

Zonal Jet Creation from Secondary Instability of Drift Waves for Plasma Edge Turbulence

Dedicated to Professor Andrew J. Majda on the occasion of his seventieth birthday

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Abstract A new strategy is presented to explain the creation and persistence of zonal flows widely observed in plasma edge turbulence. The core physics in the edge regime of the magnetic-fusion tokamaks can be described qualitatively by the one-state modified Hasegawa-Mima (MHM) model, which creates enhanced zonal flows and more physically relevant features in comparison with the familiar Charney-Hasegawa-Mima (CHM) model for both plasma and geophysical flows. The generation mechanism of zonal jets is displayed from the secondary instability analysis via nonlinear interactions with a background base state. Strong exponential growth in the zonal modes is induced due to a non-zonal drift wave base state in the MHM model, while stabilizing damping effect is shown with a zonal flow base state. Together with the selective decay effect from the dissipation, the secondary instability offers a complete characterization of the convergence process to the purely zonal structure. Direct numerical simulations with and without dissipation are carried out to confirm the instability theory. It shows clearly the emergence of a dominant zonal flow from pure non-zonal drift waves with small perturbation in the initial configuration. In comparison, the CHM model does not create instability in the zonal modes and usually converges to homogeneous turbulence.

Keywords zonal flow generation · drift wave turbulence · secondary instability · modified Hasegawa-Mima model

1 Introduction

Persistent zonal flows have been widely observed from the nature, experiments, and numerical simulations of various rotating fluids [9, 18, 23, 8, 2, 3]. In fusion plasma, poloidally extended zonal jets in the edge region of magnetically confined tokamak devices are of particular interest where the turbulent transport severely limits plasma confinement and leads to disastrous particle transport towards the boundary regime. The anomalous particle transport along the radial direction due to drift wave turbulence is found to be regulated and suppressed by the generation of poloidal zonal structures [6, 2, 12, 25, 19]. It has been suggested from several theoretical and numerical results [24, 14] that zonal flows are generated spontaneously by interacting with the drift waves. The drift wave in plasma edge turbulence is also analogous to the Rossby wave in geostrophic fluids where similar zonal jet structures are observed [9, 1].

In understanding the drift wave – zonal flow interacting dynamics, it is useful to adopt simplified models where the most relevant physical mechanism is identified. The Hasegawa-Mima (HM) [4, 1] and Hasegawa-Wakatani (HW) [5, 17] models are two groups of the simplified models which are capable to qualitatively capture the energy-conserving nonlinear dynamics for the formation of zonal jets. The HM models contain most essential physical features in the drift wave – zonal flow feedback loop mechanism, while the HW models include a drift wave instability driving the turbulence. Striking new features are generated in a newly developed *flux-balanced Hasegawa-Wakatani* (BHW) model [12, 22], where corrected treatment for the electron responses parallel to the magnetic field lines is introduced as a more physical improvement from the *modified Hasegawa-Wakatani* (MHW) model [5, 17]. One important observation from the BHW model simulations is the enhanced stronger zonal jets persistent in all the dynamical regimes even with high particle resistivity [22, 12]. In contrast, the MHW model lacks the skill to maintain such strong zonal jets and ceases to homogeneous drift wave turbulence at the low resistivity limit.

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In analyzing zonal flows from drift wave turbulence, the BHW model consists of the interplay of the linear drift wave instability and the nonlinear coupling between drift waves and zonal states. The modified Hasegawa-Mima (MHM) model, as the exact adiabatic one-state limit [12] of the BHW model, gives a cleaner setup by filtering out the linear instability, thus offers a more desirable starting model for investigating the central mechanism in flow self-organization from drift waves to coherent zonal states through nonlinear interactions. The MHM model is modified from the original Charney-Hasegawa-Mima (CHM) model [4] for plasma and geophysical flows which is also known as the quasi-geostrophic model [1,9]. Modulational instability of drift waves offers a feedback mechanism for the generation of zonal flows through the nonlinear interactions. Theories and numerical experiments have been attempted [14,24,15] for describing the emergence of zonal flows by Reynolds stress in both CHM and MHM models.

In this paper, we provide a precise explanation for the underlying mechanism in creating the dominant zonal jets observed in the flux-balanced models using secondary instability analysis about a background base state of drift wave solutions. To identify the important nonlinear impact between interactions of the drift wave states and the zonal modes, we stay in the simple one-state Hasegawa-Mima (HM) models at the adiabatic limit of the two-state BHW model, where no internal instability due to the particle resistivity is included to add extra complexity in the flow turbulence. The generation and persistence of zonal flows in the HMH model is investigated by demonstrating that: first a non-zonal drift wave base state induces strong instability in the zonal modes, implying nonlinear energy transfer to the zonal states; and then the generated zonal structure as a base state stays stable to perturbations thus is maintained in time as the system evolves. The secondary instability results are first illustrated by numerical computation of the largest growth exponent from the Floquet theory. Further, we use direct numerical simulations to confirm the jet creation mechanism. Zonal flows are induced from a pure drift wave state adding small isotropic fluctuations in the MHM model even without any dissipation effect. In the case with dissipation, selective decay principle developed in [21] helps to work together with the secondary instability mechanism to drive the state to a final purely zonal structure. In contrast in the CHM model, none of these instability and zonal jets are created due to the improper treatment in the electron flux response.

In the structure of this paper, we first briefly describe the BHM and MHM models with a balanced averaged flux creating strong zonal jets. Section 2 introduces the basic MHM model properties with its major physical interpretation. The exact single mode drift wave solution as well as the zonal mean dynamics is derived in Section 3 for the background base mode in generating the zonal states. The precise energy transfer mechanism to the zonal modes is explained through the secondary instability about the background state with numerical computations of the growth rate in Section 4. Section 5 uses direction numerical simulations with and without dissipation effects for confirming the developed theories. The conclusion and further discussion are given in the final Section 6.

1.1 The flux balanced models for plasma edge turbulence

In tokamak devices, the realistic geometry would be a circular domain with a predominant magnetic field \mathbf{B} along the toroidal z -direction. However, the shape of the plasma edge can be approximated on a slab geometry under a Cartesian coordinate where the toroidal magnetic surfaces are embedded. The Hasegawa-Wakatani models describe the drift wave – zonal flow interactions of a two state coupled system on the 2D slab geometry [12,1], with x -axis corresponding to the radial direction and y -axis representing the poloidal direction. The flux-balanced Hasegawa-Wakatani (BHW) model is introduced in [12] based on the flux-balanced potential vorticity $q = \nabla^2 \varphi - \tilde{n}$ and the density fluctuation n in the following form

$$\frac{\partial q}{\partial t} + \nabla^\perp \varphi \cdot \nabla q - \kappa \frac{\partial \varphi}{\partial y} = D \Delta q, \quad q = \nabla^2 \varphi - \tilde{n}, \quad (1a)$$

$$\frac{\partial n}{\partial t} + \nabla^\perp \varphi \cdot \nabla n + \kappa \frac{\partial \varphi}{\partial y} = \alpha (\tilde{\varphi} - \tilde{n}) + D \Delta n, \quad (1b)$$

where φ is the electrostatic potential, n is the density fluctuation from background density $n_0(x)$, and $\mathbf{u} \equiv \nabla^\perp \varphi = (-\partial_y \varphi, \partial_x \varphi)$ is the velocity field. The parameter α is for adiabatic resistivity of parallel electrons. It determines the degree to which electrons can move rapidly along the magnetic field lines. The constant background density gradient $\kappa = -\nabla \ln n_0$ is defined by the exponential background density profile near the boundary $n_0(x)$. D acts on the two states with the Laplace operator as a homogeneous damping [22,12]. The physical quantities φ and n are decomposed into zonal mean states $\bar{\varphi}, \bar{n}$ and their fluctuations about the mean $\tilde{\varphi}, \tilde{n}$ so that

$$\varphi = \bar{\varphi} + \tilde{\varphi}, \quad n = \bar{n} + \tilde{n}, \quad \bar{f}(x) = L_y^{-1} \int f(x, y) dy.$$

In the BHW model, the poloidally averaged density \bar{n} along y -direction is removed from the potential vorticity q . In contrast, the original Hasegawa-Wakatani model introduced in [5] as well as the modified version (MHW) [17] uses the ‘unbalanced’ potential density $q = \nabla^2 \varphi - n$ without removing the mean state \bar{n} in the potential vorticity, leading to problems with the convergence at the adiabatic limit $\alpha \rightarrow \infty$.

92 The BHW model offers a more realistic formulation with several desirable properties. Most importantly, it is
 93 shown from rigorous proof and numerical confirmation [12,22] that at the adiabatic limit, $\alpha \rightarrow \infty$, the BHW model
 94 converges to the following equation

$$\frac{\partial q}{\partial t} + \nabla^\perp \varphi \cdot \nabla q - \kappa \frac{\partial \varphi}{\partial y} = D \Delta q, \quad q = \nabla^2 \varphi - \tilde{\varphi}, \quad (2)$$

95 which is called the modified Hasegawa-Mima model. Notice the modification by removing zonal state in $\tilde{\varphi}$ in the
 96 definition of potential vorticity q above. On the other hand, the MHW model shows performance significantly
 97 different from the MHM model when $\alpha \rightarrow \infty$. The strong zonal jets created from the BHW and MHM model and
 98 the convergence at the adiabatic limit are discussed with explicit numerical simulations in [12] (see Fig. 4 and
 99 5 there). If we replace the potential vorticity in (2) by $q = \nabla^2 \varphi - \varphi$ without removing the zonal mean state, it
 100 recovers the Charney-Hasegawa-Mima model. The CHM model is identical to the quasi-geostrophic model with F -
 101 plane effect describing geophysical turbulence with rotation and stratification [9,18,20]. Then the rigorous theories
 102 developed for the geophysical model apply to the CHM model exactly in the same way. In this paper, we will focus
 103 on the HM models and especially changes in MHM model due to the averaged flux correction in order to analyze
 104 the unstable effect purely from a background base flow.

105 2 The Hasegawa-Mima models and their representing properties

106 To offer a better illustration with physical interpretations in the Hasegawa-Mima models, we start with the original
 107 dimensional formulation with physically related variables and derive the non-dimensionalized version using the
 108 physical scales. The *Charney-Hasegawa-Mima* (CHM) equation and the *modified Hasegawa-Mima* (MHM) equation
 109 can be formulated under the same framework by defining a switch parameter with $s = 0$ for CHM and $s = 1$ for
 110 MHM as

$$\frac{D}{Dt} \left(\frac{\zeta}{\omega_{ci}} + \ln \frac{\omega_{ci}}{n_0} - \frac{e}{T_e} (\tilde{\varphi} + \delta_{s0} \bar{\varphi}) \right) = 0, \quad (3)$$

111 where φ is the electrostatic potential, $\zeta = \nabla^2 \varphi / B_0$ is the vorticity, $\mathbf{v}_E = -\nabla \varphi \times \hat{z} / B_0$ is the $\mathbf{E} \times \mathbf{B}$ velocity.
 112 $D/Dt \equiv \partial_t + \mathbf{v}_E \cdot \nabla$ represents the material derivative along the velocity. In the parameters, T_e is the reference
 113 electron temperature, $\omega_{ci} = eB_0/m_i$ is the ion cyclotron frequency, and m_i is the ion mass [1]. For model non-
 114 dimensionalization, the new variables are introduced by

$$e\varphi/T_e \rightarrow \varphi, \quad \omega_{ci}t \rightarrow t, \quad (x, y)/\rho_s \rightarrow (x, y),$$

115 with $\rho_s = \omega_{ci}^{-1} (T_e/m_i)^{1/2} = \sqrt{m_i T_e} / eB_0$ the characteristic length scale of drift waves and ω_{ci}^{-1} the characteristic
 116 time scale from the ion frequency. Accordingly, we find the non-dimensional velocity and vorticity

$$\rho_s \frac{eB_0}{T_e} \mathbf{v}_E \rightarrow \mathbf{v}_E = \nabla^\perp \varphi, \quad \rho_s^2 \frac{eB_0}{T_e} \zeta \rightarrow \zeta.$$

117 By substituting the non-dimensionalized quantities back into the dimensional equation (3), we can rewrite the
 118 original (with $s = 0$) and modified (with $s = 1$) Hasegawa-Mima equations in the non-dimensional form as in (2)
 119 so that

$$\left(\frac{\partial}{\partial t} + \nabla^\perp \varphi \cdot \nabla \right) q + (\partial_x \ln n_0) \frac{\partial}{\partial y} \tilde{\varphi} = 0, \quad q = \zeta - (\tilde{\varphi} + \delta_{s0} \bar{\varphi}). \quad (4)$$

120 Above we introduce the new variable q as the potential vorticity, and if we assume a constant exponential decay
 121 profile in the background density $n_0 \sim \exp(-\kappa x)$ the coefficient becomes a constant $\kappa \equiv -\partial_x \ln n_0$.

122 On the magnetic surfaces, the electrons are assumed to respond adiabatically so that locally thermodynamical
 123 equilibrium (with Boltzmann distribution) is achieved on a given field surface. The electron density fluctuation
 124 does not respond adiabatically on the averaged part of the electrostatic potential $\bar{\varphi}$, thus only the flux balanced
 125 component $e\tilde{\varphi}/T_e$ follows the Boltzmann distribution. This offers the intuition for removing the zonal mean state
 126 $\bar{\varphi}$ in the MHM model. Though simple enough, the modified expansion leads to much stronger zonal jet structures
 127 and more physically consistent performance [1,14] compared with the CHM results.

128 2.1 Galilean invariance and model energetics

129 We illustrate some representative features especially from the model flux modification. First, the MHM model
 130 enhances the excitation of zonal flows with more prominent zonal structures. Consider a single mode plane wave
 131 $\varphi = A_z(x, t) \exp(i(\mathbf{k} \cdot \mathbf{x} - \omega t))$ decomposed into a slowly varying zonal mean and fast fluctuation. The slow mode

132 A_z is assumed to be zonal and gives a constant zonal mean flow $\mathbf{v}_E = \bar{v}\hat{y}$. We can find the linearized dispersion
133 relations for CHM and MHM separately as

$$\begin{aligned} \text{CHM: } \omega &= \frac{k_y \kappa}{1+k^2} + \frac{k^2}{1+k^2} k_y \bar{v}, \\ \text{MHM: } \omega &= \frac{k_y \kappa}{1+k^2} + k_y \bar{v}. \end{aligned}$$

134 Without the mean flow \mathbf{v}_E , the HM models generate no instability with the same dispersion relation in the first
135 term on the right side. In small scales $k \gg 1$, the CHM and MHM models have similar dispersion relations. In
136 large scales $k \lesssim 1$ (that is, near the scale of ρ_s), the modified model gets a stronger feedback from the fluctuation
137 (due to the simple Doppler shift $\mathbf{k} \cdot \mathbf{v}_E = k_y \bar{v}$). In the unmodified model, the Doppler shift is reduced by a factor
138 $k^2/(1+k^2)$. Detailed discussions about mean flow interaction in the CHM model can be found in [9].

139 Second, the MHM model is Galilean invariant under boosts in the y (poloidal) direction as desired for the
140 symmetry in the poloidal direction of tokamak devices. If we introduce a poloidal boost V in the flow, the new
141 states become

$$y' = y - Vt, \quad \varphi' = \varphi - VB_0x.$$

142 Notice that only the fluctuation in the electrostatic potential is invariant, $\tilde{\varphi}' = \tilde{\varphi}$, while the zonal mean is not
143 invariant, $\bar{\varphi}' = \bar{\varphi} - VB_0x$, under the change of coordinate. The CHM model (and also the QG model in geophysics)
144 does not maintain this invariance due to the last term $e\bar{\varphi}/T_e$ with $s = 0$.

145 At last, we describe the model energetics. In the MHM model, two important conserved quantities [21,17] can
146 be found as the energy E and the enstrophy W

$$E = \frac{1}{2} \int \tilde{\varphi}^2 + |\nabla \varphi|^2, \quad W = \frac{1}{2} \int q^2 = \frac{1}{2} \int (\tilde{\varphi} - \nabla^2 \varphi)^2. \quad (5)$$

147 The nonlinear term in (4) does not alter the value of both energy and enstrophy. Thus the evolution of energy and
148 enstrophy can be purely determined by the dissipation effects. Especially with the homogeneous damping form
149 $D\Delta q$ in (2), we can derive the dynamical equations

$$\begin{aligned} \frac{dE}{dt} &= -D \int |\nabla \tilde{\varphi}|^2 + |\nabla^2 \varphi|^2, \\ \frac{dW}{dt} &= -D \int |\nabla q|^2 = -D \int |\nabla \tilde{\varphi}|^2 + 2|\nabla^2 \varphi|^2 + |\nabla^3 \varphi|^2. \end{aligned}$$

150 Similarly, the CHM model also maintains two invariants with $\tilde{\varphi}$ in the definition (5) and equations replaced by
151 φ . The energetic equations play important roles in showing the stability and decay properties. In particular, a
152 selective decay to a single dominant mode can be discovered based on the energetics [13,21].

153 2.2 Selective decay in the flux balanced model

154 The persistence of the zonal jets in the MHM model can be first explained in a rigorous mathematical approach
155 using the selective decay principle [21,13]. It states that proper dissipation operator can dissipate all the non-zero
156 drift wave states at a much faster rate except a single selected dominant zonal state in the MHM model. Precisely
157 speaking, we have the convergence to one of the selective decay zonal states $\bar{\varphi}_k$ for the normalized potential
158 function in the H^1 sense

$$\lim_{t \rightarrow \infty} \|\nabla \phi - \nabla \bar{\varphi}_k\|_0 = 0, \quad \phi = \frac{\varphi}{\|\nabla \varphi\|_0}. \quad (6)$$

159 In the CHM model, the selective decay state φ_k in a single wavenumber can be also reached under the dissipation
160 operator, while the final converged state is one drift wave mode without zonal structure. Proof for the selective decay
161 results using different dissipation operators including the Landau damping with detailed numerical simulations are
162 shown in [21]. Still, the generation of the zonal structures from any arbitrary initial states is directly related with
163 the nonlinear interaction mechanism between different modes before the selective decay effect takes over.

164 3 Exact drift wave solutions and the zonal mean dynamics

165 Now we propose the precise model framework for analyzing the instability, creation and stabilization of zonal
166 jets through the nonlinear interacting mechanism with the background base states. First, we introduce additional
167 rescaling for the HM models so that the important parameters that determine the solution structures are identified.
168 Starting with the previous model formulation (4)

$$\frac{\partial q}{\partial t} + \nabla^\perp \varphi \cdot \nabla q - \kappa \frac{\partial \varphi}{\partial y} = D\Delta q, \quad q = \nabla^2 \varphi - (\tilde{\varphi} + \delta_{s0} \bar{\varphi}),$$

169 with $s = 1$ for the MHM model and $s = 0$ for the CHM model, we propose the rescaled set of variables $(q', \varphi', \mathbf{x}', t')$
 170 based on the characteristic length scale L and the characteristic flow velocity scale U

$$\mathbf{x} = L\mathbf{x}', \quad \mathbf{u} = \nabla^\perp \varphi = U\mathbf{u}', \quad t = Tt', \quad \varphi = \Phi\varphi', \quad q = Qq'.$$

171 The scales of the other variables can be found based on the values of L, U as

$$T = \frac{L}{U}, \quad \Phi = UL, \quad Q = \frac{\Phi}{L^2} = \frac{U}{L}.$$

172 With the above rescaling, the unit wavenumber mode \mathbf{p} , $|\mathbf{p}| = 1$ for the new state represents the inverse length scale
 173 L^{-1} , and the flow state with unit amplitude $u = \exp(\mathbf{p} \cdot \mathbf{x})$ represents the velocity with strength U . Accordingly,
 174 we derive the rescaled HM models for the normalized states $(q', \varphi', \mathbf{x}', t')$ based on the proposed characteristic
 175 scales

$$\frac{\partial q'}{\partial t'} + \nabla_{\mathbf{x}'}^\perp \varphi' \cdot \nabla_{\mathbf{x}'} q' - \kappa' \frac{\partial \varphi'}{\partial y'} = D' \Delta_{\mathbf{x}'} q', \quad q' = \nabla_{\mathbf{x}'}^2 \varphi' - L^2 (\hat{\varphi}' + \delta_{s0} \bar{\varphi}'). \quad (7)$$

176 The above rescaled equation (7) is not changed much just with new non-dimensional parameters κ', D' . Notice
 177 that the length scale L now appears explicitly in the potential vorticity q' . The flow solution is entirely determined
 178 by the two characteristic coefficients, $\kappa' = \frac{\kappa L^2}{U}$ and $D' = \frac{D}{UL}$. κ' has the same role as the Rhines number Rh^{-1} in
 179 geophysical flows, showing the anisotropic effect in the drift waves; and D' as the Reynolds number Re^{-1} for the
 180 dissipation effect [23, 18]. We will focus on the MHM model with $s = 1$ in (7) and neglect the primes on the states
 181 in the rest part of the paper (the CHM case can be easily implied and detailed theories for the CHM model have
 182 actually been developed in the geophysical literatures [9–11]).

183 3.1 General base flow state from the exact solution

184 Exact drift wave solutions of the MHM equation in (7) can be found by considering a single mode base state. We
 185 assume the base mode in the electrostatic potential and the potential vorticity for a single wavenumber $\mathbf{k} = (k^x, k^y)$
 186 as

$$\varphi(\mathbf{x}, t) = \hat{\varphi} \exp(i(\mathbf{k} \cdot \mathbf{x} - \omega(\mathbf{k})t)), \quad q(\mathbf{x}, t) = -\left[k^2 + L^2(1 - \delta_{k^y, 0})\right] \hat{\varphi} \exp(i(\mathbf{k} \cdot \mathbf{x} - \omega(\mathbf{k})t)), \quad (8)$$

187 where $\omega(\mathbf{k})$ is the dispersion relation in the drift waves. The Kronecker delta operator is introduced for the MHM
 188 model modification in the zonal modes $k^y = 0$. The normalized wavenumber length $k = |\mathbf{k}|$ is compared with
 189 the characteristic scale L , i.e., wavenumbers $k < 1$ characterize the scales larger than L and wavenumbers $k > 1$
 190 for scales smaller than L . Since we only consider a single mode solution in the above form, the nonlinear term
 191 $\nabla^\perp \varphi \cdot \nabla q$ vanishes in the equation through the self interaction $\mathbf{k}^\perp \cdot \mathbf{k} = 0$. The dispersion relation can be found as

$$\omega(\mathbf{k}) = \kappa' \frac{k^y}{k^2 + L^2}, \quad \kappa' = \frac{\kappa L^2}{U}. \quad (9)$$

192 Notice that the above dispersion relation $\omega(\mathbf{k})$ is valid for both the MHM and CHM models. In fact, the MHM
 193 model only adds modifications for the zonal modes with $k^y = 0$. In the zonal modes, the dispersion relation becomes
 194 $\omega = \kappa' k^y / k^2 = 0$. The formula (9) is still valid for both the MHM ($s = 1$) case and CHM ($s = 0$) case. Next, we
 195 will consider the instability of fluctuation perturbations added on top of the representative exact solutions in the
 196 form of (8).

197 3.2 Dynamical equation of the zonal mean state

198 Before proceeding to the detailed discussion about secondary instability to induce zonal structures, it is useful to
 199 check the exact dynamical equation for the zonal state to achieve a first intuition about the nonlinear interacting
 200 mechanism. Evolutions of the zonal components \bar{q} can be extracted from the above MHM model by directly taking
 201 the zonal average about the original equation (7)

$$\frac{\partial \bar{q}}{\partial t} + \frac{\partial}{\partial x} \bar{u} \bar{q} = -D' \frac{\partial^4 \bar{q}}{\partial x^4}, \quad u = -\partial_y \varphi, \quad (10)$$

202 with u the zonal velocity fluctuation. The background density gradient term $\kappa' \partial_y \varphi$ vanishes after the average along
 203 y -direction. If there is no non-zero zonal mode $k^y = 0$, we can check that the advection term vanishes

$$\bar{u} \bar{q} = \overline{(-ik^y \hat{\varphi} e^{i(\mathbf{k} \cdot \mathbf{x} - \omega t)}) (- (k^2 + L^2) \hat{\varphi} e^{i(\mathbf{k} \cdot \mathbf{x} - \omega t)})} + c.c. = \overline{ik^y (k^2 + L^2) \hat{\varphi}^2 e^{2i(\mathbf{k} \cdot \mathbf{x} - \omega t)}} + c.c. = 0,$$

after the integration along the y -direction, with *c.c.* as the complex conjugate part. Therefore single non-zonal fluctuation mode makes no contribution to the zonal mean structure in the above form, where an exact solution (8) can be reached.

On the other hand, zonal wave could be generated through the interactions between different wavenumbers. If we consider a general solution with multiple drift wave modes, the mean state equation can be derived in the following form

$$\left(\frac{d}{dt} + Dk^2\right) \bar{\varphi}_{\mathbf{k}}(t) = k^{-2} \sum_{\mathbf{m}+\mathbf{n}=\mathbf{k}} C_{\mathbf{k}\mathbf{m}\mathbf{n}}(t) (n^2 - m^2) \overline{\varphi_{\mathbf{m}}\varphi_{\mathbf{n}}},$$

with the zonal mode $\mathbf{k} = (k^x, 0)$ and the fluctuation feedback $\overline{\varphi_{\mathbf{m}}\varphi_{\mathbf{n}}}$, $m^y \neq 0, n^y \neq 0$ to the zonal state. The coupling parameter $C_{\mathbf{k}\mathbf{m}\mathbf{n}}(t)$ is time-dependent on the dispersion relations (9) and models the triad coupling between the interacting drift wave modes

$$C_{\mathbf{k}\mathbf{m}\mathbf{n}}(t) = k^x n^x e^{i(\omega(\mathbf{k}) - \omega(\mathbf{m}) - \omega(\mathbf{n}))t}, \quad m^x + n^x = k^x, \quad m^y + n^y = 0.$$

The right hand side of the above equation describes the nonlinear flux to the mean mode for the generation of a zonal jet. In the next section, we illustrate in a rigorous way how this nonlinear coupling term transfers the fluctuating energy in the non-zonal drift wave modes to the zonal directions, and maintains the zonal structures through the secondary stability mechanism.

4 Secondary instability from a base flow state

In this section, we provide a precise description for the energy transfer mechanism from the drift wave states (with $k^y \neq 0$) to zonal flows (with $k^y = 0$). From the discussion in the last section, energy in the fluctuation drift wave modes is first transferred to the zonal directions through the resonant triad interactions; next the accumulation of energy in the zonal modes gets saturated and stabilized with the large-scale stability of a zonal base state. The drift wave – zonal flow interactions are characterized by the secondary instability analysis based on a base flow state. A brief summary for the main result achieved for the MHM model is: a fluctuation drift wave base state will induce strong instability along a wide range of zonal modes, implying strong transport of energy from non-zero drift waves to the zonal directions; in contrast a purely zonal flow base state will add no instability to zonal modes or the drift waves, showing stability in the generated zonal mean structure.

4.1 Formulation of the secondary instability from a base state

First notice that drift wave instability is filtered out in the one-state Hasegawa-Mima models (7), which enables us to focus on the nonlinear interaction mechanism from a background state. Below we derive the secondary instability based on the MHM model (the CHM case can be derived in a similar fashion). The development is motivated by the secondary instability analysis carried out in [7] for geophysical turbulence on beta-plane and in [16] for 2D Navier-Stokes equations with a Kolmogorov base flow using *Floquet theory*. However, the main focus here is the changes introduced through the flux modification in the MHM model potential vorticity $q = \nabla^2 \varphi - \tilde{\varphi}$.

For simplicity in the MHM model, we consider a single mode base state with wavevector $\mathbf{p} = (\cos \theta_p, \sin \theta_p)$ of unit length (then θ_p defines the characteristic direction of the base flow with $\theta_p = 0$ for the zonal flow state and $\theta_p = \frac{\pi}{2}$ for the pure drift wave state). The single mode base flow potential Φ_p and vorticity $Q_p = \nabla^2 \Phi_p - L^2 \tilde{\Phi}_p$ with the unit length wavenumber \mathbf{p} can be defined from the exact solution formula (8) as

$$\Phi_p = -\frac{1}{2} e^{i(\mathbf{p} \cdot \mathbf{x} - \omega(\mathbf{p})t)} - \frac{1}{2} e^{-i(\mathbf{p} \cdot \mathbf{x} - \omega(\mathbf{p})t)}, \quad Q_p = -\left[1 + L^2(1 - \delta_{p^y, 0})\right] \Phi_p, \quad (11)$$

with the dispersion relation $\omega(\mathbf{p}) = \kappa' \frac{p^y}{1+L^2}$ defined in (9). From the rescaled equation (7) using the characteristic scales (L, U) , the base solution $U_p = \nabla^\perp \Phi_p$ with unit wavenumber $p = 1$ represents the characteristic length scale L and the characteristic flow velocity U . For the corresponding CHM model solution, we just need to remove the delta functions in the above formula. With no additional internal instability in the HM models, the above solution can be simply generated by a combined forcing and damping effect.

The Floquet theory considers a fluctuation solution with a characteristic multiplier $e^{\mu t}$. To study the secondary instability at each wavenumber \mathbf{k} based on the single-mode base flow (11), we introduce fluctuations, $\varphi_p = \varphi - \Phi_p$, $q_p = q - Q_p$, on top of the base state in the form

$$\begin{aligned} \varphi_p &= e^{\mu t} e^{i(\mathbf{k} \cdot \mathbf{x} - \omega(\mathbf{k})t)} \sum_{l=-N}^N \hat{\varphi}_l e^{il(\mathbf{p} \cdot \mathbf{x} - \omega(\mathbf{p})t)}, \\ q_p &= e^{\mu t} e^{i(\mathbf{k} \cdot \mathbf{x} - \omega(\mathbf{k})t)} \sum_{l=-N}^N -\left[q^2(l) + L^2(1 - \delta_{q^y, 0})\right] \hat{\varphi}_l e^{il(\mathbf{p} \cdot \mathbf{x} - \omega(\mathbf{p})t)}. \end{aligned} \quad (12)$$

246 The perturbation is added on the wavenumber \mathbf{k} with dispersion relation $\omega(\mathbf{k})$. μ is the *Floquet exponent* characterizing the growth or decay rate in this perturbed mode. $\mathbf{q}(l) = \mathbf{k} + l\mathbf{p}$ with $q^2(l) = |\mathbf{q}(l)|^2$ is the full wavenumber
 247 with the perturbation. The perturbation mode $\hat{\varphi}_l$ is added on each of the directions as multiples of the characteristic
 248 wavenumber $l\mathbf{p}$. It refers to a multiplicative perturbation along the directions of the base state flow. Here we
 249 truncate the perturbed states within the leading modes up to N multiples of the base state. Thus the perturbation
 250 coefficients $\{\hat{\varphi}_l\}$ form a finite dimensional system of $2N + 1$ real states. The problem can be further extended to an
 251 infinite dimensional system as the truncation size $N \rightarrow \infty$. Still, since we are mostly interested in the instability
 252 among the largest scales in the zonal states, a finite size truncation N will be sufficient for characterizing this
 253 instability (see the instability results illustrated in Section 4.2).
 254

255 By subtracting the base flow solution (Φ_p, Q_p) from the MHM equation (7), the fluctuation equation for the
 256 perturbed component q_p of potential vorticity can be derived in the following form

$$\frac{\partial q_p}{\partial t} + \nabla^\perp \Phi_p \cdot \nabla \left[q_p + \varphi_p + L^2 (1 - \delta_{p^y,0}) \varphi_p \right] + \nabla^\perp \varphi_p \cdot \nabla q_p - \kappa' \frac{\partial \varphi_p}{\partial y} = D' \Delta q_p, \quad q_p = \nabla^2 \varphi_p - \tilde{\varphi}_p,$$

257 where the relation for the single mode base state vorticity and potential, $Q_p = -[1 + L^2(1 - \delta_{p^y,0})]\Phi_p$, is used.
 258 Still in the secondary stability analysis, we focus on the secondary instability induced by the interactions between
 259 the background base state (Φ_p, Q_p) and the fluctuation modes (φ_p, q_p) with small perturbations. The higher-order
 260 nonlinear term between fluctuation modes, $\nabla^\perp \varphi_p \cdot \nabla q_p$, is assumed to stay small in the starting transient state and
 261 is neglected in this instability analysis. The unstable growth in perturbations due to the background base state is
 262 represented by the growth parameter μ in the fluctuation states (12). With a positive value in the real part of μ ,
 263 it infers exponential growth of the perturbation mode in the transient state on top of the base mode due to the
 264 interactions. The equation for calculating the secondary growth μ can be achieved by substituting the fluctuation
 265 modes (12) into the above fluctuation equation for q_p . It becomes an eigenvalue problem for each perturbation
 266 direction \mathbf{k} individually with interactions between the triad neighboring modes $(\hat{\varphi}_{l-1}, \hat{\varphi}_l, \hat{\varphi}_{l+1})$

$$\begin{aligned} & \left[\mu(\mathbf{k}) - i\omega(\mathbf{k}) - il\omega(\mathbf{p}) + i\omega(\mathbf{q}(l)) + Dq^2(l) \right] \hat{\varphi}_l \\ & + \frac{1}{2q^2(l)} (\mathbf{p} \times \mathbf{k}) \cdot \hat{z} \left[q^2(l-1) - 1 - L^2(1 - \delta_{p^y,0}) \right] \hat{\varphi}_{l-1} = 0, \quad l = -N, \dots, N. \\ & - \frac{1}{2q^2(l)} (\mathbf{p} \times \mathbf{k}) \cdot \hat{z} \left[q^2(l+1) - 1 - L^2(1 - \delta_{p^y,0}) \right] \hat{\varphi}_{l+1} \end{aligned} \quad (13)$$

267 with the combined wavenumber $\mathbf{q}(l) = \mathbf{k} + l\mathbf{p}$, $q^2(l) = |\mathbf{q}(l)|^2$, and the coupling coefficient $(\mathbf{p} \times \mathbf{k}) \cdot \hat{z} = p^x k^y - p^y k^x$.
 268 The first row above includes the effects from the dispersion relation and damping, and the second and third rows
 269 are due to the interactions with the neighboring modes through the background state. The equations in (13) form
 270 a $(2N + 1) \times (2N + 1)$ tri-diagonal system (with N the number of base mode perturbations added) based on the
 271 perturbed modes $\hat{\varphi}_l$ for each wavenumber \mathbf{k} . The solution of μ reflecting instability can be achieved by computing
 272 the eigenvalues of the corresponding tri-diagonal matrix. The maximum positive eigenvalue in real part of $\mu^{\max}(\mathbf{k})$
 273 refers to the most unstable growth rate in the mode \mathbf{k} according to the base flow in direction \mathbf{p} .
 274

275 As a comment for the general case, the base flow can be generalized to a combination of the single drift wave
 276 solutions (8) with a group of characteristic wavevectors $\{\mathbf{p}_j\}_{j=1}^J$. The base states for the electrostatic potential Φ
 277 and the potential vorticity $Q = \nabla^2 \Phi - L^2 \tilde{\Phi}$ then can be defined as the combination of all the modes

$$\Phi(\mathbf{x}, t; \{\mathbf{p}_j\}_{j=1}^J) = \sum_{j=1}^J A_j e^{i(\mathbf{p}_j \cdot \mathbf{x} - \omega(\mathbf{p}_j)t)}, \quad Q(\mathbf{x}, t; \{\mathbf{p}_j\}_{j=1}^J) = - \sum_{j=1}^J A_j \left[p_j^2 + L^2(1 - \delta_{p_j^y,0}) \right] e^{i(\mathbf{p}_j \cdot \mathbf{x} - \omega(\mathbf{p}_j)t)}.$$

278 Above J is the total number of characteristic modes that are combined in the background base state Φ . Accordingly
 279 for the fluctuation state about the combined base solution, we need to combine perturbations on mode \mathbf{k} for each
 280 component of the base state in the form

$$\varphi(\{\mathbf{p}_j\}) = e^{\mu t} e^{i(\mathbf{k} \cdot \mathbf{x} - \omega(\mathbf{k})t)} \cdot \sum_{\mathbf{l}} \hat{\varphi}_{\mathbf{l}} \exp \left[\sum_{j=1}^J il_j (\mathbf{p}_j \cdot \mathbf{x} - \omega(\mathbf{p}_j)t) \right],$$

281 where the index $\mathbf{l} = (l_1, l_2, \dots, l_J) \in \mathbb{Z}^J$ goes through all the J -multiples in the summation. Similar result can be
 282 derived just with much more complicated formulas. We will leave the multiple base mode case in future investigation
 283 and focus on the central issue about the generation of zonal jets through a single drift wave mode.

284 4.2 Secondary instability about a single mode base state

285 Now we check the secondary growth rate μ according to different types of the base flows through simple numerical
 286 tests. Especially, we consider the single-mode drift wave base state $\mathbf{p}_1 = (0, 1)$ and the zonal flow base state
 287 $\mathbf{p}_2 = (1, 0)$. The number of perturbed modes in the fluctuation state (12) is fixed at $N = 20$ (thus it forms
 288 a 41×41 matrix). Larger truncation sizes of N have been checked and show no significant difference for the
 289 instability in the zonal modes.

290 From the scale analysis, we find that the flow solutions can be determined by the two non-dimensional pa-
 291 rameters, $\kappa' = \frac{\kappa L^2}{U}$ and $D' = \frac{D}{UL}$. Especially, L represents the characteristic scale in the base mode \mathbf{p} and U
 292 defines the strength in the base flow. In the numerical tests, the strategy is that we fix the model parameters
 293 as $\kappa = 0.5, D = 5 \times 10^{-4}$, and change the scale parameters L and U . In general, we choose a computational
 294 domain size $L_D = 40$ used in the direct numerical simulations in Section 5. The characteristic length scale L
 295 represents a drift wave state with wavenumber $s = L_D/L$. We tested two different length scales $L = 10, 20$ and
 296 two velocity scales $U = 1, 0.1$. Correspondingly, it gives the non-dimensional parameters $\kappa' = 50, 200, 500$ and
 297 $D' = 5 \times 10^{-5}, 2.5 \times 10^{-5}, 5 \times 10^{-4}$ in the three test cases shown in Figure 1.

298 4.2.1 Instability about drift wave mode $\mathbf{p} = (0, 1)$

299 In the first test case, we consider the secondary instability due to a single-mode drift wave state $\mathbf{p} = (0, 1)$ varying
 300 only along the k^y -direction. In this case, it illustrates the transfer of energy from the purely drift wave modes
 301 to the zonal jet states through the nonlinear interactions between the background state and the fluctuations.
 302 Especially, from the drift wave linear instability in the two-state HW model, the most linearly unstable modes are
 303 always along the k^y axis (see Fig. 2 in [22] for the linear instability result). Therefore, starting from the HM model
 304 framework, the pure drift wave background state represents the first excited states due to the (unresolved) linear
 305 instability effect. We investigate the secondary instability induced due to this background state from linear drift
 306 wave instability.

307 In Figure 1, we plot the contours for the maximum growth rate in the real part of the exponent μ with different
 308 wavenumbers \mathbf{k} in the spectral domain. Results with balanced vorticity $q = \nabla^2 \varphi - L^2 \tilde{\varphi}$ in the MHM model and
 309 original vorticity $q = \nabla^2 \varphi - L^2 \varphi$ in the CHM model are compared. With drift wave mode $(0, 1)$ as the basic flow,
 310 strong positive growth rate is generated in the large-scale zonal modes with $k^y = 0$ in the MHM model uniformly
 311 for all the tested model scales L, U . All the largest positive growth rates are located near the zonal direction. This
 312 corresponds to the rapid energy transfer from the drift waves to form up zonal structures. In small scales the real
 313 parts of the eigenvalues become negative due to the much stronger damping in the smaller scales. In contrast,
 314 the CHM model without the flux modification gets no instability but only has negative decaying effect in the real
 315 part of μ . This shows the inability of creating zonal structures of the CHM model. Direct numerical simulations
 316 starting from drift waves will be shown in Section 5 for an explicit illustration of the difference between the two
 317 models.

318 For a more detailed comparison about the model instability changing with characteristic scales, the maximum
 319 growth with different parameter values L and U are compared. As shown in the three rows of Figure 1, larger
 320 characteristic scale $L = 20$ in the drift waves induces larger number of unstable zonal modes in higher wavenumbers
 321 and stronger growth rate. In comparison, the CHM model results have little change with only negative eigenvalues
 322 for stability. Figure 2 shows the maximum growth rate μ depending on different model scales U and L from
 323 secondary stability analysis along the zonal modes $k^y = 0$. More clearly, the MHM model gets unstable zonal
 324 modes from the drift wave state, while the CHM model has no instability at all along the zonal direction. For the
 325 MHM model, the largest growth gets saturated at large values of flow amplitude U ; and with decreasing amplitudes
 326 of U , the growth rate drops slowly and will finally vanish at the extremely small value $U < 0.01$ where the effective
 327 dissipation $D' = \frac{D}{UL}$ becomes strong. On the other hand for the CHM model, at most weak instability is induced
 328 for large value of U around the largest scales. The positive growth rate quickly vanishes as the flow strength U
 329 decreases in value. This is also reflected in the contour plots in Figure 1.

330 Finally, we test the small characteristic length $L = 0.01$, that is, the background drift wave state is in a
 331 very small scale. At this small length scale limit, both MHM and CHM models converge to the barotropic model
 332 with infinite (or large) deformation frequency, $q \rightarrow \nabla^2 \varphi$ [9, 10]. Figure 3 shows the maximum secondary growth
 333 rate from both the MHM and CHM model results. As expected, the growth rates from the two models perform
 334 similarly at this limit where the balanced flux correction for $\tilde{\varphi}$ becomes negligible. Also the maximum growth of
 335 the perturbations takes place at the largest scales $k < 1$ compared with the small characteristic scale L in the
 336 background drift wave state. Besides, the maximum secondary growth stays in small amplitude with just weak
 337 instability among all the wavenumbers. This agrees with the results in [7] for barotropic turbulence.

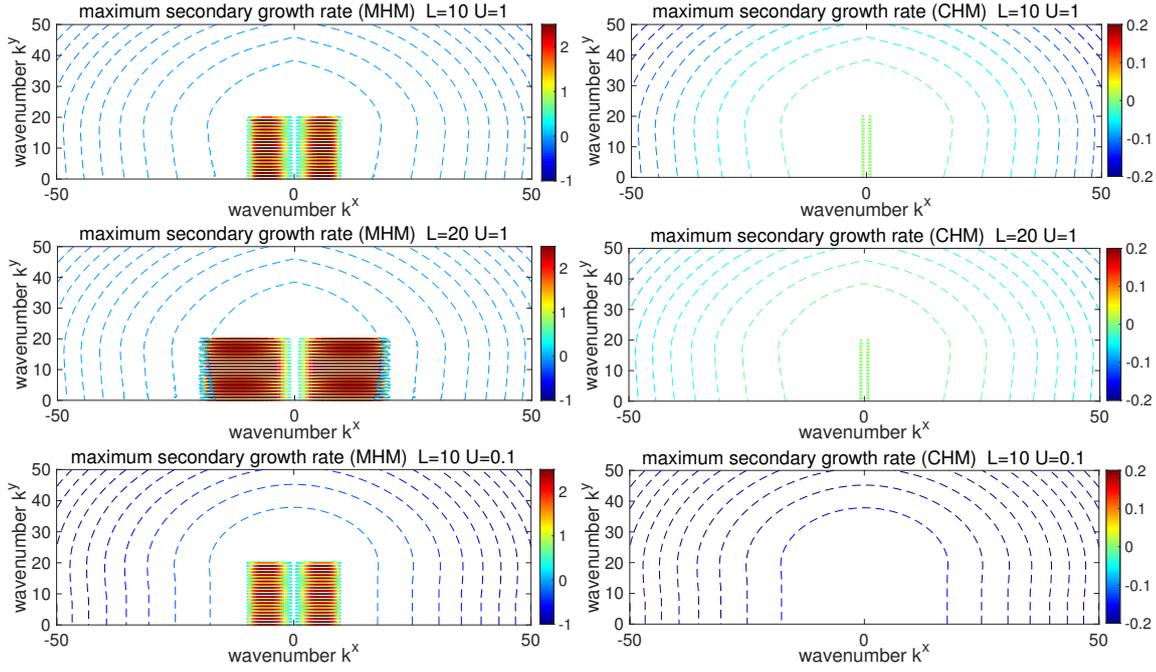


Fig. 1: Maximum growth rate at each spectral mode from the secondary stability analysis according to the drift wave base flow $\mathbf{p} = (0, 1)$. Solid lines are for positive growth rates and dashed lines are for negative damping rates. Results for the MHM model (left) and for the CHM model (right) in same parameter domains are compared. Different characteristic scales for (L, U) are compared. The other parameters used are $\kappa = 0.5, D = 5 \times 10^{-4}$. Notice the large amplitudes in the MHM model and small values in the CHM model from the colorbars.

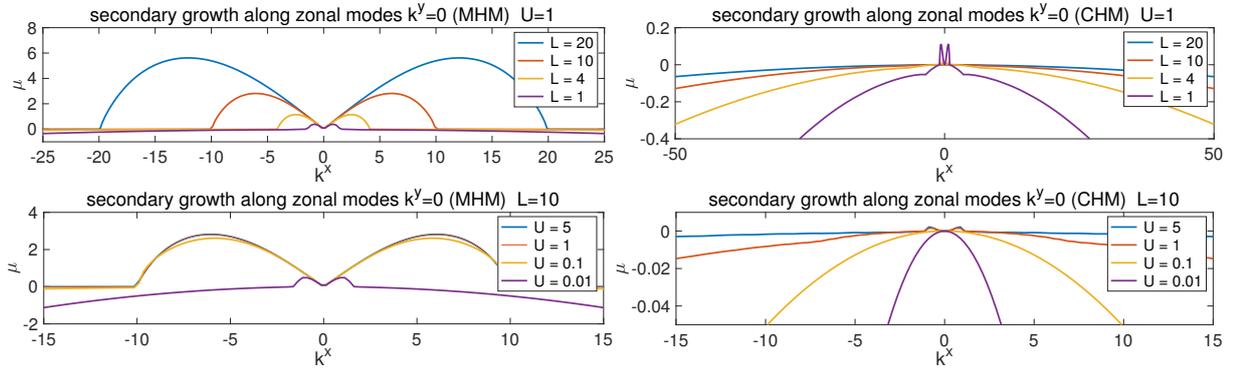


Fig. 2: Maximum growth rate from secondary stability analysis along the zonal mode direction with $k^y = 0$ for the MHM and CHM models according to the drift wave base flow $\mathbf{p} = (0, 1)$. Results with different characteristic length scale L and background flow strength U are compared.

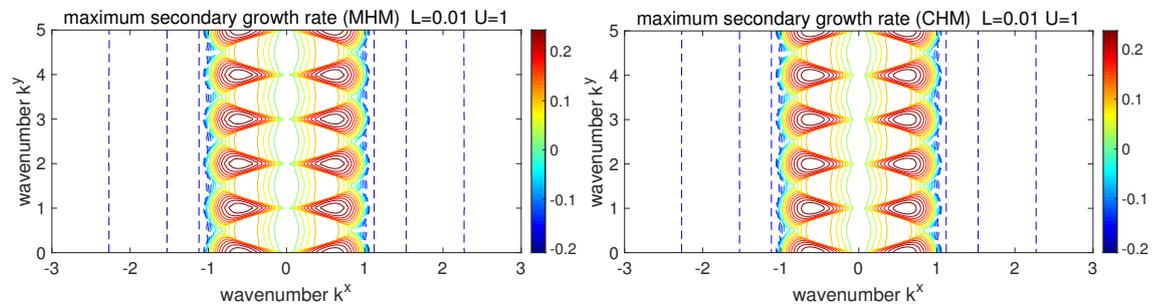


Fig. 3: Real part of the secondary growth rate μ according to background base drift wave flow $\mathbf{p} = (0, 1)$ with small characteristic length $L = 0.01$ and $U = 1$. Solid lines are for positive growth rates and dashed lines are for negative damping rates. Results for the MHM model (left) and for the CHM model (right) are compared. The other parameters used are $\kappa = 0.5, D = 5 \times 10^{-4}$.

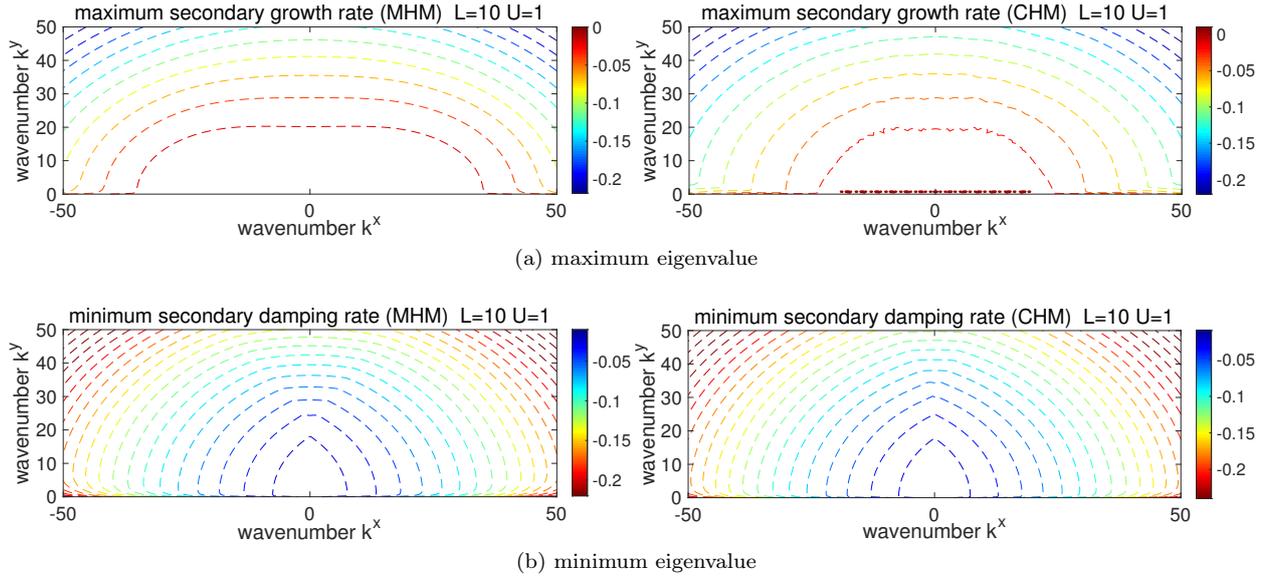


Fig. 4: Maximum growth rates and minimum damping rates for each spectral mode from largest and smallest eigenvalues in the secondary stability analysis according to zonal jet base flow $\mathbf{p} = (0, 1)$. All the eigenvalues are negative in the MHM model shown in dashed lines, while the CHM model has positive growth rates near the zonal axis $k^y = 0$. Results for the MHM model (left) and for the CHM model (right) are compared. The parameters used are $L = 10$, $U = 1$, and $\kappa = 0.5$, $D = 5 \times 10^{-4}$.

338 4.2.2 Stability about zonal flow mode $\mathbf{p} = (1, 0)$

339 In this second test case, we consider the secondary instability due to a background zonal jet state with wave
 340 direction $\mathbf{p} = (1, 0)$. In this case, a positive growth rate along the zonal modes implies further growth of the
 341 fluctuations in zonal states until higher order nonlinear interactions between modes take over. Usually, if the zonal
 342 jet amplitude keeps growing, it will finally break down due to the nonlinear interactions and cascade to the smaller
 343 scales to get dissipated. On the other hand, no instability in the zonal directions implies that the zonal jet structure
 344 is maintained stable since perturbations in zonal modes will not grow and jet structure will persist in time. This
 345 case with a zonal flow base state then characterizes the stability of the zonal structures, which can be created by
 346 the instability from drift waves through the instability analysis shown in Section 4.2.1.

347 In Figure 4, the maximum and minimum eigenvalues from the secondary stability analysis according to zonal
 348 jet base flow are plotted. It can be observed that there exists no positive growth rate in μ for instability at all
 349 throughout the spectral regime due to zonal flows in the MHM model. This supports the intuition described above
 350 that the induced zonal jets can be maintained stable in time in response to additional wave perturbations. In the
 351 CHM model case, the growth rate is in a similar shape but gets small positive growth near the zonal direction
 352 $k^y = 0$. This instability in the zonal modes may imply the less stable zonal jets and the possible break down of the
 353 zonal structures due to perturbations in the CHW model. In additional, we also compare the minimum eigenvalue
 354 for the strongest damping rate in each mode from the stability analysis. The zonal modes get stronger damping at
 355 smaller scales. This again confirms the stability of the zonal jets, so that small perturbations in the zonal direction
 356 will be quickly damped down from the stabilizing effect in the background base mode.

357

358 Combining the conclusions of instability in the drift wave state and stability in the zonal modes, we can draw
 359 a complete picture about the energy mechanism due to the nonlinear transfer of energy in the transient state. By
 360 adopting the Hasegawa-Mima model (2) without drift wave linear instability, we start with a drift wave state that
 361 could be generated from the drift wave turbulence in the higher level Hasegawa-Wakatani model (1). The secondary
 362 instability in the drift wave base mode induces strong growth particularly along the zonal mode direction, which
 363 infers the strong transport of energy from the drift wave modes to the zonal states. After the formation of the
 364 zonal structure, the strong negative damping with no instability about the zonal jet background state shows the
 365 persistence of the zonal structure to perturbations. The zonal jets will emergence even without the help of the
 366 selective decay in dissipating small scale fluctuations described in [21].

5 Direct numerical simulations to confirm the generation of zonal jets

In this final part, we use direct numerical simulations of the MHM and CHM models (4) to confirm the theory from the secondary instability induced by the background base mode discussed in the previous section. A pseudo-spectral code with a 3/2-rule for de-aliasing the nonlinear term [21,20] is applied on the square domain with side length $L_D = 40$ and resolution $N = 256$. For the time integration, a 4th-order Runge-Kutta scheme is adopted. Small time integration step $\Delta t = 1 \times 10^{-4}$ is taken in all the simulations to ensure the conservation properties especially for the non-dissipative case. The same model parameter values are taken as in Section 4.2 for instability analysis.

To check the energy transfer mechanism from drift waves, the initial state of the simulations is set as a pure drift wave adding homogenous perturbations

$$\varphi_0 = \frac{L_D}{s} \cos\left(\frac{2\pi s}{L_D} y\right) + \epsilon \sum_{|\mathbf{k}| \leq \Lambda} k^{-2} \hat{\xi}_{\mathbf{k}} e^{i\mathbf{k} \cdot \mathbf{x}}, \quad (14)$$

with $\hat{\xi}_{\mathbf{k}} \sim \mathcal{N}(0, 1)$ sampled independently from the standard normal distribution. In practice, we add the perturbations up to wavenumber $\Lambda = 5$ and the perturbation amplitude is set in a small value $\epsilon = 0.01$. The parameter s determines the scale of the background drift wave. We test two values with $s = 2$ representing drift waves of two wavelengths and $s = 10$ representing drift waves of wavenumber 10 (see the electrostatic potential φ_0 shown in Figure 6 for these two initial states). Besides, we consider two different situations without and with the dissipation operator $D\Delta q$ in the model.

5.1 Time evolution of energy and enstrophy in full and zonal modes

In the first numerical simulations, we introduce no dissipation effect $D = 0$ in the model. Thus the conservation of total kinetic energy E and potential enstrophy W defined in (5) should be guaranteed. We run the model in this way so that the selective decay effect [21] will be excluded. Therefore if zonal structures are generated in the final steady state from the simple drift wave initial state (14), the mechanism can only be the secondary interactions between the initial background state φ_0 and the perturbed small modes.

We need to confirm in the first place that the numerical dissipations have little effect in changing the model energy and enstrophy and offer no contribution in the final state of the model. For checking the conservation in the model simulations, the first two rows of Figure 5 plot the time-series of the total energy E and enstrophy W from the direct model simulations. In both the MHM and CHM model results, the total energy and enstrophy are conserved in time with at most small decrease in the enstrophy due to the numerical dissipation strongest at the smallest scales. Further we plot the energy and enstrophy only contained in the zonal state $\bar{\varphi}$ and \bar{q} . The ratio of energy in zonal velocity $v^2 / (u^2 + v^2)$ is used to characterize the flow structure, which reaches 1 when the purely zonal flow is reached. In the MHM model, the zonal energy and enstrophy start near zero in the initial time due to the initial setup, then the secondary instability takes over and the zonal energy and enstrophy jump to a large non-zero value through the nonlinear interaction. This infers the strong instability from drift wave modes and stability in the zonal jet states. In contrast, the zonal energy and enstrophy in the CHM stay in small values near zero throughout the time evolution. Then no zonal state is excited in the CHM model from the nonlinear effect.

In the last two rows of Figure 5, the time-series of energy and enstrophy as well as the zonal energy ratio with a small dissipation $D = 5 \times 10^{-4}$ are plotted. In comparison with the non-dissipative case before, the energy and enstrophy are no longer conserved. Especially, the enstrophy decays in a faster rate than the energy, implying that the dissipation is stronger on damping the smaller scale modes. Still, the MHM case induces strong zonal structures as the system approaches the final state. The CHM model still lacks the skill in generating zonal jets, while a pure single drift wave mode is converged consistent with the selective decay principle [21,13].

Most importantly, observe that in the time-series of enstrophy for the MHM model, the zonal enstrophy starts to rise at $t = 2$ (marked by dashed line in the figure) while the total enstrophy begins to drop at a later time at $t = 5$ (marked by dotted-dashed line). This illustrates the competition between the secondary instability and selective decay: i) during the starting time $t < 2$, the initial state maintains with no linear instability; ii) between the time $2 < t < 5$, the secondary instability comes into effect to generate a strong zonal structure while the dissipation has no obvious effect on the smaller scale modes; iii) finally after time $t > 5$, the selective decay becomes dominant and strongly dissipates the smaller scale fluctuations while maintains the created zonal jet.

5.2 Convergence to the final steady state without dissipation

We check the explicit flow structures from the direct numerical simulations of the MHM and CHM models. Figure 6 plots the snapshots of the electrostatic potential functions φ from both the MHM and CHM model simulations

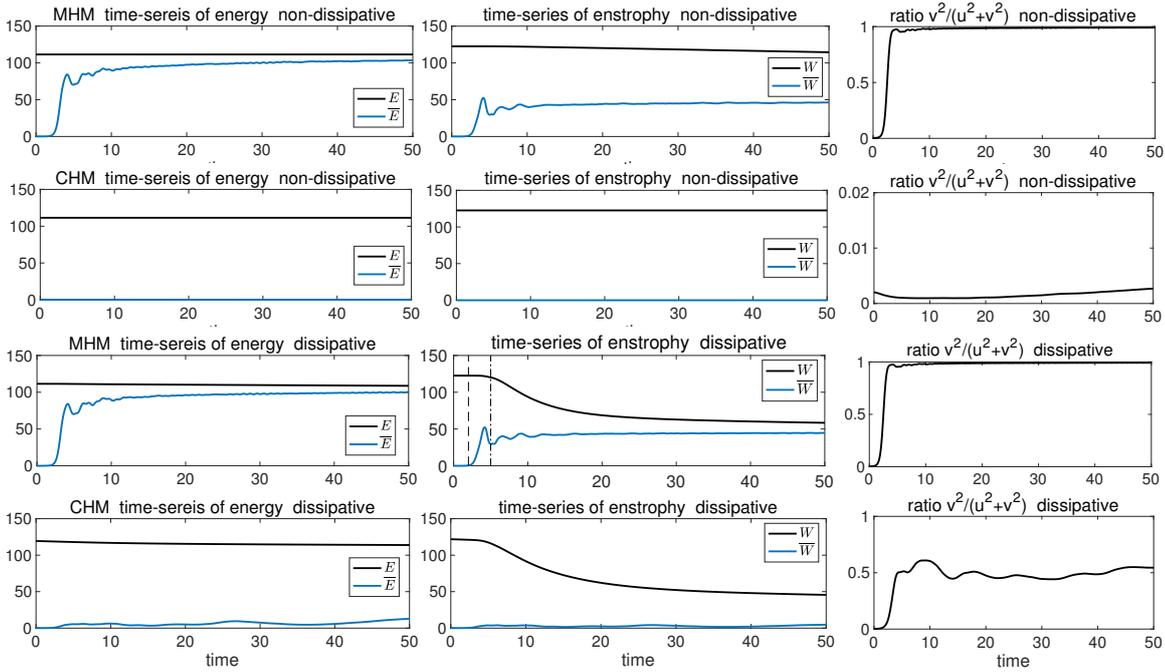


Fig. 5: Time-series of the total energy E and total enstrophy W as well as the energy and enstrophy contained in the zonal state $(\bar{\varphi}, \bar{q})$ from the MHM and CHM model simulations. The first two rows show the results for MHM and CHM models without dissipation effect $D = 0$, and the last two rows are the results for both models including weak dissipation $D = 5 \times 10^{-4}$. The last column plots the ratio of energy in the zonal state $v^2/(u^2 + v^2)$.

417 starting from the same initial drift wave structure (14). First in the MHM model in the second column, starting
 418 from the pure drift wave state with only small isotropic perturbations (first column of Figure 6), the solutions
 419 always generate strong zonal jets in the end. This confirms the transfer of energy through the secondary instability
 420 shown in Figure 1 since no dissipation and other effect are included to generate the zonal structures. Also since
 421 there is no dissipation, the small scale non-zonal fluctuations always exist in the system on top of the zonal jets.

422 In comparison, the CHM model shown in the third column of Figure 6 has difficulty in generating the zonal
 423 structures. In the first case with a large scale wavenumber of two, the drift wave structure is maintained in time as
 424 the system evolves. This is consistent with the secondary stability result in the drift wave case where no positive
 425 growth rate is observed in the CHM model with a large scale base flow (see the right column of Figure 1). In
 426 the second test case with smaller scale drift wave state of wavenumber ten initially, finally the pure drift wave
 427 structure is destroyed due to the relatively stronger instability with smaller scale drift waves. Still the flow breaks
 428 into homogeneous drift wave turbulence without any zonal jet structure. This confirms the little instability in the
 429 zonal modes (only in the largest scales) in the CHM model case.

430 5.3 Combined effects with secondary instability and selective decay

431 In the previous test cases, we run the models without dissipation effect. As described in [21], the damping operator
 432 usually adds stronger selective effect on the non-zonal modes and drives the flow solution to a single wavenumber
 433 state at the long time limit. Thus in this final test case, we consider the combined contributions from both the
 434 secondary instability and the selective damping. For simplicity, we introduce the simple dissipation operator, $D\Delta q$,
 435 as in (2) for both MHM and CHM models. The damping rate is kept in small value $D = 5 \times 10^{-4}$. In the last two
 436 rows of Figure 5, we already show the time-series of total energy and enstrophy in this case with dissipation. Both
 437 energy and enstrophy are no longer conserved in time. Still the energy/enstrophy in the zonal mean goes to the
 438 same level as the total energy/enstrophy at the long time limit in the MHM case, while the CHM only gets little
 439 energy in the zonal state.

440 Again, we check the final dissipated solutions from the direct numerical simulations. We plot in Figure 7 the
 441 snapshots of the electrostatic potential φ at the final simulation time starting from the same two initial states
 442 with different drift wave scales with the inclusion of dissipation effect. Comparing with the the non-dissipative
 443 results in Figure 6, the fluctuating small-scale structures are damped down in this case while the dominant zonal
 444 structure is still maintained in the MHM model. This is the typical selective decay solution shown in Fig. 3 of
 445 [21], while here the detailed energy exchanging mechanism is discovered by the secondary instability. Observe that

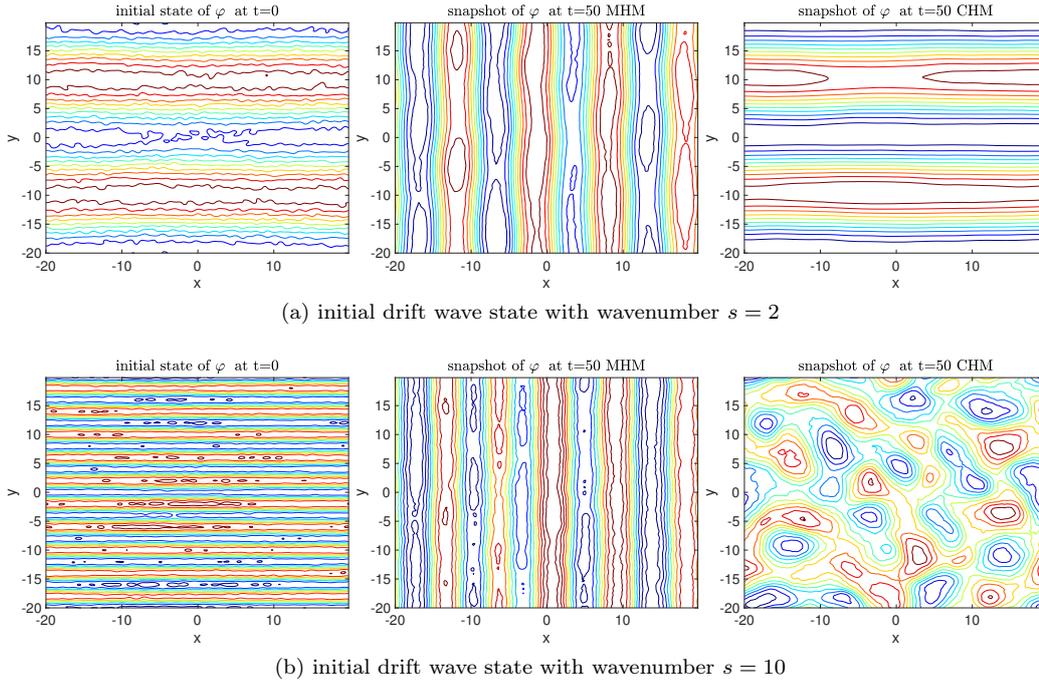


Fig. 6: Snapshots of the electrostatic potential function φ in the initial state (left) and the final numerical states from MHM (middle) and CHM (right) model simulations without dissipation $D = 0$. Two different initial drift wave states with $s = 2$ wavelengths (upper) and with $s = 10$ wavelengths (lower) are compared.

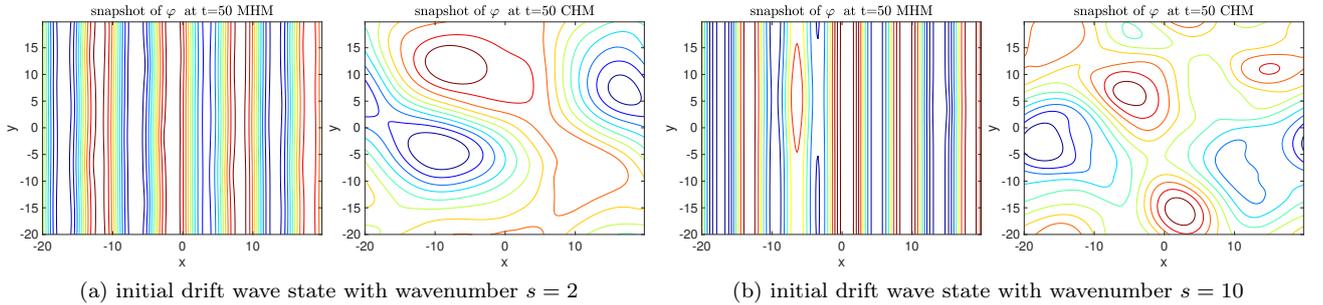


Fig. 7: Snapshots of the electrostatic potential function φ in the final states from MHM and CHM model simulations with dissipation $D = 5 \times 10^{-4}$. The same two initial drift wave states with $s = 2$ and $s = 10$ are compared.

446 same number of zonal jets in the two cases is reached as in the non-dissipative results. This shows (together with
 447 the time-series of enstrophy in Figure 5) that the instability determines the final zonal structure in the first place,
 448 then the selective decay takes over to dissipate all the other non-zonal fluctuation modes to reach a clean single
 449 mode zonal state. In comparison, the CHM model converges to a drift wave selective decay state without zonal
 450 flows. With long enough time, the CHM flows will finally converge to a single drift wave mode. More detailed
 451 results about selective decay in the CHM model can be found in Fig. 1 of [21] and [13].

452 5.3.1 Energy transfer mechanism in the decaying process

453 We offer a more detailed illustration about the energy transfer mechanism in time between different modes in the
 454 MHM model through comparing the energy spectra. In Figure 8, we plot the energy spectra in the two initial cases
 455 with and without dissipation from the MHM model simulation results. To display the transfer of energy from the
 456 non-zonal drift wave modes to the zonal modes, we compare the spectra in radially averaged modes in the first row
 457 and in zonal modes $k^y = 0$ only in the second row. The initial spectra get a dominant second or tenth wavenumber
 458 from the initial setup (14) with small fluctuations and a high wavenumber truncation. The energy will gradually
 459 cascade to smaller scales in the transient state. A dominant zonal mode with largest energy emerges finally. With
 460 dissipation, the selective damping effect only strongly dissipates the smaller scale modes. The dominant zonal

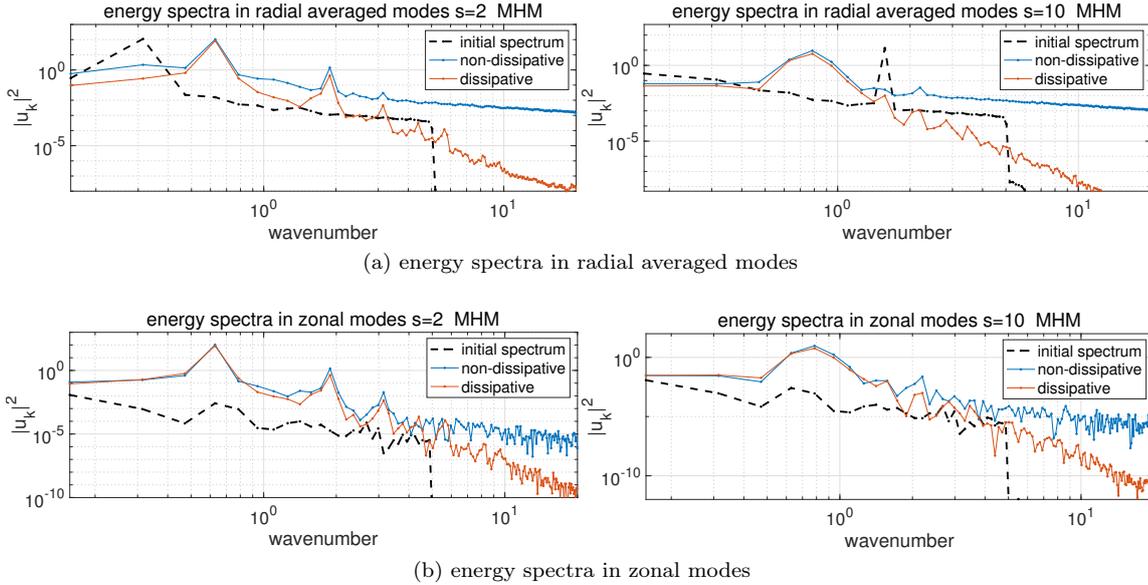


Fig. 8: Energy spectra in radial averaged modes (upper) and purely zonal modes (lower) from the MHM model simulations. The initial energy spectra in the two test cases are shown together with the final spectra achieved with and without dissipation effect.

461 mode gets maintained at exactly the same wavenumber as the non-dissipative case and stays with large energy for
 462 all the time.

463 Finally, to offer a complete picture about the creation of pure zonal jet structure through the combined effects
 464 of secondary instability and the selective damping, we plot the normalized energy ratio for the zonal modes $k^y = 0$
 465 and the non-zonal fluctuation modes at several time instant in Figure 9. In the initial state (shown in dashed black
 466 lines), all the energy is contained in the pure drift wave mode with wavenumber two ($s = 2$, left) or wavenumber
 467 ten ($s = 10$, right). As the starting transient state (at around time $t = 3$, see also the time-series of energy and
 468 enstrophy in Figure 5), the energy in the zonal modes $k^y = 0$ begins to grow due to the secondary instability
 469 induced by the interactions between the drift waves and zonal modes. At later time (starts from time $t = 5$), the
 470 energy in the non-zonal drift wave modes begins to cascade to smaller scales and gets dissipated by the selective
 471 damping. In accordance with the time-series of energy plotted in Figure 5, the energy in the zonal modes grows
 472 rapidly between the time window $t \in [3, 5]$. Then the selective damping effect takes over to drive the state to
 473 purely zonal jets. In addition, it can be observed from the energy ratios in the zonal modes that there exist several
 474 intermediate metastable saddle points which the solution visited before the convergence to the final stable single
 475 selective decay zonal mode (see [21] for a complete description of the selective decay).

476 6 Concluding discussion

477 In this paper, we perform secondary instability analysis about a background base state to explain the zonal
 478 jet creation mechanism generally observed in plasma edge turbulence. The one-state modified Hasegawa-Mima
 479 model without internal drift wave instability is adopted to identify the central drift wave – zonal flow nonlinear
 480 interactions, and the results are compared with the Charney-Hasegawa-Mima model. Together with the selective
 481 decay principle developed previously in [21], a complete picture for the generation and persistence of a dominate
 482 zonal jet structure can be drawn. Starting from a drift wave base state created from the first linear drift wave
 483 instability, secondary instability due to nonlinear coupling with the fluctuation modes gradually takes over and
 484 transfers the energy in the non-zonal drift wave states to the zonal states. The induced zonal mode as a background
 485 state is further stabilized from the negative secondary damping effect from interacting with the perturbations. The
 486 small scale fluctuations from the initial state are maintained if no dissipation exists in the system, otherwise the
 487 selective decay effect will strongly dissipate the smaller scale modes while it does not alter the dominant zonal
 488 structure created from the instability. Direct numerical simulations of the MHM model display the creation of
 489 zonal flows from a pure non-zonal drift wave state with only small perturbation and without the effect of selective
 490 decay. When dissipation is also added, secondary instability is effective before the selective decay to generate the
 491 same number of zonal jets, and the selective decay effect finally drives the state to a clean single mode zonal jet
 492 structure. In contrast, the CHM model cannot create zonal flows automatically and has no instability along the
 493 zonal model direction.

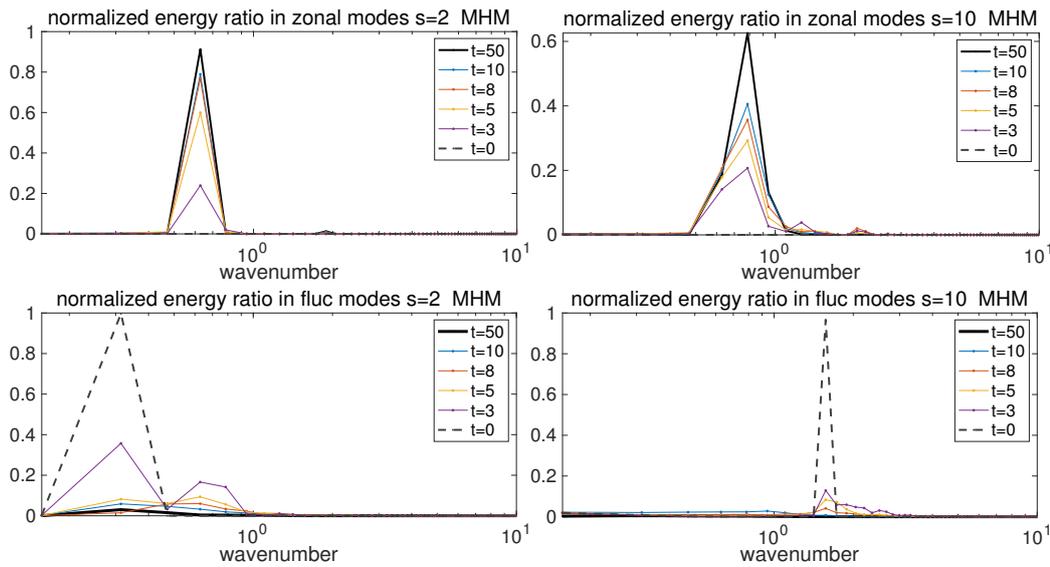


Fig. 9: Selective decay of the fluctuation modes to pure zonal modes in the MHM model from the energy ratios captured at several representative time instant during the model evolution. The total energy is normalized to one to emphasize the portion of energy in each mode.

494 Here we focus on the main energy mechanism for the creation of zonal structures, thus a single mode base
 495 mode is always used throughout the paper in illustrating the central role of secondary instability. As an immediate
 496 generalization, the secondary instability with combined effects with multiple background modes can be investigated.
 497 The multiple background base modes should show stronger dominant exponential growth along the zonal direction
 498 since different base modes enforce the instability in the zonal modes together and have cancellation effect in the
 499 non-zonal directions. As a further generalization, it is useful to consider the instability in the two-state Hasegawa-
 500 Wakatani models. There we need to consider the first linear instability in the base mode together with the secondary
 501 instability on top of the stable/unstable base modes. Especially, it is interesting to investigate the regime with
 502 large values of adiabatic resistivity α , where the model is on its way to approach the MHM model discussed here.

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505 References

- 506 1. Dewar, R.L., Abdullatif, R.F.: Zonal flow generation by modulational instability. In: *Frontiers in Turbulence and Coherent*
 507 *Structures*, pp. 415–430. World Scientific (2007)
- 508 2. Diamond, P.H., Itoh, S., Itoh, K., Hahm, T.: Zonal flows in plasma—a review. *Plasma Physics and Controlled Fusion* **47**(5),
 509 R35 (2005)
- 510 3. Fujisawa, A.: A review of zonal flow experiments. *Nuclear Fusion* **49**(1), 013001 (2009). URL [http://stacks.iop.org/0029-](http://stacks.iop.org/0029-5515/49/i=1/a=013001)
 511 [5515/49/i=1/a=013001](http://stacks.iop.org/0029-5515/49/i=1/a=013001)
- 512 4. Hasegawa, A., Mima, K.: Pseudo-three-dimensional turbulence in magnetized nonuniform plasma. *The Physics of Fluids* **21**(1),
 513 87–92 (1978)
- 514 5. Hasegawa, A., Wakatani, M.: Plasma edge turbulence. *Physical Review Letters* **50**(9), 682 (1983)
- 515 6. Horton, W.: Drift waves and transport. *Rev. Mod. Phys.* **71**, 735–778 (1999). DOI 10.1103/RevModPhys.71.735. URL
 516 <https://link.aps.org/doi/10.1103/RevModPhys.71.735>
- 517 7. Lee, Y., Smith, L.M.: Stability of rossby waves in the β -plane approximation. *Physica D: Nonlinear Phenomena* **179**(1-2),
 518 53–91 (2003)
- 519 8. Lin, Z., Hahm, T.S., Lee, W.W., Tang, W.M., White, R.B.: Turbulent transport reduction by zonal flows: Mas-
 520 sively parallel simulations. *Science* **281**(5384), 1835–1837 (1998). DOI 10.1126/science.281.5384.1835. URL
 521 <http://science.sciencemag.org/content/281/5384/1835>
- 522 9. Majda, A.: *Introduction to PDEs and Waves for the Atmosphere and Ocean*, vol. 9. American Mathematical Soc. (2003)
- 523 10. Majda, A.J.: *Introduction to turbulent dynamical systems in complex systems*. Springer (2016)
- 524 11. Majda, A.J., Qi, D.: Strategies for reduced-order models for predicting the statistical responses and uncertainty quantification
 525 in complex turbulent dynamical systems. *SIAM Review* **60**(3), 491–549 (2018)
- 526 12. Majda, A.J., Qi, D., Cerfon, A.J.: A flux-balanced fluid model for collisional plasma edge turbulence: Model derivation and basic
 527 physical features. *Physics of Plasmas* **25**(10), 102307 (2018). DOI 10.1063/1.5049389. URL <https://doi.org/10.1063/1.5049389>
- 528 13. Majda, A.J., Shim, S.Y., Wang, X.: Selective decay for geophysical flows. *Methods and applications of analysis* **7**(3), 511–554
 529 (2000)
- 530 14. Manfredi, G., Roach, C., Dendy, R.: Zonal flow and streamer generation in drift turbulence. *Plasma physics and controlled*
 531 *fusion* **43**(6), 825 (2001)

- 532 15. Manz, P., Ramisch, M., Stroth, U.: Physical mechanism behind zonal-flow generation in drift-wave turbulence. *Phys. Rev. Lett.*
533 **103**, 165004 (2009). DOI 10.1103/PhysRevLett.103.165004. URL <https://link.aps.org/doi/10.1103/PhysRevLett.103.165004>
- 534 16. Meshalkin, L.: Investigation of the stability of a stationary solution of a system of equations for the plane movement of an
535 incompressible viscous liquid. *J. Appl. Math. Mech.* **25**, 1700–1705 (1962)
- 536 17. Numata, R., Ball, R., Dewar, R.L.: Bifurcation in electrostatic resistive drift wave turbulence. *Physics of Plasmas* **14**(10),
537 102312 (2007)
- 538 18. Pedlosky, J.: *Geophysical fluid dynamics*. Springer Science & Business Media (2013)
- 539 19. Pushkarev, A.V., Bos, W.J.T., Nazarenko, S.V.: Zonal flow generation and its feedback on turbulence production in drift wave
540 turbulence. *Physics of Plasmas* **20**(4), 042304 (2013). DOI 10.1063/1.4802187. URL <https://doi.org/10.1063/1.4802187>
- 541 20. Qi, D., Majda, A.J.: Low-dimensional reduced-order models for statistical response and uncertainty quantification: Two-layer
542 baroclinic turbulence. *Journal of the Atmospheric Sciences* **73**(12), 4609–4639 (2016)
- 543 21. Qi, D., Majda, A.J.: Transient metastability and selective decay for the coherent zonal structures in plasma edge turbulence.
544 submitted to *Journal of Nonlinear Science* (2018)
- 545 22. Qi, D., Majda, A.J., Cerfon, A.J.: A Flux-Balanced Fluid Model for Collisional Plasma Edge Turbulence: Numerical Simulations
546 with Different Aspect Ratios. submitted to *Physics of Plasmas* arXiv:1812.00131 (2018)
- 547 23. Rhines, P.B.: Waves and turbulence on a beta-plane. *Journal of Fluid Mechanics* **69**(3), 417–443 (1975)
- 548 24. Smolyakov, A., Diamond, P., Malkov, M.: Coherent structure phenomena in drift wave–zonal flow turbulence. *Physical review*
549 *letters* **84**(3), 491 (2000)
- 550 25. Xanthopoulos, P., Mischchenko, A., Helander, P., Sugama, H., Watanabe, T.H.: Zonal flow dynamics and control of tur-
551 bulent transport in stellarators. *Phys. Rev. Lett.* **107**, 245002 (2011). DOI 10.1103/PhysRevLett.107.245002. URL
552 <https://link.aps.org/doi/10.1103/PhysRevLett.107.245002>