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## Applications of Stokes' theorem and Divergence Theorem.

### Conservation Laws.

Let  $\vec{v}$  be a  $C^1$  vector field, where now  $\vec{v}(t, x, y, z)$  may depend upon time as well as position. Let  $\rho(t, x, y, z)$  be the mass density of a fluid having velocity  $\vec{v}$ . Conservation of mass says that, for any region  $\Omega$ , the rate of change of total mass in  $\Omega$  equals the rate at which mass flows in to  $\Omega$ .

$$\frac{d}{dt} \iiint_{\Omega} \rho \, dx \, dy \, dz = - \iint_{\partial\Omega} \vec{J} \cdot \vec{n} \, dS .$$

$$\vec{J} = \rho \vec{v}$$

$$\iiint_{\Omega} \frac{\partial \rho}{\partial t} \, dx \, dy \, dz = - \iiint_{\Omega} \operatorname{div} \vec{J} \, dx \, dy \, dz .$$

$$\iiint_{\Omega} \left( \frac{\partial \rho}{\partial t} + \operatorname{div} \vec{J} \right) dx \, dy \, dz = 0 .$$

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Since this is true for every  $\Omega$ , it must be that the integrand is 0; that is,

$$\boxed{\frac{\partial \phi}{\partial t} + \operatorname{div}(\rho \vec{v}) = 0}$$

← one of  
the Euler's  
equation

In the special case when  $\phi = \text{constant}$ ,  $\frac{\partial \phi}{\partial t} = 0$ , we have  $\operatorname{div}(\rho \vec{v}) = 0$ , again, since  $\phi$  is constant, the equation simplifies to:

$$\boxed{\operatorname{div} \vec{v} = 0} \quad (\text{incompressible fluid})$$

Finally, suppose that besides  $\phi = \text{constant}$ , the flow is irrotational; i.e.,  $\nabla \times \vec{v} = 0$ .

$$\nabla \times \vec{v} = 0 \Rightarrow \vec{v} = \nabla \psi.$$

Thus,  $\operatorname{div} \vec{v} = 0$  simplifies to:

$$\operatorname{div}(\nabla \psi) = 0$$

$$\therefore \Delta \psi = 0,$$

Hence,  $\psi$  is a harmonic function.

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### Heat equation

Let  $T(t, x, y, z)$  be a  $C^2$  function giving temperature. Then  $\nabla \vec{T}$  points in the direction of increasing temperature, so the vector giving heat flow would be

$$\vec{F} = -\nabla T$$

If  $\rho_0$  (constant) is the mass density and  $c$  is the specific heat, then  $c\rho_0 T$  gives the energy per unit volume, and the energy flux is  $\vec{J} = k \vec{F}$ , where  $k$  is the conductivity.

Using conservation of energy as with conservation of mass, for arbitrary  $\Omega$ ,

$$\frac{d}{dt} \iiint_{\Omega} c\rho_0 T \, dx \, dy \, dz = - \iint_{\partial\Omega} \vec{J} \cdot \vec{n} \, dS.$$

$$\iiint_{\Omega} c\rho_0 \frac{\partial T}{\partial t} \, dx \, dy \, dz = - \iiint_{\Omega} k \operatorname{div} \vec{F} \, dx \, dy \, dz$$

$$\iiint_{\Omega} c\rho_0 \frac{\partial T}{\partial t} + k \operatorname{div} (-\nabla T) \, dx \, dy \, dz = 0.$$

$$c\rho_0 \frac{\partial T}{\partial t} = k \operatorname{div} (\nabla T) \Rightarrow \boxed{\frac{\partial T}{\partial t} = \frac{k}{c\rho_0} \Delta T} \Bigg|_{\text{heat equation}}$$

If  $T$  is time indep.  $\Rightarrow \Delta T = 0$

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Maxwell's equations

 $\vec{E}$  electric field $\vec{H}$  magnetic field $\vec{J}_c$  conduction current density $\rho$  charge density

$$\textcircled{1} \quad \operatorname{div} \vec{E} = \rho \quad (\text{Gauss' law})$$

$$\textcircled{2} \quad \operatorname{div} \vec{H} = 0 \quad (\text{no magnetic sources})$$

$$\textcircled{3} \quad \nabla \times \vec{E} + \frac{\partial \vec{H}}{\partial t} = \vec{0} \quad (\text{Faraday's law})$$

$$\textcircled{4} \quad \nabla \times \vec{H} - \frac{\partial \vec{E}}{\partial t} = \vec{J}_c \quad (\text{Ampere's law}).$$

$\textcircled{1}, \textcircled{2}, \textcircled{3}, \textcircled{4} \hookrightarrow$  Wave equation

$$\frac{\partial^2 \phi}{\partial t^2} = \Delta \phi$$

Partial Differential equations (ODE)

Isaac Newton: "All in nature reduces to differential equations".

Max Planck: "Present day physics, as far as it is theoretically organized, is completely governed by a system of space-time differential equations".