \mathcal{I} = \mathcal{P}\mathcal{L}^2(M), \tag{7.2}$$

$$\text{Im } \mathcal{I}^* \subset \mathcal{S}\mathcal{L}^2(M). \tag{7.3}$$

Equality (7.2) is just the claim of Theorem 5.3. Equality (7.3) follows, as soon as $\text{Im } \mathcal{I}^*$ is in the orthogonal complement to $\ker \mathcal{I}$ and the orthogonal complement to $\mathcal{P}\mathcal{L}^2(M)$ is $\mathcal{S}\mathcal{L}^2(M)$.

Consider the operator

$$\mathcal{N} = \mathcal{I}^* \mathcal{I} : \mathcal{L}^2(M) \rightarrow \mathcal{L}^2(M),$$

which can be written down as

$$\mathcal{N}[v, \varphi] = [\mathcal{N}_{11}v + \mathcal{N}_{10}\varphi, \mathcal{N}_{01}v + \mathcal{N}_{00}\varphi].$$

Considerations similar to those in the proof of 4.1 show the following:

Proposition 7.2. \mathcal{N} is a Ψ DO in M^{int} of order -1 with principal symbol

$$\begin{aligned} \sigma_p(\mathcal{N}) &= \text{diag}(\sigma_p(\mathcal{N}_{11}), \sigma_p(\mathcal{N}_{00})), \\ \sigma_p(\mathcal{N}_{11})^j_i(x, \xi) &= c_n |\xi|^{-1} (\delta_i^j - \xi^j \xi_i / |\xi|^2), \\ \sigma_p(\mathcal{N}_{00})(x, \xi) &= c_n |\xi|^{-1}, \end{aligned}$$

where $|\xi|^2 = g^{ij}(x)\xi_i \xi_j$ and $\xi^j = g^{ji}(x)\xi_i$.

Define

$$C_\alpha^\infty(\partial_+ SM) = \{w \in C^\infty(\partial_+ SM): w^\sharp \in C^\infty(SM)\},$$

where w^\sharp is the function that is constant along the orbits of the magnetic flow and equals w on $\partial_+ SM$. This space can be described in terms of the scattering relation \mathcal{S} alone; by the forthcoming Lemma 7.6

$$C_\alpha^\infty(\partial_+ SM) = \{w \in C^\infty(\partial_+ SM): Aw \in C^\infty(\partial(SM))\}, \tag{7.4}$$

where

$$Aw(x, \xi) = \begin{cases} w(x, \xi), & (x, \xi) \in \partial_+ SM, \\ w \circ \mathcal{S}^{-1}(x, \xi), & (x, \xi) \in \partial_- SM. \end{cases}$$

The following is an analog of [29, Theorems 3.3.3, 3.3.4] (see also [32, Theorem 1.4], [31, Theorems 4.1, 4.2]).

Theorem 7.3. *Let (M, g, α) be a simple magnetic system. Then, for every pair $[v, \varphi] \in C^\infty(M)$, there exist $w \in C_\alpha^\infty(\partial_+ SM)$ and $f \in C^\infty(M)$ such that*

$$[v + \nabla f, \varphi] = \mathcal{I}^* w.$$

(Note that if v is solenoidal then, in view of (7.3), f is harmonic in M .)

Proof. Our argument is the same as in [29,31,32]. Embed M into a closed manifold \tilde{M} and extend g and α smoothly to a Riemannian metric g and a 1-form α on \tilde{M} .

If $U \subset \tilde{M}$ is an open neighborhood of M with smooth boundary, then (\bar{U}, g, α) is also a simple magnetic system if ∂U is close enough to ∂M . Henceforth such an U is assumed to be fixed.

Denote the magnetic ray transform for (\bar{U}, g, α) by \mathcal{I}_U , and denote by r_M the operator of restriction to M . We have the following analog of [29, Theorems 3.3.1, 3.3.2] (see also [32, Theorem 3.1], [31, Theorem 4.3]):

Lemma 7.4. *For every pair $[v, \varphi] \in \mathcal{H}^{s+1}(M)$, $s \geq 0$, there exists a pair $[u, \psi] \in \mathcal{H}^s(U)$ and $f \in H^{s+2}(U)$ such that*

$$[v + \nabla f, \varphi] = r_M \mathcal{I}_U^* \mathcal{I}_U [u, \psi].$$

Again, the proof is similar to the one in [29,31,32]; we will merely sketch it.

Cover \tilde{M} by finitely many open sets U_k such that $U = U_1$, $U_k \cap M = \emptyset$ for $k \geq 2$, and (\bar{U}_k, g, α) is a simple magnetic system for every k . Let $\{h_k\}$ be a subordinate partition of unity such that $h_1|_M = 1$. Consider the operators $\mathcal{I}_{(k)}, \mathcal{I}_{(k)}^*$ for (\bar{U}_k, g, α) , and define the following operator on the bundle $T\tilde{M} \oplus (\tilde{M} \times \mathbb{R})$:

$$P[u, \psi] = \sum_k h_k \mathcal{I}_{(k)}^* \mathcal{I}_{(k)}([u, \psi]|_{U_k}).$$

In view of Proposition 7.2, P is a Ψ DO of order -1 with principal symbol

$$c_n \operatorname{diag} \left(\frac{\delta_i^j}{|\xi|} - \frac{\xi^j \xi_i}{|\xi|^3}, \frac{1}{|\xi|} \right).$$

The operator $\Lambda : C^\infty(T\tilde{M}) \rightarrow C^\infty(T\tilde{M})$,

$$\Lambda = -c_n \nabla (-\Delta)^{-3/2} \delta,$$

is a Ψ DO of order -1 with principal symbol

$$c_n \frac{\xi^j \xi_i}{|\xi|^3}.$$

Therefore, the principal symbol of the operator

$$P + \operatorname{diag}(\Lambda, 0)$$

equals

$$c_n \operatorname{diag}(\delta_i^j |\xi|^{-1}, |\xi|^{-1}),$$

which means that $P + \operatorname{diag}(\Lambda, 0)$ is an elliptic Ψ DO of order -1 .

Now, the same arguments as in [29,31,32] show that the operator

$$r_M(P + \operatorname{diag}(\Lambda, 0)) : \mathcal{H}^s(U) \rightarrow \mathcal{H}^{s+1}(M)$$

has closed range and finite codimension.

Since on $\mathcal{H}^s(U)$

$$r_M P = r_M \mathcal{I}_U^* \mathcal{I}_U, \tag{7.5}$$

(7.3) and the argument of [29,31,32] show that the equation

$$r_M(P + \operatorname{diag}(\Lambda, 0))[u, \psi] = [v, \varphi]$$

has a solution $[u, \psi] \in \mathcal{H}^s(U)$ for every pair $[v, \varphi] \in \mathcal{H}^{s+1}(M)$. Then by (7.5)

$$[v, \varphi] + [\nabla f, 0] = r_M \mathcal{I}_U^* \mathcal{I}_U [u, \psi]$$

with $f = c_n (-\Delta)^{-3/2} \delta u$, which proves the lemma.

Continuing the proof of Theorem 7.3, observe that by Lemma 7.4, for every pair $[v, \varphi] \in C^\infty(M)$ there exist $[u, \psi] \in C^\infty(U)$ and $f \in C^\infty(M)$ such that

$$[v + \nabla f, \varphi] = r_M \mathcal{I}_U^* \mathcal{I}_U [u, \psi]. \tag{7.6}$$

For $(x, \xi) \in SU$, define $\ell^\pm(x, \xi)$ by

$$\pi(\psi^{\ell^\pm}(x, \xi)) \in \partial U, \quad \ell^- < 0, \quad \ell^+ > 0.$$

Define

$$w^+(x, \xi) = \int_0^{\ell^+(x, \xi)} (u_i(\gamma_{x, \xi}(t)) \dot{\gamma}_{x, \xi}^i(t) + \psi(\gamma_{x, \xi}(t))) dt,$$

$$w^-(x, \xi) = \int_{\ell^-(x, \xi)}^0 (u_i(\gamma_{x, \xi}(t)) \dot{\gamma}_{x, \xi}^i(t) + \psi(\gamma_{x, \xi}(t))) dt.$$

These functions belong to $C^\infty(SU)$. Let $w = (w^+ + w^-)|_{\partial_+ SM}$. It is easy to see that $(w^+ + w^-)$ is constant on the orbits of the magnetic flow; therefore, $w^\sharp = (w^+ + w^-)|_{SM}$ and $w^\sharp \in C^\infty(SM)$. Since $\mathcal{I}_U[u, \psi] = (w^+ + w^-)|_{\partial_+ SU}$, we see from (7.6) that $\mathcal{I}^*w = [v + \nabla f, \varphi]$, which completes the proof of the theorem. \square

7.2. Scattering relation and folds

The main aim of this section is to prove the characterization (7.4) for the space $C^\infty_\alpha(\partial_+ SM)$ (Lemma 7.6). We will proceed in the same way as in [32].

Preserving the notations of the previous section, define the map

$$\Phi : \partial(SM) \rightarrow \partial_- SU$$

by

$$\Phi(x, \xi) = \psi^{\ell^+(x, \xi)}(x, \xi), \quad (x, \xi) \in \partial(SM). \tag{7.7}$$

Since ℓ^+ is smooth in SU , Φ is smooth as well.

Lemma 7.5. (Cf. [32, Theorem 4.1].) Φ is a fold map with fold $S(\partial M)$.

We recall that a smooth map $f : \mathcal{M} \rightarrow \mathcal{N}$ between two smooth manifolds \mathcal{M} and \mathcal{N} of the same dimension is said to be a *Whitney fold* (with fold L) at a point $m \in L$ if f drops rank by one simply at m , so that $\{x : df(x) \text{ is singular}\}$ is a smooth hypersurface near m and $\ker(df(m))$ is transverse to $T_m L$.

Example. Let Σ be a manifold, \mathcal{R} a smooth function on Σ having 0 as a regular value, and $\mathcal{M} = \{\mathcal{R}^{-1}(0)\}$. Let X be a nonzero vector field on Σ such that $X\mathcal{R}(m) = 0$, and $XX\mathcal{R}(m) \neq 0$ for a point $m \in \mathcal{M}$. Let \mathcal{N} be a hypersurface in Σ transversal to X such that the projection $f : \mathcal{M} \rightarrow \mathcal{N}$ along the integral curves of X is well defined. Then f is a Whitney fold at m with fold $\mathcal{M} \cap \{(X\mathcal{R})^{-1}(0)\}$.

Indeed, we may assume without loss of generality that Σ is a domain in \mathbb{R}^n , $X = \partial/\partial x^n$, and \mathcal{N} is the hyperplane $\{x_n = 0\}$. In this case, we easily verify the claim by straightforward calculations.

Proof of Lemma 7.5. Let ρ be a defining function of M in U with $\text{grad } \rho(x) = \nu(x)$ for $x \in \partial M$. Then

$$\begin{aligned} \mathbf{G}_\mu(\rho \circ \pi)(x, \xi) &= \langle \xi, \text{grad } \rho(x) \rangle, \\ \mathbf{G}_\mu^2(\rho \circ \pi)(x, \xi) &= \langle \xi, \nabla_\xi \text{grad } \rho(x) \rangle + \langle Y_x(\nu), \text{grad } \rho(x) \rangle. \end{aligned}$$

Therefore, for $(x, \xi) \in S(\partial M)$ we have $\mathbf{G}_\mu(\rho \circ \pi)(x, \xi) = 0$, while $\mathbf{G}_\mu^2(\rho \circ \pi)(x, \xi) = -\Lambda(x, \xi) + \langle Y_x(\xi), \nu(x) \rangle \neq 0$ by strict magnetic convexity of ∂M .

We arrive directly at the above example if we take $\Sigma = S\tilde{M}$, $\mathcal{R} = \rho \circ \pi$, $\mathcal{M} = \partial(SM)$, $\mathcal{N} = \partial(S\bar{U})$, and $X = \mathbf{G}_\mu$. This completes the proof of the lemma. \square

Define the extension operator

$$A : C(\partial_+ SM) \rightarrow C(SM)$$

by

$$Aw(x, \xi) = \begin{cases} w(x, \xi), & (x, \xi) \in \partial_+ SM, \\ w \circ \mathcal{S}^{-1}(x, \xi), & (x, \xi) \in \partial_- SM, \end{cases}$$

where \mathcal{S} is the scattering relation.

Lemma 7.6. (Cf. [32, Lemma 1.1].) *If (M, g, α) is a simple magnetic system, then*

$$C_\alpha^\infty(\partial_+ SM) = \{w \in C^\infty(\partial_+ SM) : Aw \in C^\infty(\partial(SM))\}.$$

Proof. If $w^\sharp \in C^\infty(SM)$, then $Aw = w^\sharp|_{\partial(SM)}$ is smooth as well. Let us prove the converse. If $Aw \in C^\infty(\partial(SM))$, then from Lemma 7.5 and [17, Theorem C.4.4] we deduce the existence of a smooth function v on a neighborhood of the range $\Phi(\partial(SM))$ such that $w = v \circ \Phi$.

Let $\Psi : SM \rightarrow \partial_- SM$ denote the map

$$\Psi(x, \xi) = \psi^{\ell(x, \xi)}(x, \xi), \quad (x, \xi) \in SM,$$

and $\Psi_U : S\bar{U} \rightarrow \partial_- S\bar{U}$ the map

$$\Psi_U(x, \xi) = \psi^{\ell^+(x, \xi)}(x, \xi), \quad (x, \xi) \in S\bar{U}.$$

Note that $w^\sharp = \mathcal{S}^{-1} \circ \Psi$. Therefore, $w^\sharp = v \circ \Phi \circ \mathcal{S}^{-1} \circ \Psi$. It is easy to see that $\Phi \circ \mathcal{S}^{-1} \circ \Psi = \Psi_U|_{SM}$. Since Ψ_U is smooth on SM , we conclude that $w^\sharp \in C^\infty(SM)$, i.e., $w \in C_\alpha^\infty(\partial_+ SM)$. \square

7.3. Hilbert transform

From now on, we let (M, g) be an oriented Riemannian surface. Given $v \in T_x M$, we will denote by v_\perp the vector obtained by rotating v by $\pi/2$ according to the orientation of M . In coordinates $(v_\perp)_i = \varepsilon_{ij} v^j$, where

$$\varepsilon = \sqrt{\det g} \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix}.$$

Consider the fiberwise Hilbert transform [32, (1.4)]

$$Hu(x, \xi) = \frac{1}{2\pi} \int_{S_x M} \frac{1 + (\xi, \eta)}{(\xi_\perp, \eta)} u(x, \eta) d\sigma_x(\eta), \quad \xi \in S_x M.$$

If we fix $x \in M$ and a reference point $a \in S_x M$, any function on the fiber S_x can be treated as a function of an angular variable. Then

$$Hu(x, \theta) = \frac{1}{2\pi} \int_0^{2\pi} \cot\left(\frac{\varphi - \theta}{2}\right) u(x, \varphi) d\varphi.$$

Define

$$(\mathbf{G}_\perp u)(x, \xi) = \langle \xi_\perp, \nabla^\perp u \rangle = -\langle \xi, \nabla_\perp u \rangle,$$

where $\nabla_\perp u = \varepsilon \nabla^\perp u$. The following commutation formula holds [32, Theorem 1.5]:

$$[H, \mathbf{G}]u = \mathbf{G}_\perp(u_0) + (\mathbf{G}_\perp u)_0, \tag{7.8}$$

where

$$u_0(x) = \frac{1}{2\pi} \int_{S_x M} u(x, \xi) d\sigma_x(\xi)$$

is the average value on a fiber.

Let V be the infinitesimal generator of the action of S^1 on the fibers of the canonical projection $\pi : SM \rightarrow M$. Then the generator of the magnetic flow ψ^t is given by

$$\mathbf{G}_\mu = \mathbf{G} + \lambda V,$$

where λ is a function on M such that $\Omega = \lambda \Omega_a$, with Ω_a the area form of g .

Since H commutes with λV , we get from (7.8)

$$[H, \mathbf{G}_\mu]u = \mathbf{G}_\perp(u_0) + (\mathbf{G}_\perp u)_0. \tag{7.9}$$

Substitute $u = w^\sharp$, $w \in C^\infty_\alpha(\partial_+ SM)$, into (7.9). This yields

$$\mathbf{G}_\mu H w^\sharp = -\mathbf{G}_\perp(w^\sharp)_0 - (\mathbf{G}_\perp w^\sharp)_0. \tag{7.10}$$

Define the operator

$$B : C(SM) \rightarrow C(\partial_+ SM)$$

by

$$Bu(x, \xi) = u(x, \xi) - u \circ S(x, \xi), \quad (x, \xi) \in \partial_+ SM.$$

Clearly,

$$IG_\mu u = -Bu. \tag{7.11}$$

Using (7.11), we deduce from (7.10) the identity

$$BHAw = I(\mathbf{G}_\perp(w^\sharp)_0) + I((\mathbf{G}_\perp w^\sharp)_0), \tag{7.12}$$

since $w^\sharp|_{\partial(SM)} = Aw$.

Note that

$$\mathbf{G}_\perp(w^\sharp)_0 = -\frac{1}{2\pi} \langle \xi, \nabla_\perp \mathcal{I}_0^* w \rangle$$

and

$$(\mathbf{G}_\perp w^\sharp)_0 = -\frac{1}{2\pi} \delta_\perp \mathcal{I}_1^* w,$$

where

$$\delta_\perp v = -\delta v_\perp.$$

Hence, we can rewrite (7.12) as

$$BHAw = -\frac{1}{2\pi} \mathcal{I}[\nabla_\perp \mathcal{I}_0^* w, \delta_\perp \mathcal{I}_1^* w]. \tag{7.13}$$

7.4. Dirichlet-to-Neumann map

Given a compact Riemannian manifold (M, g) with boundary, denote the Laplace–Beltrami operator associated with g by Δ_g . Consider the Dirichlet problem

$$\Delta_g u = 0 \quad \text{on } M, \quad u|_{\partial M} = f.$$

The DN map is defined by

$$\Lambda_g(f) = \langle v, \nabla u|_{\partial M} \rangle.$$

In the two-dimensional case, the DN map can also be described as follows. Let (h, h_*) be a pair of conjugate harmonic functions on M ,

$$\nabla h = \nabla_\perp h_*, \quad \nabla h_* = -\nabla_\perp h.$$

Let h^0 and h_*^0 denote their traces on ∂M . Then

$$\Lambda(h_*^0) = \langle v, \nabla h_*^0|_{\partial M} \rangle = -\langle v, (\nabla_{\perp} h)|_{\partial M} \rangle = \langle v_{\perp}, (\nabla h)|_{\partial M} \rangle = \langle v_{\perp}, \nabla_{\partial M} h^0 \rangle, \tag{7.14}$$

where $\nabla_{\partial M}$ is the gradient w.r.t. the induced metric on ∂M .

The following theorem is an analog of [32, Theorem 1.3] and states that the scattering relation of a simple magnetic surface (M, g, α) completely determines the DN map of the metric g .

Theorem 7.7. *Let (g, α) and (g', α') be simple magnetic systems on a compact surface M with boundary such that $g|_{\partial M} = g'|_{\partial M}$. Assume that the scattering relations \mathcal{S} and \mathcal{S}' of these systems coincide. Then $\Lambda_g = \Lambda_{g'}$.*

Proof. Assume that h, h_* is a pair of smooth conjugate harmonic functions on M . Then $\mathbf{G}_{\perp} h = \mathbf{G}h_* = \mathbf{G}_{\mu} h_*$. By Theorem 7.3, there are $w \in C^{\infty}_{\alpha}(\partial_+ SM)$ and $f \in C^{\infty}(M)$ satisfying $\mathcal{I}_0^* w = h$, $\mathcal{I}_1^* w = \nabla f$. From (7.13) we then obtain

$$BHAw = -\frac{1}{2\pi} B h_*^0, \tag{7.15}$$

since $\delta_{\perp} \nabla f = 0$. Hence, the following holds:

Lemma 7.8. *(Cf. [32, Theorem 1.6].) If h, h_* is a pair of smooth conjugate harmonic functions on M , then there is $w \in C^{\infty}_{\alpha}(\partial_+ SM)$ such that $h = \mathcal{I}_0^* w$ and Eq. (7.15) holds with h_*^0 the trace of h_* on ∂M .*

In the opposite direction we have:

Lemma 7.9. *(Cf. [32, Theorem 1.6].) Suppose $h_*^0 \in C^{\infty}(\partial M)$ and $w \in C^{\infty}_{\alpha}(\partial_+ SM)$ satisfy Eq. (7.15). Define $h := \mathcal{I}_0^* w$ and let h_* be the harmonic continuation of h_*^0 to M . Then h and h_* are conjugate harmonic functions.*

Proof. Let q be an arbitrary smooth extension of h_*^0 to M . Note that

$$\mathbf{G}_{\mu} q = \mathbf{G}q = \langle \xi, \nabla q \rangle.$$

Using (7.13), we can therefore rewrite (7.15) as

$$-\mathcal{I}[\nabla_{\perp} h, \varphi] = \mathcal{I}[\nabla q, 0],$$

with $\varphi = \delta_{\perp} \mathcal{I}_1^* w$.

Thus,

$$\mathcal{I}[\nabla_{\perp} h + \nabla q, \varphi] = 0.$$

By (7.2) we then have $\varphi = 0$ and $\nabla q + \nabla_{\perp} h = \nabla p$ for some smooth function p on M with $p|_{\partial M} = 0$. Therefore, h and $q - p$ are conjugate harmonic functions. Since $(q - p)|_{\partial M} = h_*^0$, we have $q - p = h_*$, which implies that h and h_* are conjugate harmonic functions. \square

Continuing the proof of Theorem 7.7, we have the following procedure to obtain the DN map from the scattering relation. For an arbitrary given smooth function h_*^0 on ∂M we find a function $w \in C_\alpha^\infty(\partial_+ SM)$ that solves Eq. (7.15). Then the functions $h^0 = 2\pi(Aw)_0$ (notice that $2\pi(Aw)_0 = \mathcal{I}_0^* w|_{\partial M}$) and h_*^0 are traces of conjugate harmonic functions. This gives the DN map by means of (7.14). \square

7.5. Proof of Theorem 7.1

We end up with the proof of this theorem in the same way as in [32]. If two simple magnetic systems (g, α) and (g', α') on a compact surface with boundary have the same boundary action functions, then by Theorem 2.2 we may assume that $g|_{\partial M} = g'|_{\partial M}$ and, by Lemma 2.5, their scattering relations coincide. Theorem 7.7 then tells us that the DN maps of the metrics g and g' coincide. Now, the result of [19,21] implies the existence of a diffeomorphism $f : M \rightarrow M$, which is the identity on ∂M , and of a function ω such that $g' = \omega^2 f^* g$. Next, Theorem 6.1 yields $\omega = 1$ and $\alpha' = f^* \alpha + d\varphi$ for a smooth φ vanishing on ∂M . This concludes the proof of the theorem.

Appendix A. Geometry of magnetic systems

A.1. Mañé’s critical value and simplicity

Here we adapt a certain part of the theory of convex superlinear Lagrangians to the case of manifolds with boundary.

Let M be a compact Riemannian manifold with boundary and let $L : TM \rightarrow \mathbb{R}$ be a C^∞ Lagrangian satisfying the following hypotheses:

- *Convexity*: For all $x \in M$ the restriction of L to $T_x M$ has everywhere positive definite Hessian.
- *Superlinear growth*:

$$\lim_{|v| \rightarrow \infty} \frac{L(x, v)}{|v|} = +\infty$$

uniformly on $x \in M$.

The action of L on an absolutely continuous curve $\gamma : [a, b] \rightarrow M$ is

$$\mathbb{A}_L(\gamma) = \int_a^b L(\gamma(t), \dot{\gamma}(t)) dt.$$

For each $k \in \mathbb{R}$, the *Mañé action potential* $\mathbb{A}_k : M \times M \rightarrow \mathbb{R} \cup \{-\infty\}$ is defined by

$$\mathbb{A}_k(x, y) = \inf_{\gamma \in \mathcal{C}(x, y)} \mathbb{A}_{L+k}(\gamma),$$

where $\mathcal{C}(x, y) = \{\gamma : [0, T] \rightarrow M : \gamma(0) = x, \gamma(T) = y, \gamma \text{ is absolutely continuous}\}$.

The *critical level* $c = c(L)$ is defined as

$$\begin{aligned} c(L) &= \sup\{k \in \mathbb{R}: \mathbb{A}_{L+k}(\gamma) < 0 \text{ for some closed curve } \gamma\} \\ &= \inf\{k \in \mathbb{R}: \mathbb{A}_{L+k}(\gamma) \geq 0 \text{ for every closed curve } \gamma\}. \end{aligned}$$

Proposition A.1. For $k < c(L)$, $\mathbb{A}_k(x, y) = -\infty$ for all $x, y \in M$. For $k \geq c(L)$, $\mathbb{A}_k(x, y) \in \mathbb{R}$ for all $x, y \in M$.

Proof. The same as in [6, Proposition 2-1.1]. \square

One more characterization of $c(L)$ is useful. Recall that the Hamiltonian $H : T^*M \rightarrow \mathbb{R}$ associated with L is defined by the Fenchel transform

$$H(x, p) = \sup\{p(v) - L(x, v) : v \in T_x M\},$$

and the supremum is achieved at v such that $p = \frac{\partial L}{\partial v}(x, v)$.

Proposition A.2. If there exists a C^1 function $f : M \rightarrow \mathbb{R}$ such that $H(df) < k$, then $k \geq c(L)$.

Proof. The same as in [7, Lemma 5]. \square

Recall that the *energy function* $E : TM \rightarrow \mathbb{R}$ for L is defined by

$$E(x, v) = \frac{\partial L}{\partial v}(x, v) \cdot v - L(x, v),$$

and that the energy function is constant on every solution $x(t)$ of the Euler–Lagrange equation

$$\frac{d}{dt} \frac{\partial L}{\partial v}(x(t), \dot{x}(t)) = \frac{\partial L}{\partial x}(x(t), \dot{x}(t)). \quad (\text{A.1})$$

Let $\psi^t : TM \rightarrow TM$ be the Euler–Lagrange flow, defined by $\psi^t(x, v) = (\gamma(t), \dot{\gamma}(t))$, where γ is the solution of (A.1) with $\gamma(0) = x$ and $\dot{\gamma}(0) = v$. For $x \in M$ and $k \in \mathbb{R}$, the *exponential map* at x of energy k is defined to be the partial map $\exp_x : T_x M \rightarrow M$ given by

$$\exp_x^k(tv) = \pi \circ \psi^t(v), \quad t \geq 0, v \in T_x M, E(x, v) = k.$$

Then \exp_x^k is a C^1 -smooth partial map on $T_x M$ which is C^∞ -smooth on $T_x M \setminus \{0\}$.

The next proposition is similar to [8, Theorem D] and has a similar proof.

Proposition A.3. If $\exp_x^k : (\exp_x^k)^{-1}(M) \rightarrow M$ is a diffeomorphism for every $x \in M$, then $k \geq c(L)$.

Proof. Fix $q \in M$. Given $x \in M$, let $\gamma_{q,x} : [0, T_{q,x}] \rightarrow M$ be a solution of the Euler–Lagrange equation with energy k , joining q to x . Consider the function $f(x) = \mathbb{A}_k(\gamma_{q,x})$. It is easy to see

that the assumption of the lemma implies this function is smooth in $M \setminus \{q\}$. It follows from the first variation formula of [8, Lemma 4] that for $x \in M \setminus \{q\}$

$$d_x f(w) = \frac{\partial L}{\partial v}(x, \dot{\gamma}_{q,x}(T_{q,x})) \cdot w,$$

which implies

$$H(x, d_x f) = E(x, \dot{\gamma}_{q,x}(T_{q,x})) = k.$$

This last equation also implies that $|d_x f|$ is uniformly bounded for all $x \in M \setminus \{q\}$. Thus f is Lipschitz in M and a smoothing argument as in [7,14] shows that for any $\varepsilon > 0$ there exists $\tilde{f} \in C^\infty(M)$ for which $H(x, d_x \tilde{f}) < k + \varepsilon$ for all $x \in M$. Thus by Proposition A.2, $k \geq c(L)$. \square

The next proposition is an analog of [6, Proposition 3-5.1] and has the same proof.

Proposition A.4. *If $k > c(L)$ and $x, y \in M$ $x \neq y$, then there is $\gamma \in \mathcal{C}(x, y)$ such that*

$$\mathbb{A}_k(x, y) = \mathbb{A}_{L+k}(\gamma).$$

Moreover, the energy of γ is $E(\gamma, \dot{\gamma}) \equiv k$.

Now, we apply the above to the case of magnetic systems. For a simple magnetic system (M, g, α) , the magnetic flow can also be obtained as the Euler–Lagrange flow with the corresponding Lagrangian defined by

$$L(x, v) = \frac{1}{2}|v|_g^2 - \alpha_x(v).$$

Lemma A.5. *Let (g, α) be a simple magnetic system on M . For $x, y \in M$, $x \neq y$,*

$$\mathbb{A}_{1/2}(x, y) = \mathbb{A}_{L+1/2}(\gamma_{x,y}) = T_{x,y} - \int_{\gamma_{x,y}} \alpha,$$

where $\gamma_{x,y} : [0, T_{x,y}] \rightarrow M$ is the unit speed magnetic geodesic from x to y .

Proof. It is easy to see that the simplicity assumption implies that for this Lagrangian the assumptions of Proposition A.3 hold for all k sufficiently close to $1/2$. Therefore, the proposition gives $1/2 > c(L)$. Then Proposition A.4 shows that, given $x \neq y$ in M , there is $\gamma \in \mathcal{C}(x, y)$ with energy $1/2$ (i.e., γ is parametrized by arc length) such that $\mathbb{A}(x, y) = \mathbb{A}(\gamma)$. Using simplicity, one can then prove that γ is a unit speed magnetic geodesic, i.e., $\gamma = \gamma_{x,y}$. \square

A.2. Magnetic convexity

Let M be a compact manifold with boundary, endowed with a Riemannian metric g and a closed 2-form Ω . Consider a manifold M_1 such that $M_1^{\text{int}} \supset M$. Extend g and Ω to M_1 smoothly, preserving the former notation for extensions. We say that M is *magnetic convex* at $x \in \partial M$ if there is a neighborhood U of x in M_1 such that all unit speed magnetic geodesics in U , passing

through x and tangent to ∂M at x , lie in $M_1 \setminus M^{\text{int}}$. If, in addition, these geodesics do not intersect M except for x , we say that M is *strictly magnetic convex* at x . It is not hard to show that these definitions depend neither on the choice of M_1 nor on the way we extend g and Ω to M_1 .

As before, we let Λ denote the second fundamental form of ∂M and $\nu(x)$ the inward unit normal to ∂M at x .

Lemma A.6. *If M is magnetic convex at $x \in \partial M$, then*

$$\Lambda(x, v) \geq \langle Y_x(v), \nu(x) \rangle \quad \text{for all } v \in S_x(\partial M). \tag{A.2}$$

If the inequality is strict, then M is strictly magnetic convex at x .

Proof. Suppose M is convex at x . Choosing a smaller U if necessary, we may assume that there is a smooth function ρ on U such that $|\text{grad } \rho| = 1$ and $\partial M \cap U = \rho^{-1}(0)$. Further we may assume that all the above geodesics lie in $U^- = \{x: \rho(x) \leq 0\}$.

Let $v \in S_x(\partial M)$ and $\gamma(t)$ be the magnetic geodesic with $\gamma(0) = x, \dot{\gamma}(0) = v$. By assumption, $\rho \circ \gamma(t) \leq 0$ for all small t . Therefore,

$$\frac{d^2}{dt^2} [\rho \circ \gamma(t)]|_{t=0} \leq 0.$$

Since

$$\begin{aligned} \frac{d^2}{dt^2} [\rho \circ \gamma(t)] &= \frac{d}{dt} \langle \text{grad } \rho(\gamma(t)), \dot{\gamma}(t) \rangle \\ &= \langle \nabla_{\dot{\gamma}(t)} \text{grad } \rho(\gamma(t)), \dot{\gamma}(t) \rangle + \langle \text{grad } \rho(\gamma(t)), \ddot{\gamma}(t) \rangle \\ &= \text{Hess}_{\gamma(t)} \rho(\dot{\gamma}(t), \dot{\gamma}(t)) + \langle \text{grad } \rho(\gamma(t)), Y(\dot{\gamma}(t)) \rangle \end{aligned}$$

and since $\Lambda(x, \xi) = -\text{Hess}_x \rho(v, v)$ and $\text{grad } \rho(x) = \nu(x)$ when $(x, v) \in \partial(SM)$, we arrive at (A.2).

Now, assume that (A.2) is strict. Then there is $\delta > 0$ such that for every magnetic geodesic γ in N with $\gamma(0) = x$ and $\dot{\gamma}(0) = v \in S_x(\partial M)$,

$$\frac{d^2}{dt^2} [\rho \circ \gamma(t)]|_{t=0} \leq -\delta.$$

Thus, there is a small $\varepsilon > 0$ such that

$$\rho \circ \gamma(t) \leq -\frac{1}{4} \delta t^2 \quad \text{for all } t \in (-\varepsilon, \varepsilon).$$

This implies the second statement. \square

A.3. Exponential map

Lemma A.7. *The map $\exp_x^\mu : T_x M \rightarrow M$ is C^1 and C^∞ on $T_x M \setminus \{0\}$. This map is C^2 if and only if $\Omega = 0$.*

Proof. Recall that $\exp_x^\mu(tv) = \pi \circ \psi^t(x, v) := \gamma(t, x, v)$ where v has norm one and $t \geq 0$. Clearly this implies that \exp_x^μ is C^∞ on $T_x M \setminus \{0\}$. Since

$$\left. \frac{d}{dt} \right|_{t=0} \pi \circ \psi^t(x, v) = v$$

we see that the directional derivative of \exp_x^μ at 0 in the direction of v exists and equals v . In a coordinate system around x write

$$\exp_x^{\mu,i}(tv) = \gamma^i(t, x, v) \tag{A.3}$$

where γ^i depends smoothly in (t, x, v) . Differentiating with respect to v we obtain

$$\frac{\partial \exp_x^{\mu,i}}{\partial v^j}(tv)t = \frac{\partial \gamma^i}{\partial v^j}(t, x, v).$$

Since $\gamma^i(0, x, v) = x^i$, $\frac{\partial \gamma^i}{\partial v^j}(0, x, v) = 0$. Therefore,

$$\begin{aligned} \lim_{t \rightarrow 0^+} \frac{\partial \exp_x^{\mu,i}}{\partial v^j}(tv) &= \lim_{t \rightarrow 0^+} \frac{1}{t} \frac{\partial \gamma^i}{\partial v^j}(t, x, v) \\ &= \frac{\partial^2 \gamma^i}{\partial t \partial v^j}(0, x, v) = \frac{\partial^2 \gamma^i}{\partial v^j \partial t}(0, x, v) = \delta_j^i \end{aligned}$$

since $\dot{\gamma}^i(0, x, v) = v^i$. Since $S_x M$ is compact, the above limit is uniform in $v \in S_x M$ and thus \exp_x^μ has continuous partial derivatives at 0, i.e., \exp_x^μ is C^1 .

Suppose now that \exp_x^μ is C^2 . Differentiate (A.3) twice with respect to t ($v \in S_x M$)

$$v^j v^k \frac{\partial^2 \exp_x^{\mu,i}}{\partial v^j \partial v^k}(tv) = \ddot{\gamma}^i(t, x, v).$$

Using the equations of a magnetic geodesic

$$\ddot{\gamma}^i + \dot{\gamma}^j \dot{\gamma}^k \Gamma_{jk}^i = Y_k^i \dot{\gamma}^k$$

we obtain

$$v^j v^k \frac{\partial^2 \exp_x^{\mu,i}}{\partial v^j \partial v^k}(tv) = -\dot{\gamma}^j \dot{\gamma}^k \Gamma_{jk}^i + Y_k^i \dot{\gamma}^k.$$

Let $t \rightarrow 0^+$. Then

$$v^j v^k \left(\frac{\partial^2 \exp_x^{\mu,i}}{\partial v^j \partial v^k}(0) + \Gamma_{jk}^i(x) \right) = Y_k^i(x) v^k.$$

Since this holds for all $v \in S_x M$ we must have $Y_k^i(x) = 0$, i.e., $\Omega = 0$. \square

A.4. Santaló’s formula

If (M, g) is a compact Riemannian manifold with boundary, we endow its unit sphere bundle SM with its usual Liouville (local product) measure $d\Sigma^{2n-1}$, and endow the bundle $\partial_+ SM$ with its standard measure $d\Sigma^{2n-2}$ (again a local product measure where the measure on the fiber is the measure of a hemisphere). Denote by $d\mu$ the measure on $\partial_+ SM$ given by

$$d\mu(x, \xi) = \langle \xi, \nu(x) \rangle d\Sigma^{2n-2}(x, \xi),$$

where $\nu(x)$ is the inward unit normal of ∂M at a point x .

The following version of Santaló’s formula holds for magnetic flows.

Lemma A.8. *Suppose that (M, g, α) is simple. Then for every continuous function $\varphi : SM \rightarrow \mathbb{R}$ we have*

$$\int_{SM} \varphi d\Sigma^{2n-1} = \int_{\partial_+ SM} d\mu(x, \xi) \int_0^{\ell(x, \xi)} \varphi(\gamma_{x, \xi}(t), \dot{\gamma}_{x, \xi}(t)) dt. \tag{A.4}$$

Proof. The argument we use is the same as in [33]. We give it for the sake of completeness.

First, we recall the well-known fact that the Liouville measure is invariant under the magnetic flow (for example, because $\omega_0^{\wedge n} = (\omega_0 + \pi^* \Omega)^{\wedge n}$ while $\omega_0 + \pi^* \Omega$ is flow invariant). Now, let $D = \{(x, \xi; t) \in \partial_+ SM : 0 \leq t \leq \ell(x, \xi)\}$, and define $\Psi : D \rightarrow SM$ by $\Psi(x, \xi; t) = \psi^t(x, \xi)$, where ψ^t is the magnetic flow. Then

$$\int_{SM} \varphi d\Sigma^{2n-1} = \int_D (\varphi \circ \Psi) \Psi^*(d\Sigma^{2n-1}). \tag{A.5}$$

By construction, Ψ conjugates ψ^t with the flow generated by $\partial/\partial t$ on D . Since $d\Sigma^{2n-1}$ is invariant under ψ^t , $\Psi^*(d\Sigma^{2n-1})$ is invariant under the flow of $\partial/\partial t$. Then

$$\Psi^*(d\Sigma^{2n-1})(x, \xi; t) = a(x, \xi) d\Sigma^{2n-2} \wedge dt$$

for some function a on $\partial_+ SM$, so that (A.5) takes the form

$$\int_{SM} \varphi d\Sigma^{2n-1} = \int_{\partial_+ SM} a(x, \xi) d\Sigma^{2n-2}(x, \xi) \int_0^{\ell(x, \xi)} \varphi(\gamma_{x, \xi}(t), \dot{\gamma}_{x, \xi}(t)) dt.$$

We are left with proving that $a(x, \xi) = \langle \xi, \nu(x) \rangle$. To this end, it suffices to show that

$$\Psi^*(d\Sigma^{2n-1})(x, \xi; 0) = \langle \xi, \nu(x) \rangle d\Sigma^{2n-2} \wedge dt.$$

Note that $d\Sigma^{2n-1} = -d\Sigma^{2n-2} \wedge dr$ on $\partial(SM)$, where $r(x) = \text{dist}(x, \partial M)$ is the distance function from x to ∂M . Since the differential of Ψ is the identity on $T_{(x,\xi)}(\partial_+ SM)$ and takes $\partial/\partial t$ to the generator \mathbf{G}_μ of the flow ψ^t , we have

$$\Psi^*(d\Sigma^{2n-1})(x, \xi; 0) = (\mathbf{G}_\mu r) d\Sigma^{2n-2}(x, \xi) \wedge dt.$$

As soon as $\mathbf{G}_\mu r = \langle \xi, \nabla r \rangle = -\langle \xi, \nu(x) \rangle$, we are done. \square

A.5. Index form of a magnetic geodesic

Let (M, g, α) be a simple magnetic system. For every $x \in M$, $\exp_x^\mu : T_x M \rightarrow M$ is a diffeomorphism restricted to a suitable set in $T_x M$ which is diffeomorphic to a closed ball.

Let $\pi : SM \rightarrow M$ be the canonical projection and let for $v \in SM$,

$$V(v) := \ker d_v \pi,$$

which is an $(n - 1)$ -dimensional subspace of $T_v SM$, and

$$E(v) := V(v) \oplus \mathbb{R}\mathbf{G}_\mu(v). \tag{A.6}$$

Lemma A.9. *If $\gamma : [0, T] \rightarrow M$ is a unit speed magnetic geodesic, then*

$$d_{\dot{\gamma}(0)} \psi^t(E) \cap V(\dot{\gamma}(t)) = \{0\}$$

for every $t \in (0, T]$.

Proof. Take $v \in SM$ and $t \in (0, T]$. From the definition of \exp_x^μ one sees right away that

$$\text{image}(d_{tv} \exp_x^\mu) = d_{\dot{\gamma}(t)} \pi(d_{\dot{\gamma}(0)} \psi^t(E)).$$

Since $d_w \exp_x^\mu$ is a linear isomorphism for every $w \in T_x M$ at which \exp_x^μ is defined, the lemma follows. \square

Given a unit speed magnetic geodesic $\gamma : [0, T] \rightarrow M$, let \mathcal{A} and \mathcal{C} be the operators on smooth vector fields along γ defined by

$$\begin{aligned} \mathcal{A}(Z) &= \ddot{Z} + R(\dot{\gamma}, Z)\dot{\gamma} - Y(\dot{Z}) - (\nabla_Z Y)(\dot{\gamma}), \\ \mathcal{C}(Z) &= R(\dot{\gamma}, Z)\dot{\gamma} - Y(\dot{Z}) - (\nabla_Z Y)(\dot{\gamma}). \end{aligned}$$

A vector field J along γ is said to be a *magnetic Jacobi field* if it satisfies the equation

$$\mathcal{A}(J) = 0. \tag{A.7}$$

Let Λ denote the \mathbb{R} -vector space of smooth vector fields Z along γ such that $Z(0) = Z(T) = 0$. Define the quadratic form $\text{Ind} : \Lambda \rightarrow \mathbb{R}$ by

$$\text{Ind}(Z, Z) = \int_0^T \{ |\dot{Z}|^2 - \langle \mathcal{C}(Z), Z \rangle - \langle Y(\dot{\gamma}), Z \rangle^2 \} dt.$$

Note that

$$\text{Ind}(Z, Z) = - \int_0^T \{ \langle \mathcal{A}(Z), Z \rangle + \langle Y(\dot{\gamma}), Z \rangle^2 \} dt.$$

Clearly, $\text{Ind}(Z, Z)$ generalizes the index form of a geodesic in a Riemannian manifold. The next lemma is an analog of a well-known assertion from Riemannian geometry.

Lemma A.10 (Index Lemma). *If $Z \in \Lambda$ is orthogonal to $\dot{\gamma}$, then*

$$\text{Ind}(Z, Z) \geq 0,$$

with equality if and only if Z vanishes.

Proof. Note that the subbundle E defined by (A.6) is Lagrangian.

If $\xi \in E(v)$, then $J_\xi(t) = d\pi \circ d\psi^t(\xi)$ satisfies the Jacobi equation (A.7). Since

$$d\dot{\gamma}(t)\pi|_{E(\dot{\gamma}(t))} : E(\dot{\gamma}(t)) \rightarrow T_{\gamma(t)}M$$

is an isomorphism for all $t \in (0, T]$, there exists a basis $\{\xi_1, \dots, \xi_n\}$ for $E(v)$ such that $\{J_{\xi_1}(t), \dots, J_{\xi_n}(t)\}$ is a basis of $T_{\gamma(t)}M$ for all $t \in (0, T]$. Without loss of generality we may assume that $\xi_1 = \mathbf{G}_\mu(v)$, $J_{\xi_1} = \dot{\gamma}$ and $J_{\xi_i}(0) = 0$ for $i \geq 2$.

Let us set for brevity $J_i = J_{\xi_i}$. Then if Z is an element of Λ , we can write for $t \in (0, T]$

$$Z(t) = \sum_{i=1}^n f_i(t) J_i(t)$$

for some smooth functions f_1, \dots, f_n . The functions f_i can in fact be smoothly extended to $t = 0$. Indeed, for $i \geq 2$, we can write $J_i(t) = t A_i(t)$ where A_i is a smooth vector field such that $A_i(0) = \dot{J}_i(0)$. Since $\{\dot{\gamma}(t), A_2(t), \dots, A_n(t)\}$ is now a basis for all $t \in [0, T]$, there exist smooth functions g_i such that for all $t \in [0, T]$

$$Z(t) = g_1(t)\dot{\gamma}(t) + \sum_{i=2}^n g_i(t)A_i(t).$$

Therefore for $t \in (0, T]$, $g_1(t) = f_1(t)$ and for $i \geq 2$, $g_i(t) = t f_i(t)$. Since $Z(0) = 0$, $g_i(0) = 0$ for all i and the f_i 's smoothly extend to $t = 0$.

Now we can write

$$\text{Ind}(Z, Z) = - \sum_{i,j} \int_0^T \langle \mathcal{A}(f_i J_i), f_j J_j \rangle dt - \int_0^T \langle Y(\dot{\gamma}), Z \rangle^2 dt. \tag{A.8}$$

An easy computation shows that

$$\mathcal{A}(f_i J_i) = \ddot{f}_i J_i + 2\dot{f}_i \dot{J}_i - \dot{f}_i Y(J_i) + f_i \mathcal{A}(J_i).$$

Since J_i satisfies Eq. (A.7), then $\mathcal{A}(J_i) = 0$ and hence,

$$\langle \mathcal{A}(f_i J_i), J_j \rangle = \ddot{f}_i \langle J_i, J_j \rangle + 2\dot{f}_i \langle \dot{J}_i, J_j \rangle - \dot{f}_i \langle Y(J_i), J_j \rangle.$$

Observe that since E is a Lagrangian subspace,

$$\langle J_i, \dot{J}_j \rangle - \langle \dot{J}_i, J_j \rangle + \langle Y(J_i), J_j \rangle = 0,$$

and then

$$\langle \mathcal{A}(f_i J_i), J_j \rangle = \frac{d}{dt} (\dot{f}_i \langle J_i, J_j \rangle).$$

Now we can write

$$\int_0^T \langle \mathcal{A}(f_i J_i), f_j J_j \rangle dt = \langle \dot{f}_i J_i, f_j J_j \rangle \Big|_0^T - \int_0^T \langle \dot{f}_i J_i, \dot{f}_j J_j \rangle dt.$$

Combining the last equality with (A.8) we obtain

$$\text{Ind}(Z, Z) = \int_0^T \left\| \sum_{i=1}^n \dot{f}_i J_i \right\|^2 dt - \left\langle \sum_{i=1}^n \dot{f}_i J_i, Z \right\rangle \Big|_0^T - \int_0^T \langle Y(\dot{\gamma}), Z \rangle^2 dt.$$

But $Z(0) = Z(T) = 0$, therefore

$$I(V, V) = \int_0^T \left\| \sum_{i=1}^n \dot{f}_i J_i \right\|^2 dt - \int_0^T \langle Y(\dot{\gamma}), Z \rangle^2 dt. \tag{A.9}$$

Now let

$$W := \sum_{i=2}^n \dot{f}_i J_i.$$

Since $J_1 = \dot{\gamma}$ we have:

$$\left\langle \sum_{i=1}^n \dot{f}_i J_i, \sum_{i=1}^n \dot{f}_i J_i \right\rangle = \langle \dot{f}_1 \dot{\gamma} + W, \dot{f}_1 \dot{\gamma} + W \rangle = \dot{f}_1^2 + 2\dot{f}_1 \langle \dot{\gamma}, W \rangle + \langle W, W \rangle.$$

Differentiating $\langle Z, \dot{\gamma} \rangle = 0$ we get

$$\langle \dot{Z}, \dot{\gamma} \rangle + \langle Z, Y(\dot{\gamma}) \rangle = 0.$$

But

$$\langle \dot{Z}, \dot{\gamma} \rangle = \left\langle \sum_{i=1}^n \dot{f}_i J_i, \dot{\gamma} \right\rangle = \dot{f}_1 + \langle W, \dot{\gamma} \rangle$$

since $\langle \dot{J}_i, \dot{\gamma} \rangle = 0$ for all i . Therefore

$$\langle Y(\dot{\gamma}), Z \rangle^2 = f_1^2 + 2f_1 \langle W, \dot{\gamma} \rangle + \langle W, \dot{\gamma} \rangle^2.$$

Thus

$$\left\langle \sum_{i=1}^n \dot{f}_i J_i, \sum_{i=1}^n \dot{f}_i J_i \right\rangle - \langle Y(\dot{\gamma}), Z \rangle^2 = \langle W, W \rangle - \langle W, \dot{\gamma} \rangle^2.$$

If we let W^\perp be the orthogonal projection of W to $\dot{\gamma}^\perp$, the last equation and (A.9) give:

$$\text{Ind}(Z, Z) = \int_0^T \|W^\perp\|^2 dt \geq 0$$

with equality if and only if W^\perp vanishes identically. But if W^\perp vanishes, then

$$-\langle W, \dot{\gamma} \rangle \dot{\gamma} + \sum_{i=2}^n \dot{f}_i J_i = 0$$

which implies that the functions f_i are constant for $i \geq 2$. Thus Z is of the form $f_1 \dot{\gamma} + J$ where J is a magnetic Jacobi field. But $Z(T) = 0$ implies $J(T) = 0$. Since the J_i 's are linearly independent at T , J must vanish identically and since Z is orthogonal to $\dot{\gamma}$, Z must also vanish. \square

Appendix B. Study of a certain class of integral operators with singular kernels

As we mentioned before, the fact that the magnetic exponential map is smooth in polar coordinates only forces us to work in polar coordinates as well. In this appendix we study a class of operators that naturally arise in our analysis.

Let $U \subset \mathbb{R}^n$ be open and g be a smooth Riemannian metric in a neighborhood of \bar{U} .

Lemma B.1. *Let $A : C_0(U) \rightarrow C(U)$ be the operator*

$$Af(x) = \int_{S_x U} \int_{\mathbb{R}} A(x, r, \omega) f(x + r\omega) dr d\sigma_x(\omega), \tag{B.1}$$

with $A \in C^\infty(U \times \mathbb{R} \times S_x U)$. Then A is a classical Ψ DO of order -1 with full symbol

$$a(x, \xi) \sim \sum_0^\infty a_k(x, \xi),$$

where

$$a_k(x, \xi) = 2\pi \frac{i^k}{k!} \int_{S_x U} \partial_r^k A(x, 0, \omega) \delta^{(k)}(\omega \cdot \xi) d\sigma_x(\omega).$$

Proof. Notice first that if A is an odd function of (r, ω) , then $\mathcal{A}f = 0$. Therefore, we can replace A above by $A_{\text{even}}(r, \omega) = (A(r, \omega) + A(-r, -\omega))/2$. Next, it is easy to check that we can integrate over $r \geq 0$ only and double the result. Therefore,

$$\mathcal{A}f(x) = 2 \int_{S_x M} \int_0^\infty A_{\text{even}}(x, r, \omega) f(x + r\omega) dr d\sigma_x(\omega). \tag{B.2}$$

Consider now r, ω as polar coordinates for $z = r\omega$, and make also the change of variables $y = x + z$ to get

$$\mathcal{A}f(x) = 2(\det g(x))^{1/2} \int A_{\text{even}}\left(x, |y - x|_g, \frac{y - x}{|y - x|_g}\right) \frac{f(y)}{|y - x|_g^{n-1}} dy, \tag{B.3}$$

where the subscript g refers to $g(x)$. Let

$$A_{\text{even}}(x, r, \omega) = \sum_{k=0}^{N-1} A_{\text{even},k}(x, \omega) r^k + r^N R_N(x, r, \omega) \tag{B.4}$$

be a finite Taylor expansion of A_{even} in r near $r = 0$ with $N > 0$. It follows easily that $2A_{\text{even},k}(x, \omega) = A_k(x, \omega) + (-1)^k A_k(x, -\omega)$, where $k!A_k = \partial_r^k|_{r=0}A$, and in particular, $A_{\text{even},k}(x, \omega)r^k$ is even w.r.t. (r, ω) . The remainder term contributes to (B.3) an operator that maps $L^2_{\text{comp}}(U)$ into $H^{N-N_0}(U)$ with some fixed N_0 . To study the contribution of the other terms, write

$$\mathcal{A}_{\text{even},k}f(x) = 2(\det g(x))^{1/2} \int A_{\text{even},k}\left(x, \frac{y - x}{|y - x|_g}\right) |y - x|_g^{k-n+1} f(y) dy. \tag{B.5}$$

The kernel of $\mathcal{A}_{\text{even},k}$ is therefore a function of x and $z = y - x$, with a polynomial singularity at $y - x = 0$, and it is therefore a formal Ψ DO with symbol that can be obtained by taking Fourier transform in the z variable. Motivated by this, apply the Plancherel theorem to the integral above to get

$$\mathcal{A}_{\text{even},k}f(x) = (2\pi)^{-n} \int e^{ix \cdot \xi} a_k(x, \xi) \hat{f}(\xi) d\xi,$$

where

$$\begin{aligned} a_k(x, \xi) &= 2 \int e^{-iy \cdot \xi} A_{\text{even},k}\left(x, \frac{y - x}{|y - x|_g}\right) |y - x|_g^{k-n+1} (\det g(x))^{1/2} dy \\ &= 2 \int_{S_x U} \int_0^\infty e^{-ir\omega \cdot \xi} A_{\text{even},k}(x, \omega) r^k dr d\sigma_x(\omega) \\ &= \int_{S_x U} \int_{-\infty}^\infty e^{-ir\omega \cdot \xi} A_k(x, \omega) r^k dr d\sigma_x(\omega) \end{aligned}$$

$$= 2\pi i^k \int_{S_x U} A_k(x, \omega) \delta^{(k)}(\omega \cdot \check{\xi}) d\sigma_x(\omega). \tag{B.6}$$

In the third line, we used the fact that $A_{\text{even},k}(x, \omega)r^k$ is even. Note that $a_k(x, \xi)$ is homogeneous in ξ of order $-k - 1$ and smooth away from $\xi = 0$ but a distribution (in S') near zero. To deal with this, choose $\chi \in C_0^\infty$ supported in $|\xi| \leq 1$ and equal to 1 near $\xi = 0$. Write $a(x, \xi) = \chi(\xi)a(x, \xi) + (1 - \chi(\xi))a(x, \xi)$. The second term is a classical amplitude, while the first one contributes the term

$$\mathcal{A}_{\text{even},k}(\check{\chi} * f) \tag{B.7}$$

to (B.5) that is smooth, as can be easily seen by making the change of variables $z = y - x$ in (B.5), see also (B.16). \square

Remark B.2. If $A(x, r, \omega)$ and g are smooth of class C^k only, then \mathcal{A} is an Ψ DO with an amplitude of finite smoothness $l(k)$, admitting a finite expansion. If $k \gg 1$, then $l \gg 1$, and one can still construct a finite order parametrix of an elliptic Ψ DO in this class and the usual $H^{s_1} \rightarrow H^{s_2}$ estimates still hold, if $k \gg 1$, depending on s_1, s_2 . This has been used already in [37,38].

We return to the analysis of the singular operator \mathcal{A} introduced in Lemma B.1 under the assumption that A and g are analytic. Our reference for the calculus of analytic Ψ DOs is [40].

Lemma B.3. *Let $\mathcal{A}: C_0(U) \rightarrow C(U)$ be the operator (B.1) with $A(x, r, \omega)$ analytic for $(x, \omega) \in U \times S_x U$, and $r \in \mathbb{R}$ such that $x + r\omega \in U$. Then \mathcal{A} is an analytic Ψ DO of order -1 with a symbol expansion as in Lemma B.1.*

Proof. Notice first that \mathcal{A} is analytic pseudolocal, see [40, Theorem V.2.1].

By performing the change of variables $\omega' = g^{1/2}(x)\omega$, we reduce the lemma to the case where g is the Euclidean metric. Let $U' \Subset U$.

Let us estimate $a_k(x, \xi)$, see Lemma B.1. Since $A(x, r, \omega)$ is analytic, and a_k is homogeneous of order $-k - 1$, we have

$$|a_k(x, \xi)| \leq C^{k+1} k! |\xi|^{-k-1}$$

with some $C > 0$. Using the homogeneity, we get

$$|\partial_\xi^\alpha a_k(x, \xi)| \leq C^{k+|\alpha|+1} \alpha! k! |\xi|^{-k-|\alpha|-1} \tag{B.8}$$

for ξ is in a complex neighborhood of \mathbb{R} and x in a complex neighborhood of \bar{U}' . Therefore, there exists a pseudoanalytic symbol $a \sim \sum a_k$, see [40]. This symbol is defined by

$$a(x, \xi) = \sum_{k=0}^\infty \varphi_k(\xi) a_k(x, \xi), \tag{B.9}$$

where φ_k have the properties (see [40, V.2]): $0 \leq \varphi_k \leq 1$, $\varphi_k(\xi) = 0$ for $|\xi| < 2R \max(k, 1)$, $\varphi_k(\xi) = 1$ for $|\xi| > 3R \max(k, 1)$, $|D^\alpha \varphi_k| \leq (C/R)^{|\alpha|}$ for $|\alpha| \leq 2k$, where $R > 1$ is a large parameter. We will prove next that $a(x, D)$ differs from \mathcal{A} by an analytic regularizing operator.

Let $u \in \mathcal{E}'(U')$. Let $s \geq 0$ be such that u can be represented as a finite sum of derivatives of continuous functions of order not exceeding s .

By (B.4), for any $N \geq 1$,

$$\mathcal{A}u = \sum_{k=0}^{N-1} a_k(x, D)u + \tilde{\mathcal{R}}_N u, \tag{B.10}$$

where, despite the strong singularity of a_k at $\xi = 0$, $a_k(x, D)$ have regular (integrable) Schwartz kernels, and

$$\begin{aligned} \tilde{\mathcal{R}}_N u(x) &= 2 \int_{\mathbb{S}^{n-1}} \int_0^\infty r^N R_N(x, r, \omega) u(x + r\omega) dr d\sigma(\omega) \\ &= 2 \int |x - y|^{N-n+1} R_N\left(x, |x - y|, \frac{x - y}{|x - y|}\right) u(y) dy. \end{aligned} \tag{B.11}$$

We express R_N in its Cauchy form as

$$R_N(x, r, \omega) = \frac{1}{(N - 1)!} \int_0^1 \partial_r^N A_{\text{even},k}(x, tr, \omega) (1 - t)^{N-1} dt.$$

We have

$$|D^\alpha \tilde{\mathcal{R}}_N u| \leq C^N \alpha! \quad \text{in } U' \text{ for } |\alpha| \leq N - s. \tag{B.12}$$

Splitting the sum (B.9) into two parts, we write

$$a(x, D) = \text{Op}\left(\sum_{k=0}^{N-1} \varphi_k(\xi) a_k(x, \xi)\right) + \text{Op}\left(\sum_{k=N}^\infty \varphi_k(\xi) a_k(x, \xi)\right). \tag{B.13}$$

For the second term we have (see (3.15) in Chapter V in [40])

$$\left| D^\alpha \text{Op}\left(\sum_{k=N}^\infty \varphi_k(\xi) a_k(x, \xi)\right) u \right| \leq C^N \alpha! \quad \text{in } U' \text{ for } |\alpha| \leq N - s. \tag{B.14}$$

We are left to compare the first sum in (B.13) with the sum in (B.10):

$$\mathcal{B}_N u := \text{Op}\left(\sum_{k=0}^{N-1} (1 - \varphi_k(\xi)) a_k(x, \xi)\right) u = \mathcal{B}'_N u + \mathcal{B}''_N u,$$

where

$$\mathcal{B}'_N u = \text{Op} \left(\sum_{k=0}^{N-1} (\varphi_N(\xi) - \varphi_k(\xi)) a_k(x, \xi) \right) u,$$

$$\mathcal{B}''_N u = \text{Op} \left(\sum_{k=0}^{N-1} (1 - \varphi_N(\xi)) a_k(x, \xi) \right) u.$$

On $\text{supp}(\varphi_N - \varphi_k)$, we have $2Rk \leq |\xi| \leq 3RN$ provided that $k < N$ and, as always, we assume that $R \gg 1$. Using this and (B.8), we get

$$|D^\alpha \mathcal{B}'_N u| \leq C(CRN)^{|\alpha|-1+s} \quad \text{in } U' \text{ for } |\alpha| \leq N-1, \quad (\text{B.15})$$

compare with (3.17) in Chapter V in [40].

We write $\mathcal{B}''_N u$ in the form (see (B.7))

$$\mathcal{B}''_N u(x) = 2 \sum_{k=0}^{N-1} \int A_{\text{even},k} \left(x, \frac{z}{|z|} \right) |z|^{k-n+1} (1 - \varphi_N) \check{*} f(z+x) dz. \quad (\text{B.16})$$

This implies

$$|D^\alpha \mathcal{B}''_N u| \leq C^N (CRN)^{|\alpha|+s} \quad \text{in } U' \text{ for } |\alpha| \leq N-1. \quad (\text{B.17})$$

Combining (B.12), (B.14), (B.15), and (B.17), we get

$$|D^\alpha (\mathcal{A} - a(x, D))u| \leq C^N N! \quad \text{in } U' \text{ for } |\alpha| \leq N-s.$$

For $N \geq s$, choose $|\alpha| = N-s$ to conclude that $(\mathcal{A} - a(x, D))u$ is analytic in U' . \square

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